Introduction on HXMT Project, Ground Segment, Calibration and Ground Facility

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Outline

Overview on HXMT > **Collaborations with ESA** \mathbf{i} **HXMT** ground segment > **HXMT** Calibration > data flow; organization implementation ; milestones; status **Ground facility**

Overview on HXMT



Chinese Academy of Sciences, Tsinghua University Chinese Academy of Space Technology





A sunshading board will be set so that the LE and ME telescopes can work at low temperatures

Characteristics of the HXMT Mission

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Detectors	LE: SCD, 384 cm ² ;ME : Si-PIN, 950 cm ² HE : Nal/CsI, 5000 cm ²
Energy Range	LE: 1-15 keV;ME: 5-30 keV;HE: 20-250 keV
Time Resolution	HE: 25 μ s; ME: 180 μ s;LE: 1ms
Energy Resolution	LE: 2.5% @ 6 keV ME: 8% @ 17.8 keV HE: 19% @ 60 keV
Field of View of one module	LE: $6^{\circ} \times 1.6^{\circ}$; $6^{\circ} \times 4^{\circ}$; $60^{\circ} \times 3^{\circ}$; blind; ME: $4^{\circ} \times 1^{\circ}$; $4^{\circ} \times 4^{\circ}$; blind; HE: $5.7^{\circ} \times 1.1^{\circ}$; $5.7^{\circ} \times 5.7^{\circ}$; blind
Source Location	<1' (20 o source)

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100 25 200	Orbit	Altitude: ~550 km ; Inclination: ~43°
「「人」したますい	Attitude	Three-axis stabilized Control precision: $\pm 0.1^{\circ}$ Measurement accuracy: $\pm 0.01^{\circ}$
Carlor Carlor	Data Rate	LE: 3 Mbps; ME: 3 Mbps; HE: 300 kbps
and and and	Payload Mass	~1000 kg
68 18 million	Nominal Lifetime	4 years
御天ろういう	Working Mode	Scan survey, pointed observation

Observation modes

- All-sky Survey mode
- Deep imaging observation of selected sky regions (such as the Galactic center region)
- Pointed observations

HXMT in all-sky survey



During this phase, the satellite attitude is in the three-axis stabilized state with the telescopes pointed in the plane perpendicular to the Sun. The common optical axis of the telescopes scans a great circle of constant ecliptic longitude in one orbital period. (Similar to ROSAT all-sky survey)

Main advantages and key sciences of HXMT

Hard X-ray sky survey with highest sensitivity

- High precision hard X-ray full sky map: diffuse background and cosmic variance
- Discover highly obscured supermassive BHs: galaxy formation and evolution
- Discover new types of high energy objects: usual surprises of new surveys

High precision pointed observations of high energy objects

- > Space-time in strong gravitational field: dynamics and radiation near BH horizons of stellar mass and supermassive BHs
- > Equation of state in strong magnetic field: neutron star and its surface properties
- > High energy particle acceleration: AGN, SNR, shock and relativistic jets
- > Large scale structure: through hard X-ray detection of galaxy clusters

Status on the overall project

- Officially approved by CNSA in March 2011
- Entered Phase-B in the end of 2011

Milestones

Phase-B (qualification module): finishing at the end of 2012Phase-C (flight module):from 2013 until mid 2014Launch:late 2014 to early 2015

Schedule of HXMT Operation plan:

- Year I: All-sky survey (No AO time)
- Year II: Pointed Observations and small region imaging.
- Year III: Pointed Observations and imaging
- Year IV: Pointed Observations
- ≥Year V: Pointed Observations

Status on the overall project

HXMT mission will be dominated :

By building the hardware (in Phases B and C) :

By ground segment after launch :



The automatic gain control of HE works very well



2.8keV @60keV, room temperature

1.7keV @60 keV, -28℃

Energy resolution of ME



Radiation damage experiment of the LE detector

Collaborations with ESA

ESA Assistances and helps upon:

HXMT ground segment; HXMT calibration;

(documentation, software, experienced consultant etc.)

HXMT ground segment

Configuration for HXMT Ground Segment

The HXMT Ground Segment is composed of three parts:

- (1) Track , Telemetry and Control Ground Segment (TT&CGS)
- (2) Mission Operation Ground Segment (MOGS)
- (3) Science Ground Segment (SGS)

In China, the functions of Mission Operation Center (MOC) is divided into two parts, TT&CS and MOGS.



The responsibilities of TT&CGS

The responsibilities of TT&CGS are to exchange between platform and ground for those:

(1) telemetry data(2) uploading commands(3) payload data

The responsibilities of HXMT SGS

- Observational plan and, accordingly, platform attitude and payload condition be figured out
 Contemporary payload health and performance
 Inflight calibration and calibration database
 Scientific and auxiliary data productions
 Software maintenance/assistance to both onboard and user
- (6) ToO and multi-wavelength campaign

The Progress of SGS

1. Generator to produce data analogue to inflight one: under test and mostly ready for HE.

2. Optimize strategy upon source observation: have the source and background be derived contemporarily

3、Tool to estimate the sky coverage for survey under breaks of SAA and earth occultation.

4、Lessons for the inflight background modeling from the pioneer missions



Optimized source (located at center) observations (pointing toward the write circles)



Sky coverage of survey under breaks of SAA and earth occultation

Inflight background: lessons from Suzaku/XIS



HXMT Calibration

Calibration production (Concept)

PCF (primary calibration files) : raw ground and inorbit calibration data set, most of which are of not direct interest to users in their scientific data analysis;

BCF (basic calibration files): the lowest level calibration datasets potentially required by analysis software, can be considered as the atom units of the instrument calibration;

>CPF (calibration product files): the convolution of the information stored within BCFs customized for a specific analysis task or observations.



Organization Instrument teams (The Core prior to launch): LE, ME, HE Science Ground Segment (The Core post launch) : SOC, SU&DC Project committee (The head of HXMT): every aspect of the overall project Project management (The officers): coordinating the implementation; Small technique groups (The working bees): simulation, thermal control, structure etc.



Calibration milestones

Phase B (2012-2013) (instrumental teams mostly involved):

Ground calibration plan (for PCF, by ins. team); Preliminary inflight calibration plan (for PCF, by ins. team); Preliminary calibration plan for producing BCF (by SU&DC); Figure out interface between PCF and BCF (by SU&DC);

Tests for performance of instruments (by ins. team);

Calibration milestones Phase C (2013-2014) (ground segment mostly involved):

Inflight calibration plan (for PCF, by ins. team); Calibration plan for producing BCF (by SU&DC); Calibration plan for producing CPF (by SU&DC); Inferface between ground PCF and inflight PCF (by SU&DC);

Carry out the ground calibration and produce PCF (by ins. team);

Calibration milestones

After launch (2015-) (ground segment dominated):

Implementation of inflight calibration and produce PCF; Produce BCF and CPF datasets;



Scheme: Implementation

Instrument and Component Calibration (Ongoing)

Collimator: psf, parrellel, etc.

PMT: efficiency, degrading etc.

CCD: QE, CTI, energy response, etc.

Crystal: Nal/Csl, energy response, etc.

Dependences on temperature, magnetic field, etc.

Sub-System and System Calibrations (not yet done)

The performance of the integrated sub-system: (LE, ME and HE) and system (the sum of all) Calibration Error Budgets (Not yet touched)

Scientific drivers for calibration -> the goals of the calibration accuracies of parameters.

To exam the component calibrations, make sure these goals are met and maintained by establishing error budget to work to.

In-Orbit Calibrations (plan not yet available) **Standard candles: Crab, and others see IACHEC Embedded calibration sources:** Am-241, Fe-55 **Environmental fluorescence lines :** Si, etc. **Operational conditions:** Gain, temperature etc. **Collimator PSF; energy response; effective area;**

timing etc.

Simulations (Working bees busy for years)

Cover almost whole procedure of the calibration



Contributions of various components to the total background (HXMT/HE)



Simulated in-space background of HXMT/HE

Ground facilities

Ground facilities

Radioactive sources:

Advantages: known energy and flux: energy linearity and resolution

Disadvantages: mostly constrained by weak flux and short life time; and less available at hard X-rays

Radioactive Sources (HE)

Half-life	Line Energies/keV
433y	26.3 59.54
271d	122 136.5
453d	88 24.9 22
30y	32 36.6
46.6d	279.2
59d	27.3 35.5
	Half-life 433y 271d 453d 30y 46.6d 59d

Ground facilities

Synchrotron radiation facility:

Perfect for calibration:

good energy resolution, high flux level, good parallelism, energy continuity

Constraints:

available time very limited,

small beam dispersion -> long calibrations

Facility in China	Energy band	Light spot size	Availability
Shanghai synchrotron radiation facility	< 73 keV	~ 0.3 (H) mm X 0.3 (V) mm	very low
Beijing synchrotron radiation facility	< 25 keV	~ 25 (H) mm X15 (V) mm	low
Hefei synchrotron radiation facility	< 12 keV	~ 15 (H) mm X 1 (V) mm	ok

Ground facilities

X-ray apparatus under construction at IHEP:

0.5-10 keV (multi-anode, fluorescence lines + filters for monochromatic light)

Beam line 3 m (relatively poor parallel of the produced X-rays)

Standard detector (monitor beam intensity)

Vacuum detector chamber (10⁻³pa)



X-ray apparatus

(Manson Model 05 from McPherson company)

Vacuum detector chamber

(diagram from Shanghai Shuguang company)

Ground facilities

Designed for upgrade in future:



X-ray apparatus: voltage >60 kV, power >10 kw, Xrays up to 60 keV: cover hard X-rays

Double crystal monochromator : more energy points Beam line extended to > 30m: better parallelity Utility of the ground facilities: the perspectives

Radioactive sources: supplementary to all, HE in special

Synchrotron radiation facility: apply time to cover LE, ME, and partially HE

The X-ray apparatus: takes care of the LE and partially ME

Relatively safe for LE but still largely open for HE and ME

THANK YOU !