



A Rapidly Prepared Analysis of Ionosphere Dynamics

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1 Scientific Goals

Solar Wind and IMF

- Data and Analysis
- DMSP and SuperDARN
- NOAA POES
- Magnetometers
- Sondestrum ISR
- EISCAT

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Can we observe signatures of energetic electrons from the Duskside Relativistic Electron Losses (DREP) as ionization in the D region. Study the auroral oval at altitudes 90-700 km





Sondrestrom and EISCAT incoherent radars, for the time period t=22:00 UT-23:30 UT:

- Perform three, 24 minute long, elevation sweeps along magnetic meridian. Use pulse option 1 for Sondrestrom and CP3 for EISCAT
- Collect data for: electron and ion temperature, ion radial velocity, and electron density.



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Supporting measurements:

- Satellite based: ACE, NOAA POES, DMSP
- Radar: SuperDarn
- Ground based magnetometer network: west Greenland and Norway.



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Solar Wind and IMF



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Solar Wind and IMF



Figure: Traveling time between the L1 point and the Earth: approx. 40 min.

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Ion velocities and ion energy



Convection map movie

Ion velocities and ion energy



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Ion velocities and ion energy



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NOAA POES Total Energy Detector





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Magnetometers



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Sondestrum Ne 22:06:34 - 22:30:34



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Sondestrum N_e 22:31:35 - 22:55:35



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Sondestrum Ne 22:56:36 - 23:20:36



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Sondestrum T_e 22:06:34 - 22:30:34



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Sondestrum V_{los} 22:06:34 - 22:30:34



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Sondestrum V_{los} 22:31:35 - 22:55:35



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Sondestrum V_{los} 22:56:36 - 23:20:36



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EISCAT Scan 1 (and 2)





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EISCAT Scan 3





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EISCAT Scan 3 Zoomed



EISCAT Electron Density



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- Overall southward movement of plasma in agreement with convection pattern
- Energetic particle precipitation
- Clear F region
- Clear E region over EISCAT site.
- Northward movement of the auroral event over Greenland
- No high energy electrons detected on the duskside; high electron densities on the dawn side at lower altitudes.

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