

Bias in Chisq Estimation

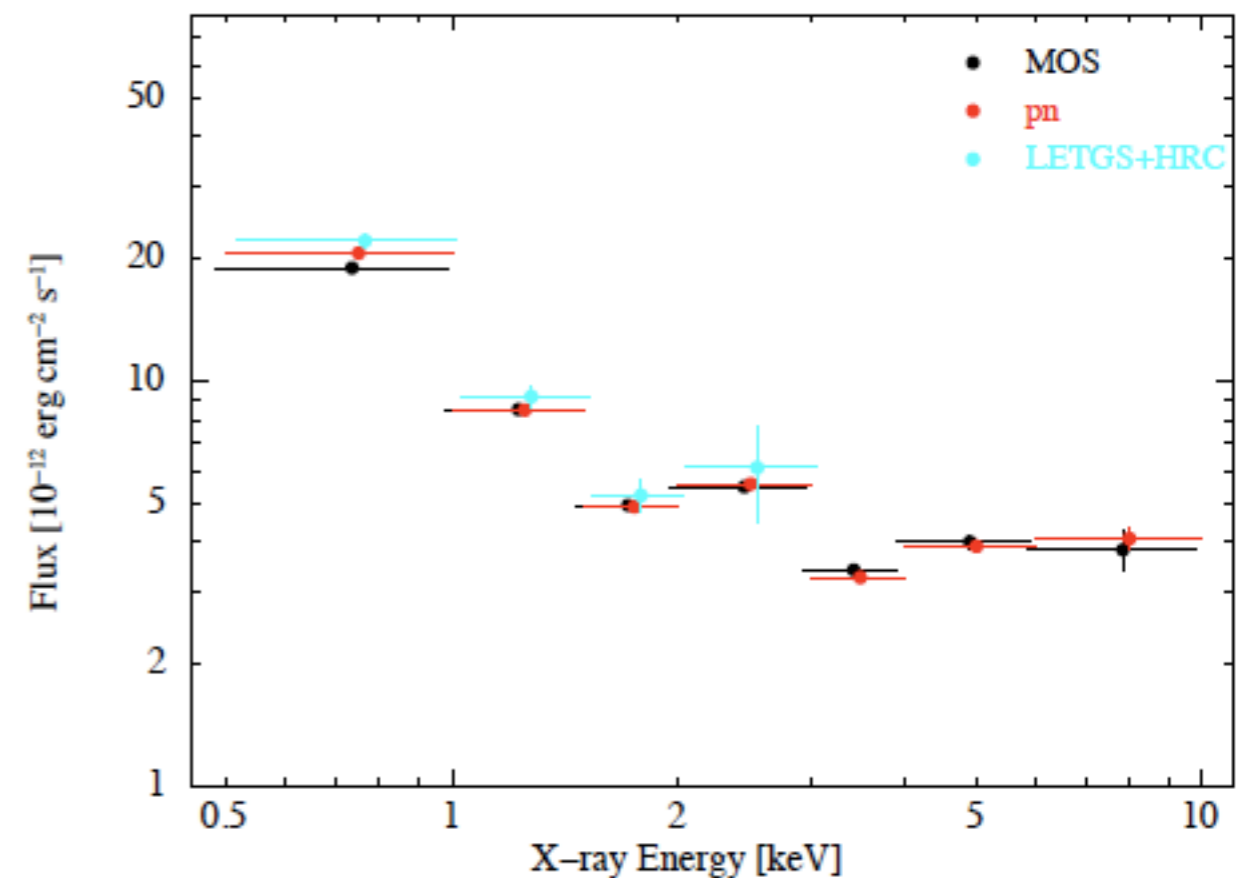
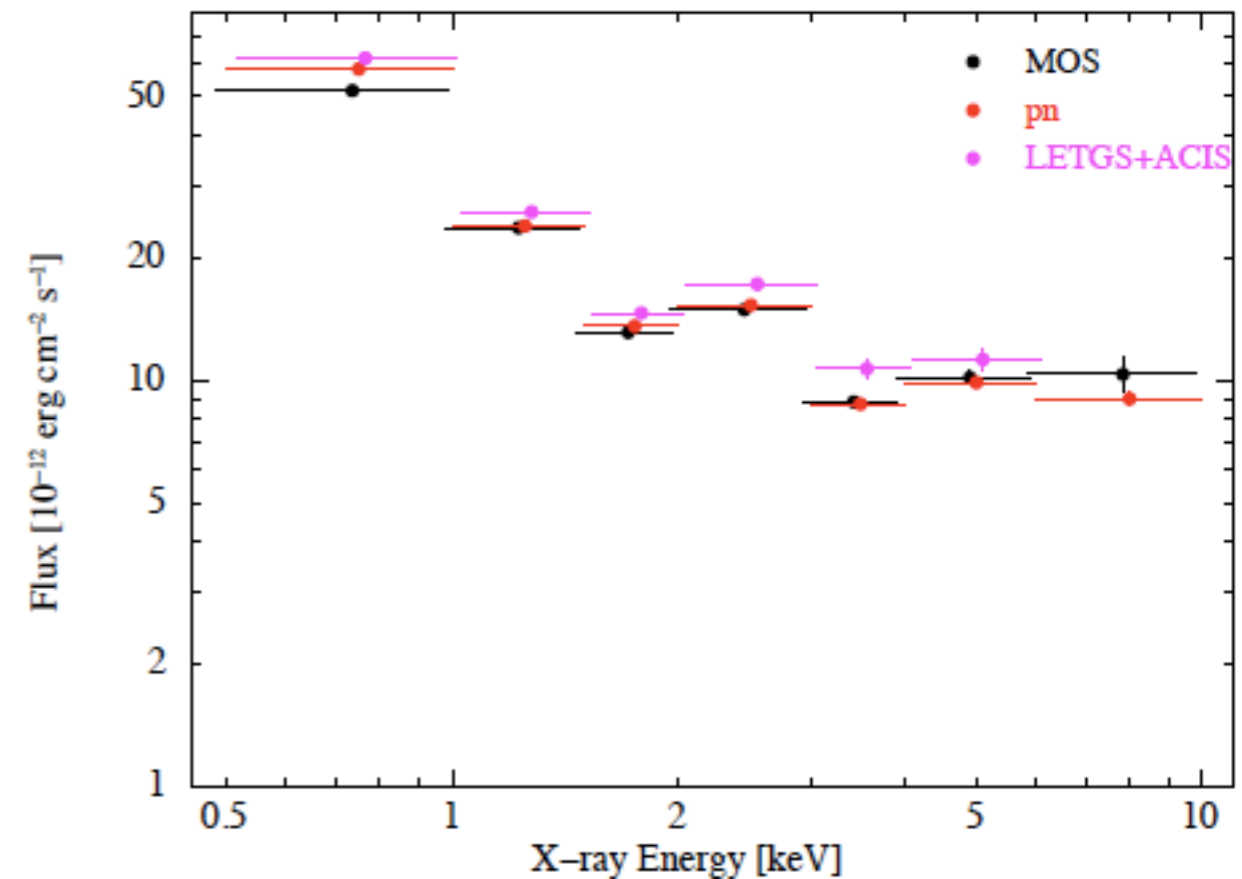
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Fitting Power Laws in Narrow Energy Ranges

- Objective: Coarse characterization of systematic errors
- Method:
 - Define narrow energy bands
 - Fit power law to spectrum in each band
 - Compute flux in each band using model
 - Compute confidence interval for each flux
 - Compare fluxes for different instruments
- Claim: flux is robust to error in model
- Shortcut for grating spectra: straight sums

Cross-calibration with PKS 2155

- Ishida et al (2011)
- Direct result of IACHEC
- Joint Suzaku, XMM, & Chandra
- Each combination examined
- Overall fits to power law
- Fluxes in bands (by PL fits)
- No conclusion yet....



Application to HETGS

- Cross-check results with direct method
- Data = $\{C_i, E_i\}$, measured in time t
- Effective area = A_i
- Estimator:

$$F(E_{\min}, E_{\max}) = \sum_{E_{\min}}^{E_{\max}} \frac{C_i E_i}{t A_i}$$

- Is this the best estimator?
- Is it biased?

Estimation Methods

- Consider simple situation (Case 1)
 - Source has invariant photon flux n
 - Observe twice with effective area A
 - Exposure times are t_1, t_2 , counts C_1, C_2
- One estimate of n (χ^2):

$$n = \frac{n_1/\sigma_1^2 + n_2/\sigma_2^2}{1/\sigma_1^2 + 1/\sigma_2^2}, n_1 = \frac{C_1}{At_1}, n_2 = \frac{C_2}{At_2}, \sigma_1 = \frac{\sqrt{C_1}}{At_1}, \sigma_2 = \frac{\sqrt{C_2}}{At_2}$$

- Maximum Likelihood (Poisson) estimate of n :

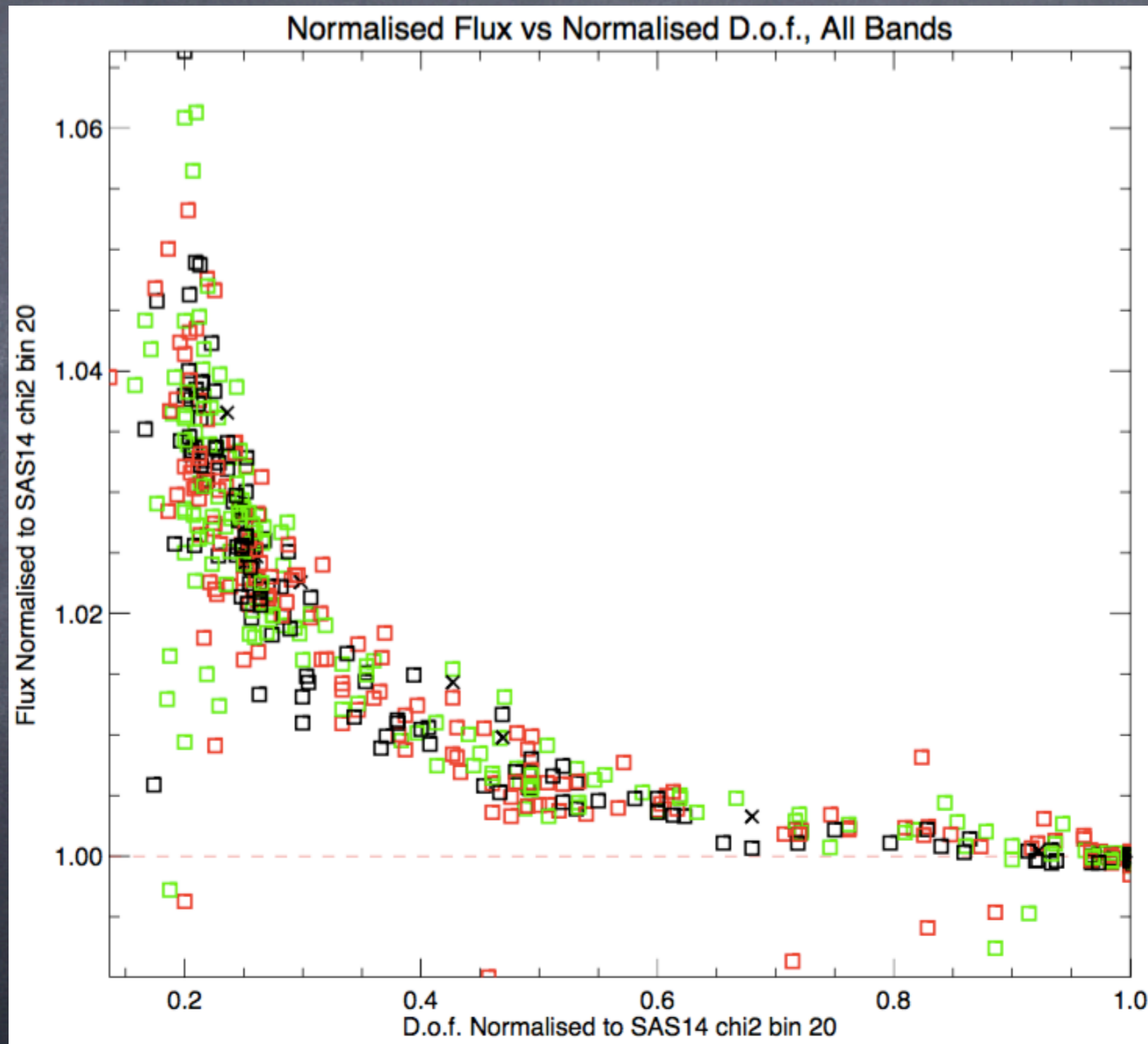
$$n = \frac{C_1 + C_2}{A(t_1 + t_2)}$$

Chisq v. ML

- Bevington (p. 248)
 - model: $y = \alpha e^{-\beta x^2} + \gamma$; data: $(y_i, \sigma_i), y \sim P(y[x_i])$
 - Fit using χ^2 stat giving $(\alpha', \beta', \gamma')$
 - Define $A = \sum y_i, A' = \sum y[x_i; \alpha', \beta', \gamma']$
 - Then using χ^2 stat gives $A' = A - \chi_{\min}^2$
 - If $\sigma_i = \sigma$, then $A' = A$ (but unexplained)
- ML treatment gives $A = A'$
 - Simple case: $y \sim P(\alpha)$, M equal bins: $\alpha' = N/M$
 - Fit using χ^2 : $1 - A'/A = \chi_{\min}^2/N \approx (M-1)/N \approx 1/(\text{SNR})^2$
 - Fit using $\chi^2, \sigma_i = \alpha$: $A' = A + \chi_{\min}^2/2$
 - Also true for $y \sim P(\eta_i \alpha)$, with known η_i

Cross-Cal Case

- Example from fitting XMM spectra in bands



Simple Cases

- Case 2: two observations, different areas and exposures:

$$n = \frac{C_1 + C_2}{A_1 t_1 + A_2 t_2}$$

- Case 3: estimate narrow band energy flux (two observations, same band)

$$F = E \frac{C_1 + C_2}{A_1 t_1 + A_2 t_2}$$

Extending Chisq v. ML

- Case 4, analogous to counts in HETGS

- Model: $y = \omega_i \mu_i F, \sum \mu_i = 1$

- $\mu_i =$ unknown fractional flux in bin i (at energy E_i) of M

- $\omega_i = TA_i/E_i =$ known flux/count scaling, total count is $N = \sum C_i$

- ML: $F' = N / \sum \omega_i \mu_i, \mu'_i = C_i / (F' \omega_i)$

- using $\sum \mu_i = 1$, then $F' = \sum C_i / \omega_i = \sum C_i E_i / (TA_i)$

- flux is sum of flux estimates in each bin

- Uncertainty: $\sigma_F = F / \sqrt{N}$

- χ^2 : Same answer!

- M unknowns (F, μ_i) , $N_{\text{DoF}} = 0 \rightarrow \chi^2_{\text{min}} = 0$

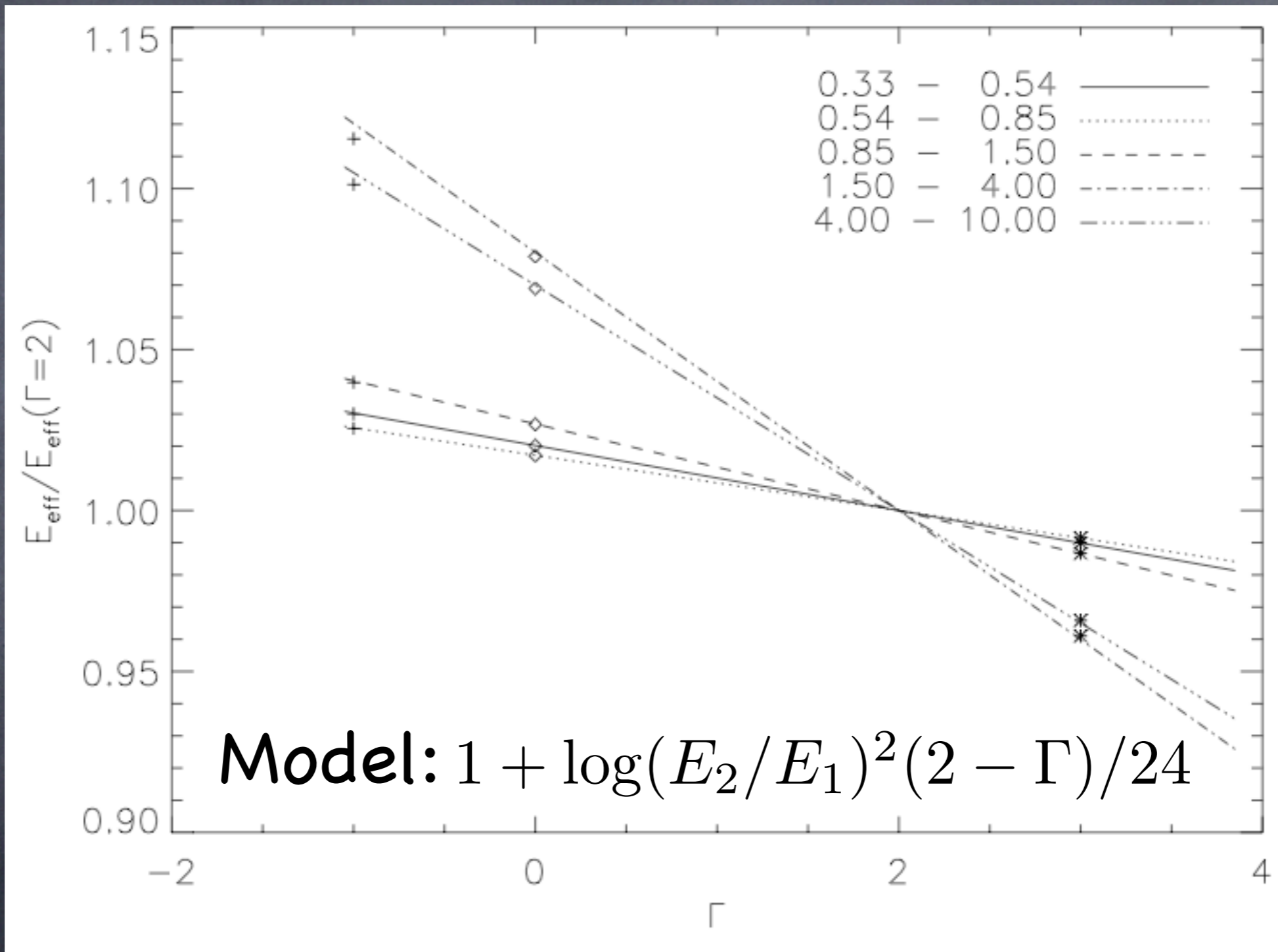
The Case of Interest

- PL spectral model, want broad-band flux
 - known Γ , $n(E) = K(E/\hat{E})^{-\Gamma}$
 - data: counts in equal bandpasses of size ΔE

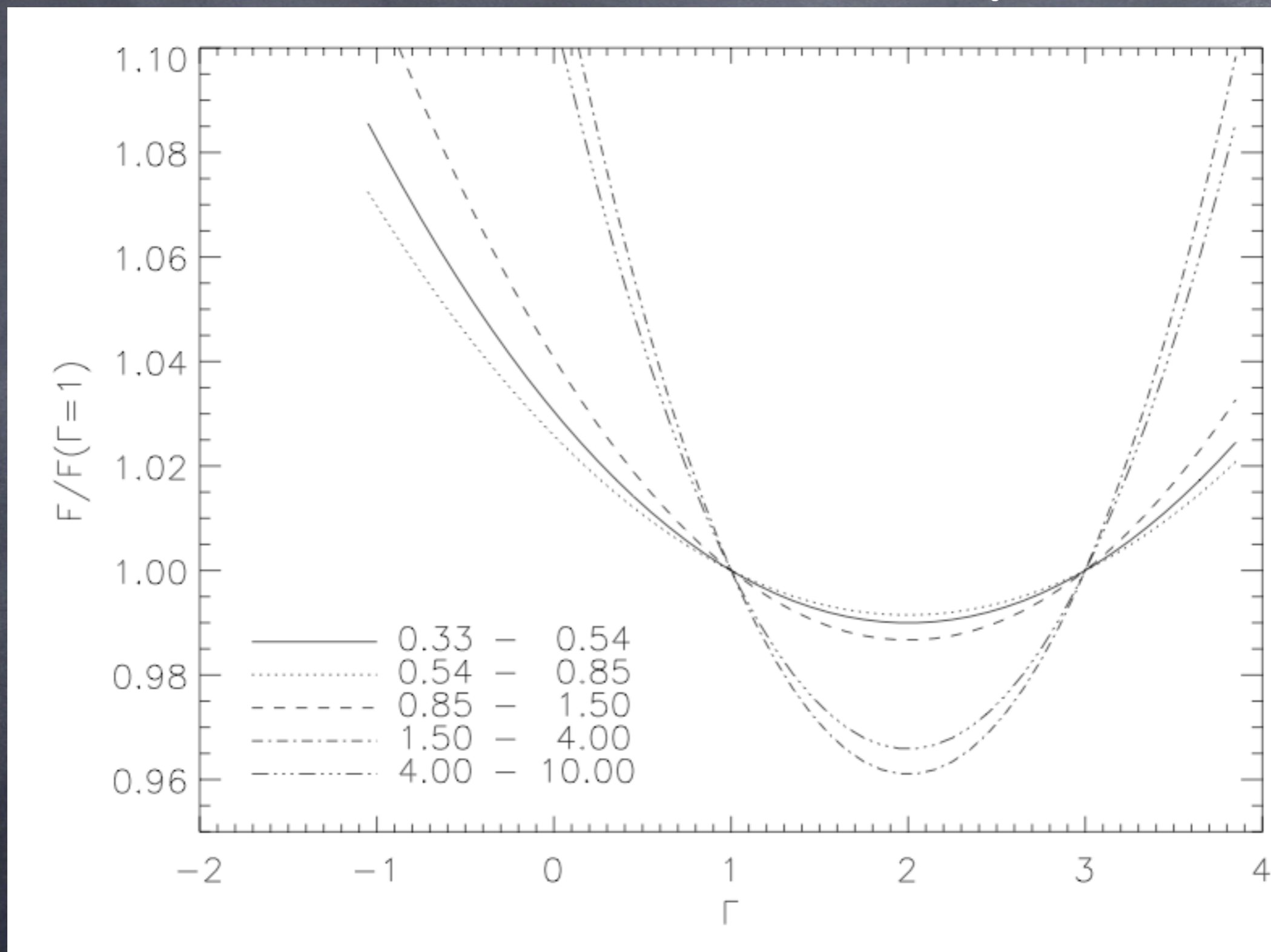
$$K = \frac{\sum C_i}{\Delta E \sum A_i t_i (E_i/\hat{E})^{-\Gamma}}, F = \frac{K \hat{E}^2}{2-\Gamma} \left[(E_{\max}/\hat{E})^{2-\Gamma} - (E_{\min}/\hat{E})^{2-\Gamma} \right]$$

- χ^2 : fractional error in $F = \chi^2_{\min}/N \approx (M-1)/N$
- Set reference energy to $\hat{E} = \log E_{\max}/E_{\min}$
 - How does \hat{E} depend on assumed Γ ($=2$)?
- What's the error in F if assumed Γ is wrong?

Central Energy



Flux Sensitivity



Summary

- Chisq fits: systematically low flux estimates
 - Fractional flux bias is $\sim 1/(\text{cnt/bin})$
 - Applies to fluxes in lines as well
 - emission lines: underestimated
 - optical depths: overestimated
 - Results from approx. model of stat. variations
- Maximum likelihood fluxes are unbiased
- Flux summing method is same for ML and χ^2
 - Not “best” estimator if spectral shape is known
 - Biased if full band is not represented
 - e.g. PL model of 4–10 keV is larger than sum of 4–8 keV
 - “Best” if spectrum is not easily characterized