

# 1ES 0229 with XMM-Newton

And its impact on calibration

Felix Fürst for the EPIC IDT at ESAC

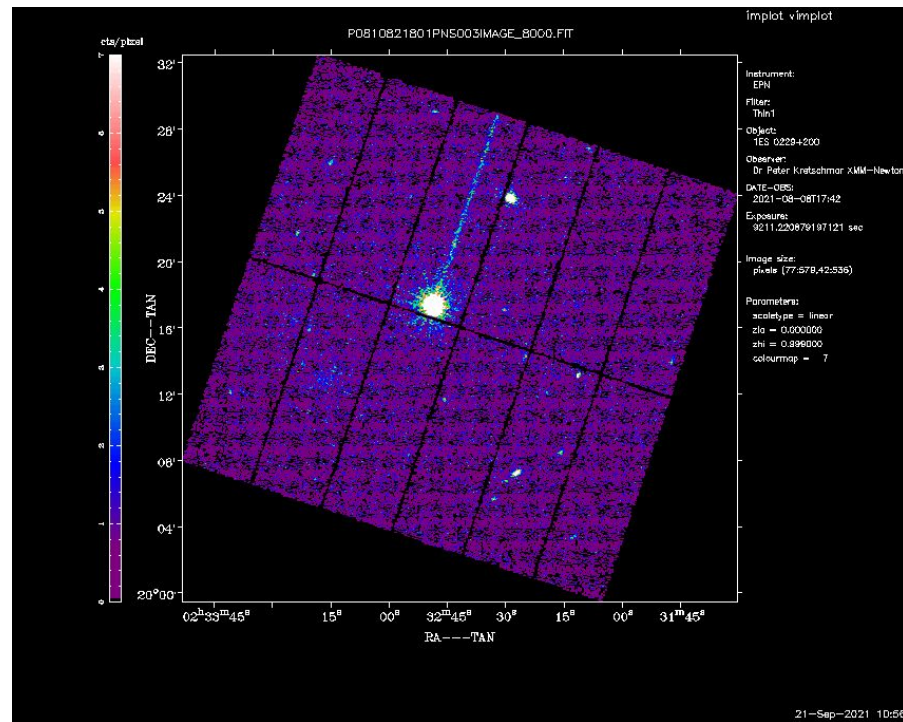
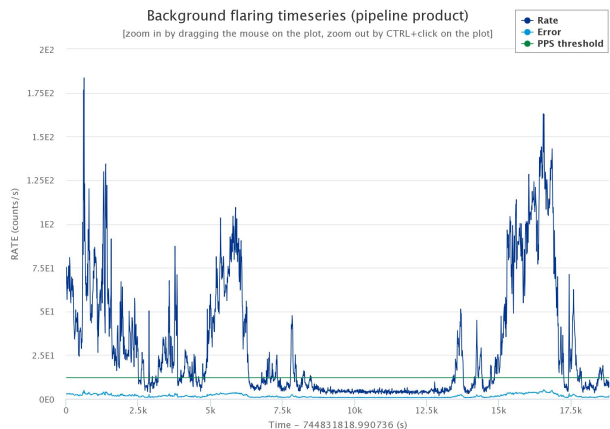
2021/10/22

IACHEC Coordinated Obs WG meeting

# ObsID 0810821801 on 2021-08-08 to 08-10

For EPIC-pn

- Full Frame exposure with 18.8ks and very bad background flaring
- **Small Window** exposure with 103ks, resulting in 60ks GTI



# pn-only fits

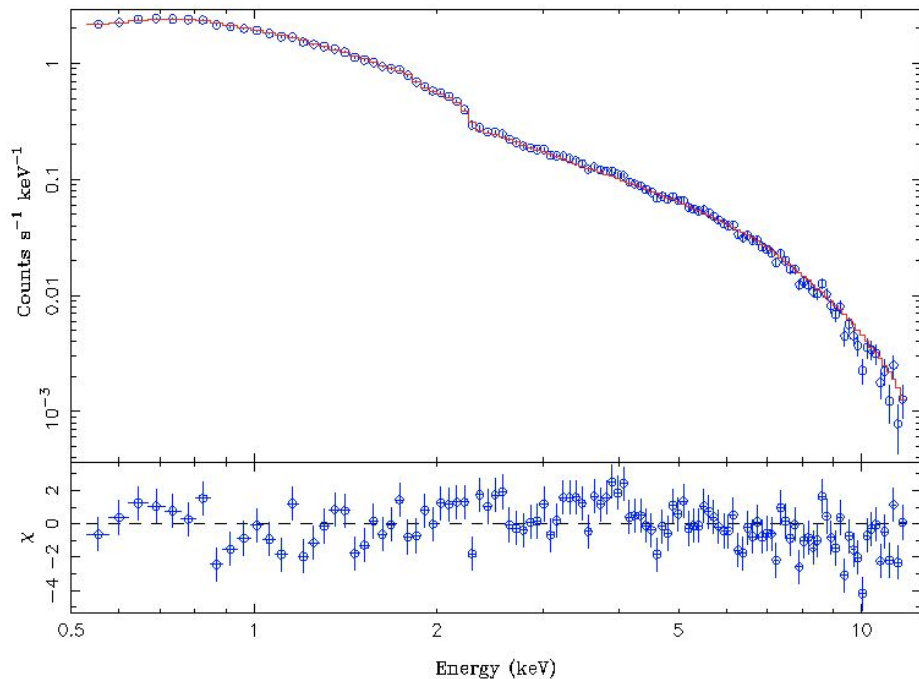
tbnew\*powerlaw

$N_H = 0.140 \pm 0.005 \text{ e}22$

$\Gamma = 2.101 \pm 0.012$

Flux (2-10keV) =  $5.36 \pm 0.06 \text{ e-}12$   
erg/s/cm<sup>2</sup>

$\chi^2_{\text{red}} = 1.64 (184/112)$



# pn-only fits

Gainshift ( tbnew\*powerlaw )

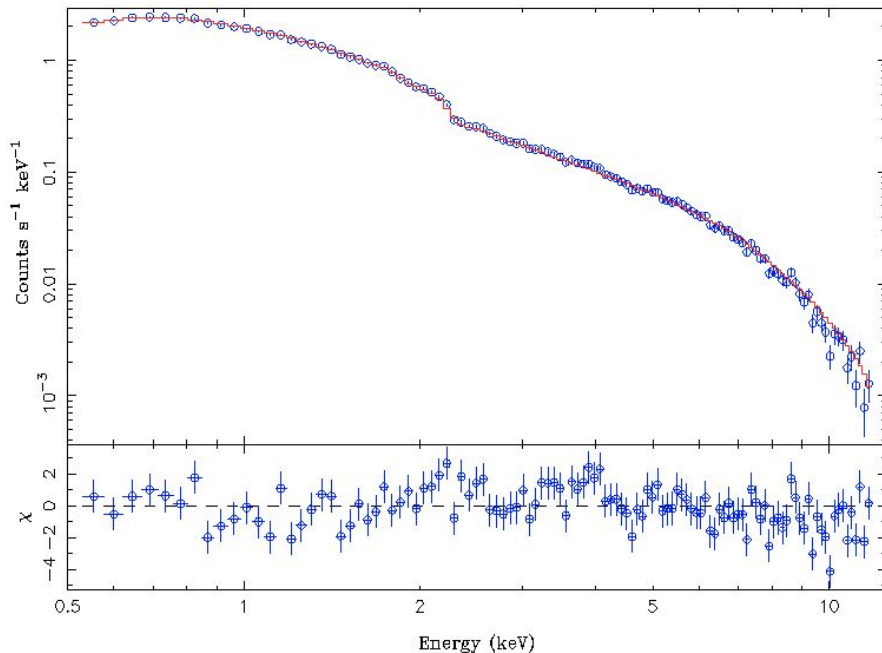
$N_H = 0.150 \pm 0.009 \text{ e}22$

$\Gamma = 2.115 \pm 0.015$

Flux (2-10keV) =  $5.43 \pm 0.07 \text{ e-}12$   
erg/s/cm<sup>2</sup>

Intercept =  $-0.017 \pm 0.010$

$\chi^2_{\text{red}} = 1.57 (174/111)$



# pn-only fits

tbnew\*cutoffpl

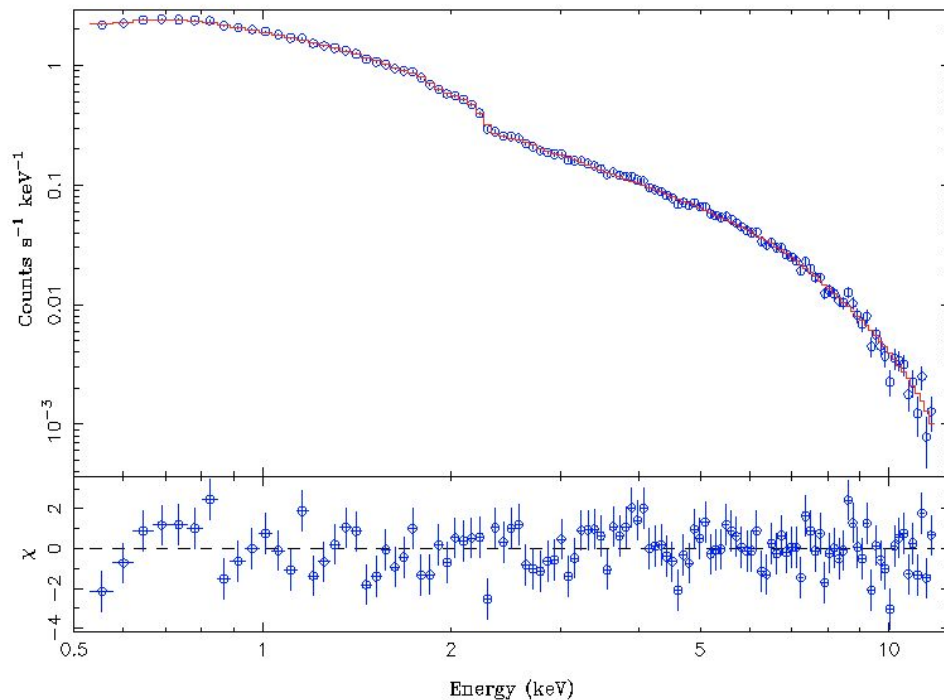
$N_H = 0.114 \pm 0.008 \text{ e}22$

$\Gamma = 1.93 \pm 0.05$

$E_{\text{cut}} = 19^{+6}_{-4} \text{ keV}$

Flux (2-10keV) =  $5.28 \pm 0.06 \text{ e-}12$   
erg/s/cm<sup>2</sup>

$\chi^2_{\text{red}} = 1.20 \text{ (133/111)}$



# Joint fits (NuSTAR + pn)

tbnew\*powerlaw

CALDB v20210908

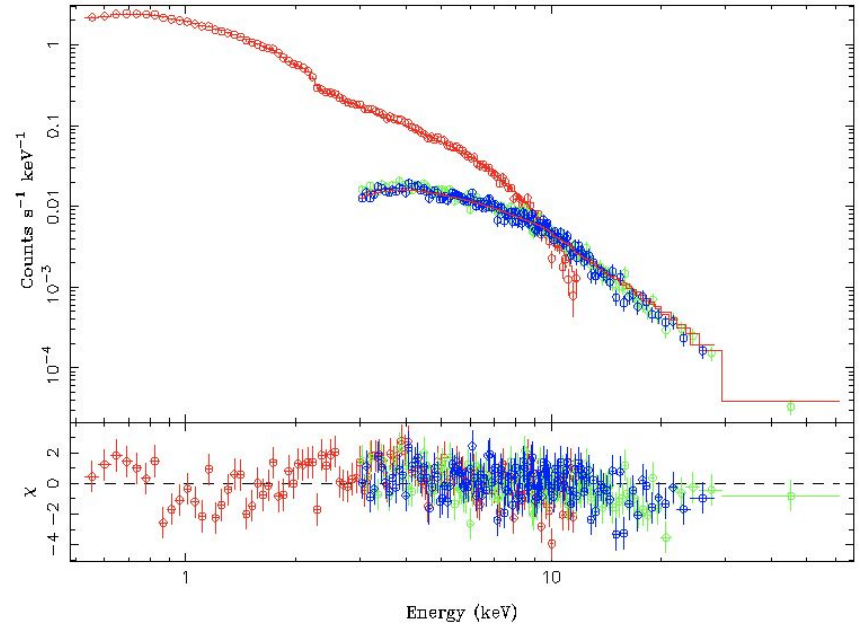
$N_H = 0.147 \pm 0.005 \text{ e}22$

$\Gamma = 2.123 \pm 0.011$

$C_{\text{FPMA}} = 1.036 \pm 0.024$

$C_{\text{FPMB}} = 1.069 \pm 0.025$

$\chi^2_{\text{red}} = 1.26 (477/379)$



# Joint fits (NuSTAR + pn)

tbnew\*powerlaw

CALDB v20210806 (NEW)

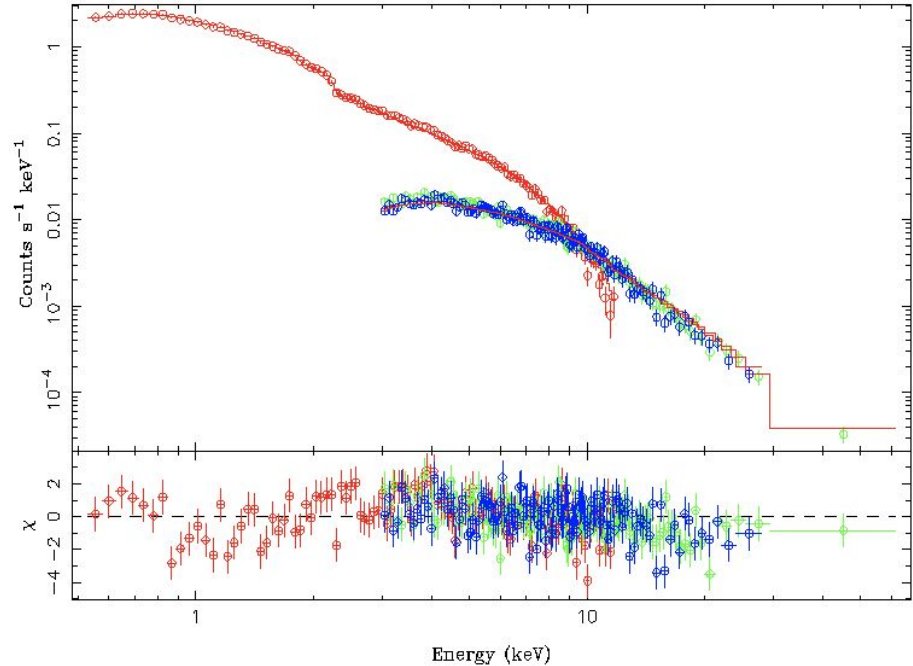
$N_H = 0.148 \pm 0.005 \text{ e}22$

$\Gamma = 2.125 \pm 0.011$

$C_{\text{FPMA}} = 1.131 \pm 0.026$

$C_{\text{FPMB}} = 1.131 \pm 0.026$

$\chi^2_{\text{red}} = 1.28 (487/379)$



# Joint fits (NuSTAR + pn)

tbnew\*cutoffpl

CALDB v20210806 (NEW)

$N_H = 0.125 \pm 0.006 \text{ e}22$

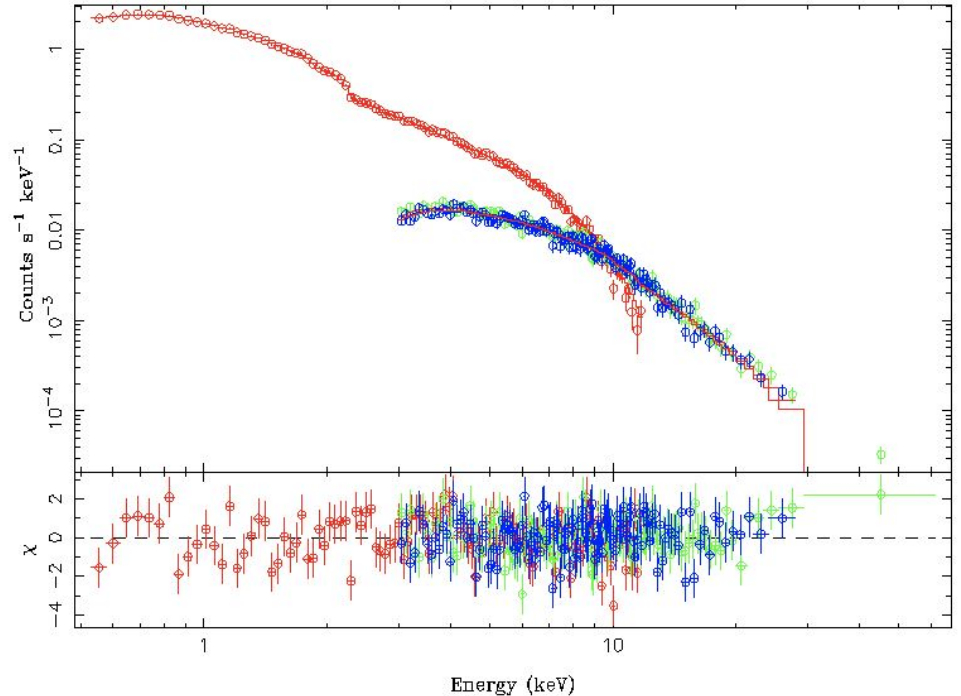
$\Gamma = 2.000 \pm 0.024$

$E_{\text{cut}} = 32^{+7}_{-5} \text{ keV}$

$C_{\text{FPMA}} = 1.157 \pm 0.026$

$C_{\text{FPMB}} = 1.152 \pm 0.026$

$\chi^2_{\text{red}} = 0.99 \text{ (374.5/378)}$



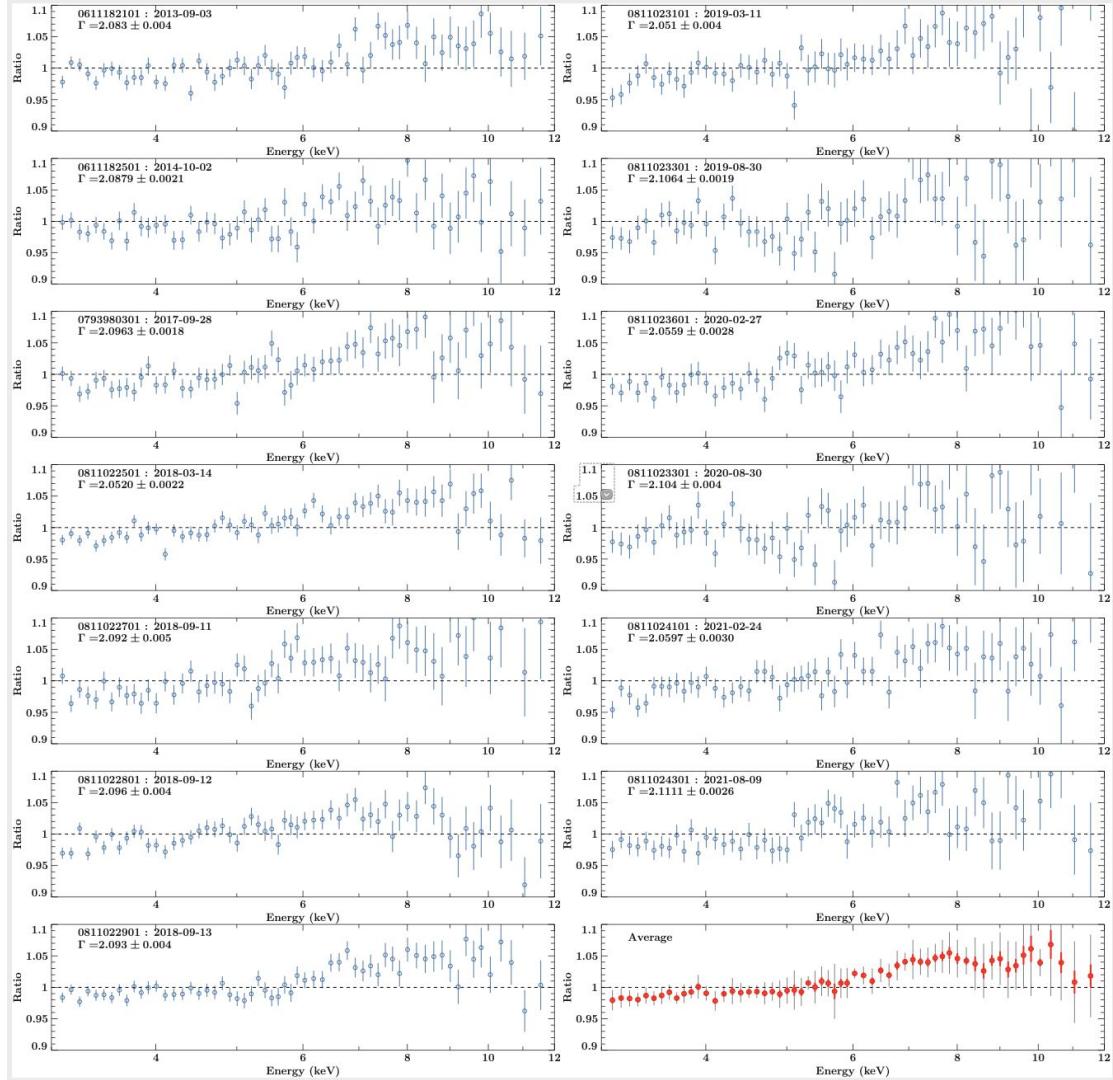


# Crab Calibration

Joint observations every 6 months.

EPIC-pn in burst mode, which causes all kinds of difficulties.

Used to create a correction function of EPIC-pn ARF, but can't tell us about absolute flux difference.

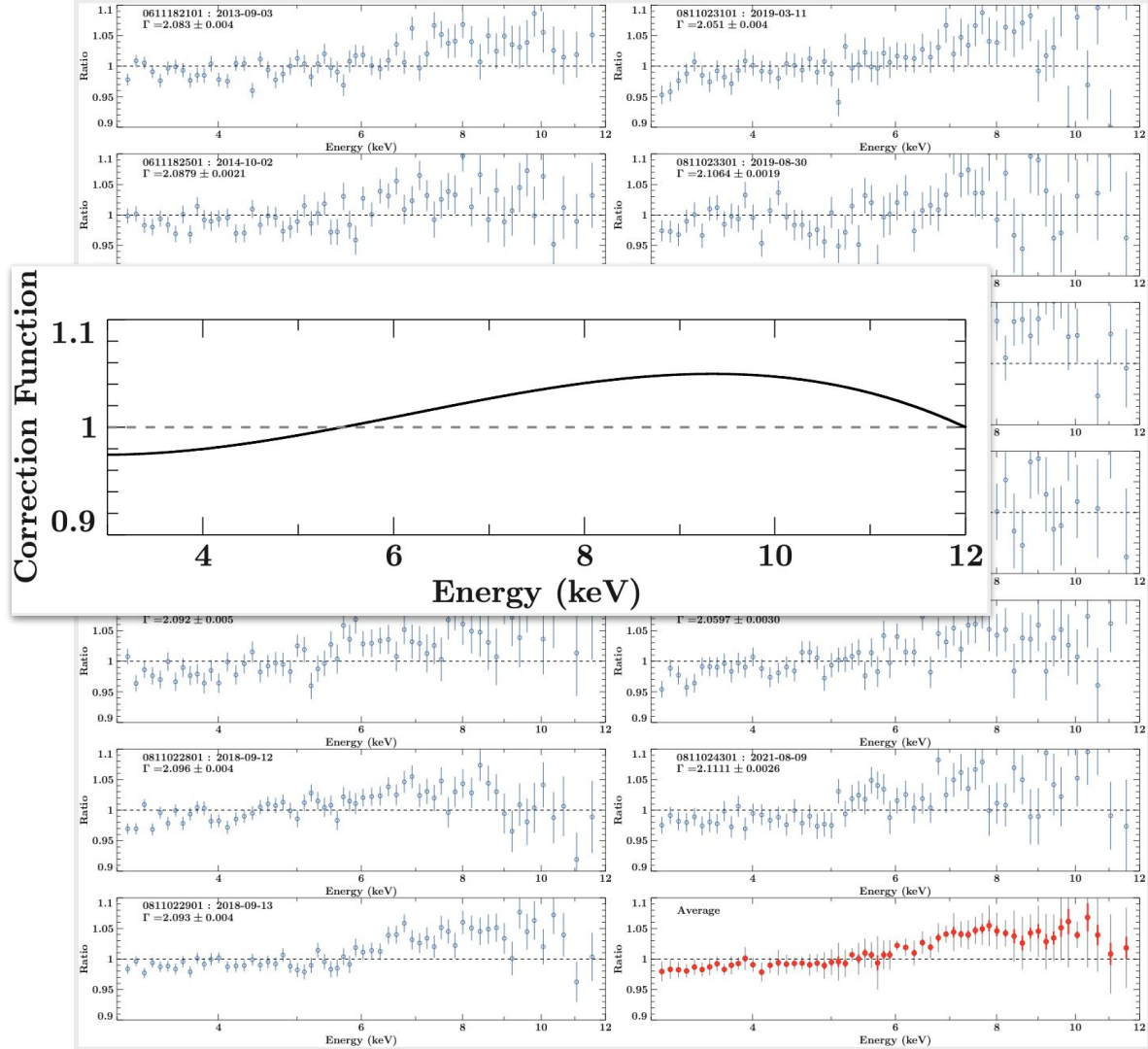


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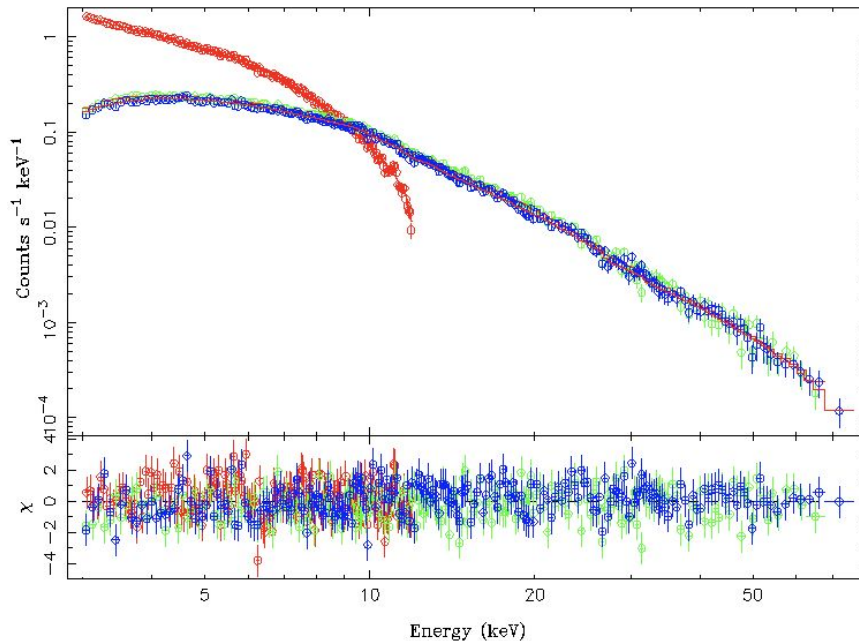


# 3C 273: ye olde calibration targhet

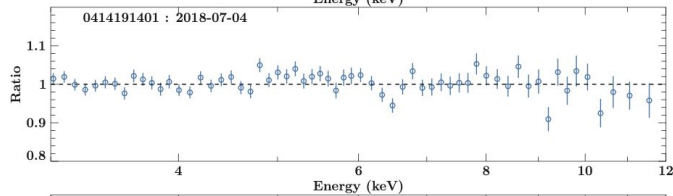
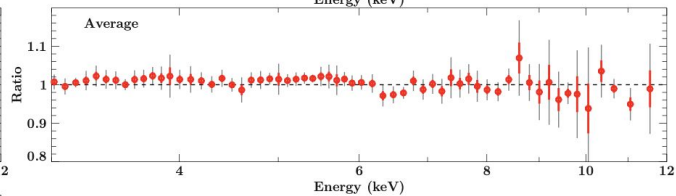
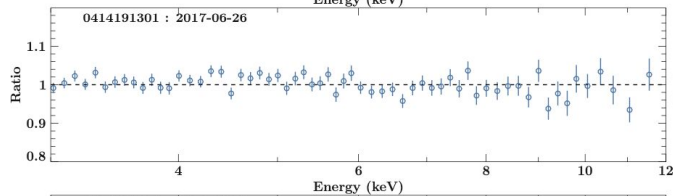
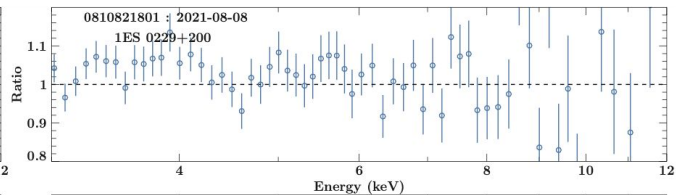
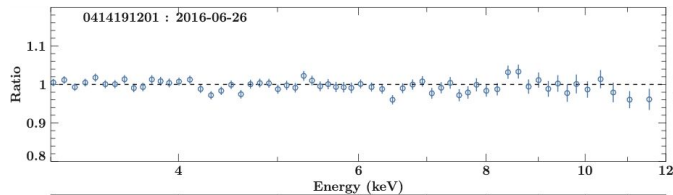
7 epochs of 3C 273 data between 2012-2020.

Fitted with `cutoffpl + xillver` (xiller normalization fixed in all epochs, distant reflection, see Madsen et al. 2015).

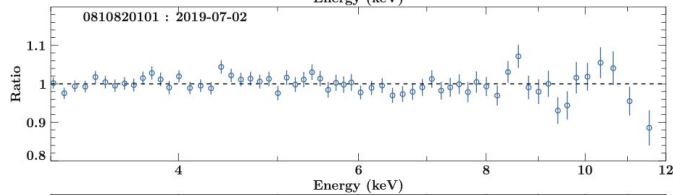
Complex model makes calibration more tricky, flux is a very close to / above pile-up limit



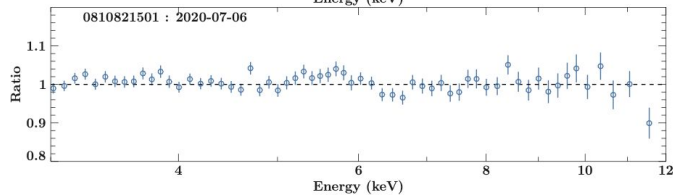
# 3C273 + 1ES 0229: the new dream team?



1ES 0229 (right) has clearly worse S/N in this observation than the average 3C 273 observation (left column).



Effects of correction likely lost within the noise.



This is for CALDB v20210806 (NEW)

$$pn = 0.91 * FPMA$$

Absolute flux correction with current available calibration

$$pn = 0.83 * FPMA$$

Absolute flux correction with new NuSTAR calibration

# Conclusion

(preliminary)

- 1ES 0229+200 has a nice, simple spectrum
  - However, its so much fainter than 3C 273 that it is hard to build up enough S/N
  - Unclear if 1ES has a different cross-calibration than in 3C 273
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# Outlook

(future work)

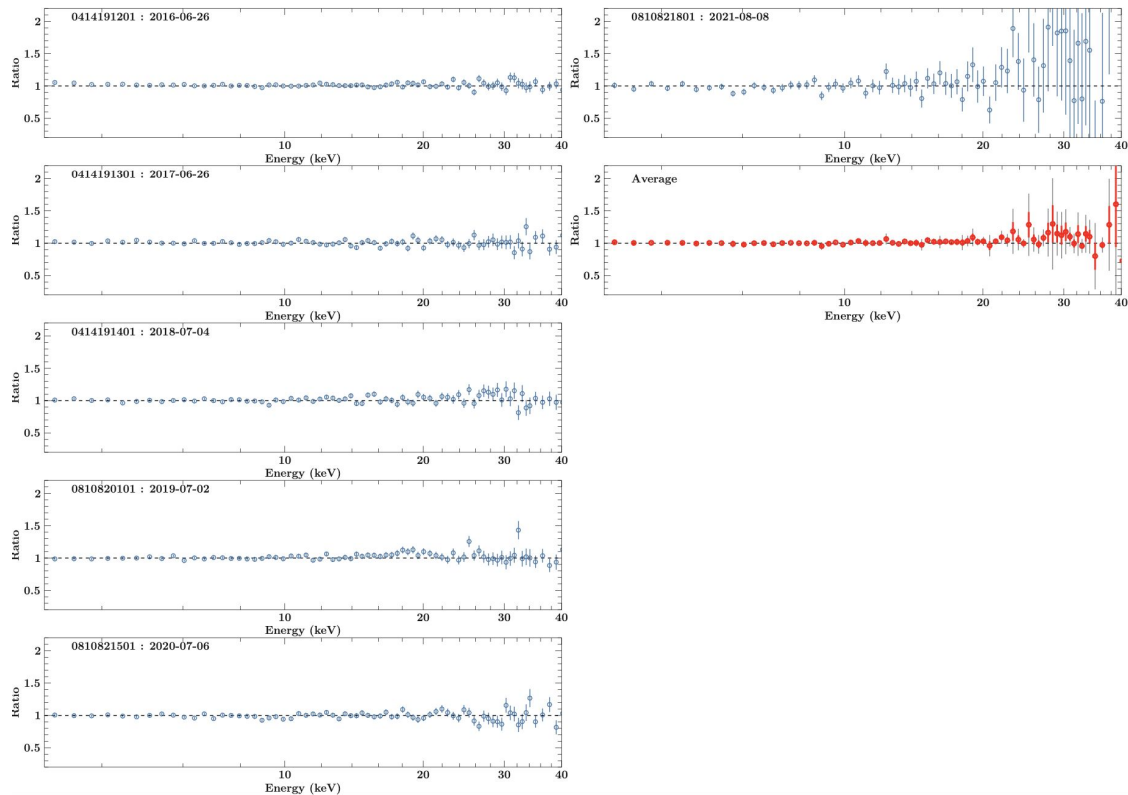
- Power Law slope of 1ES 0229 needs to be validated with MOS and OM, as well as Swift/XRT and UVOT
  - Correction function needs to be vetted with larger array of observations and uncertainties calculated with MCMC
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# Appendix



# 3C273 +1ES 0229 for NuSTAR/FPMA



# Crab is extended

EPIC-pn is operated in timing mode.

We need to make sure that *NuSTAR* data is extracted from same source region as CCD4 pn footprint on the sky.

Seasonal change between Northern and Southern are of the remnant.

