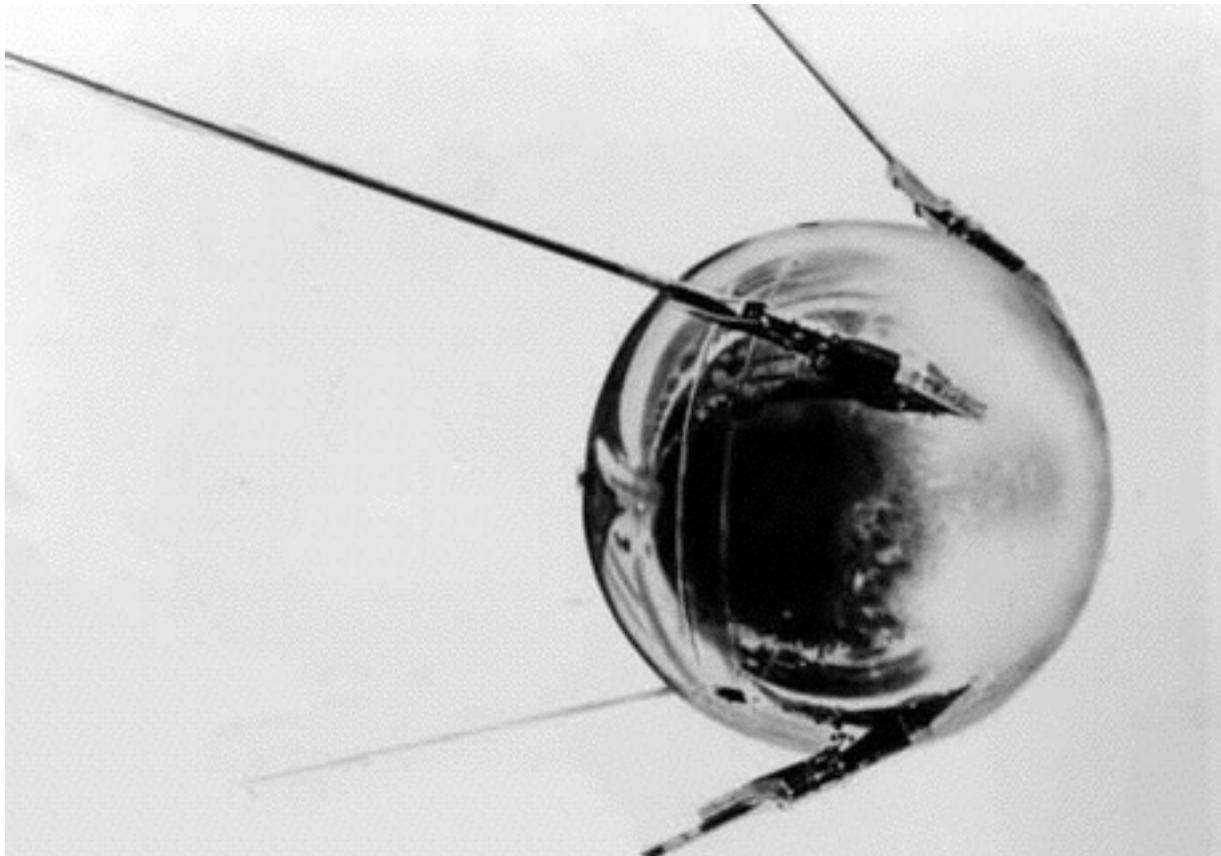


# Early History of the Jicamarca Radio Observatory

Ronald Woodman  
Instituto Geofísico del Perú

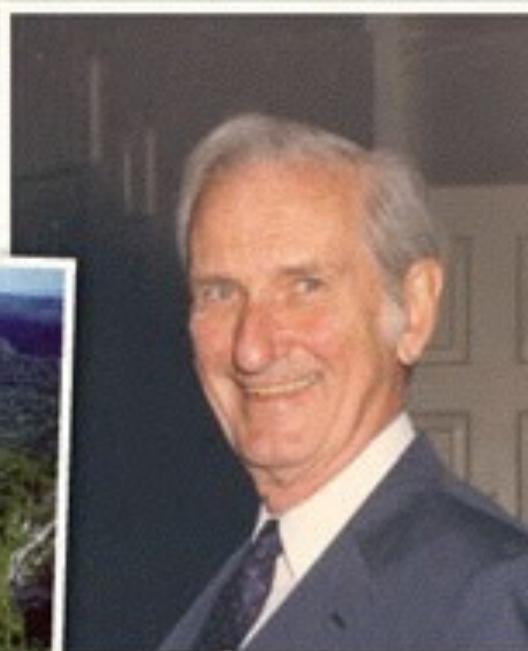
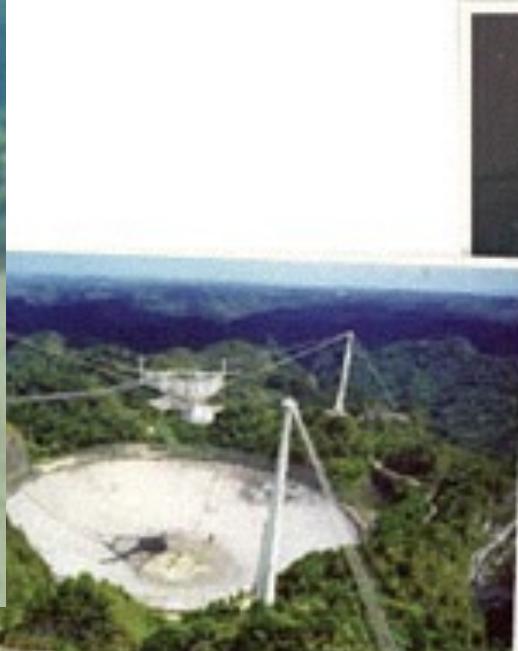
# Spunick, 1958

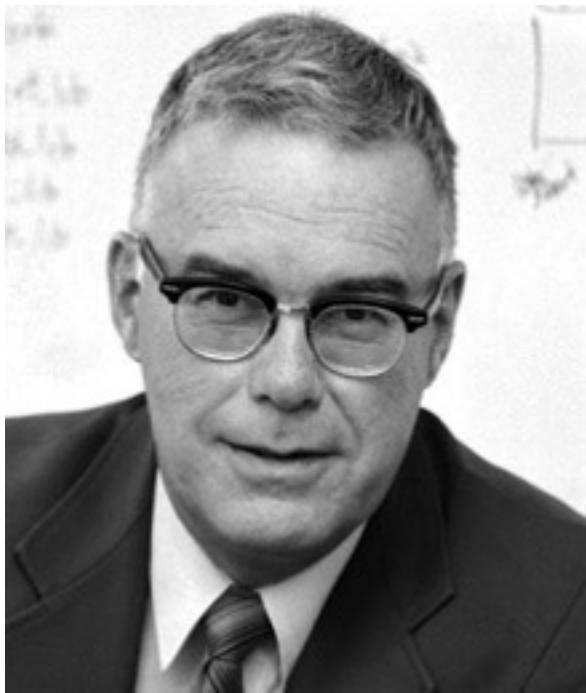


# William E. Gordon



1958 paper, proposes the technique

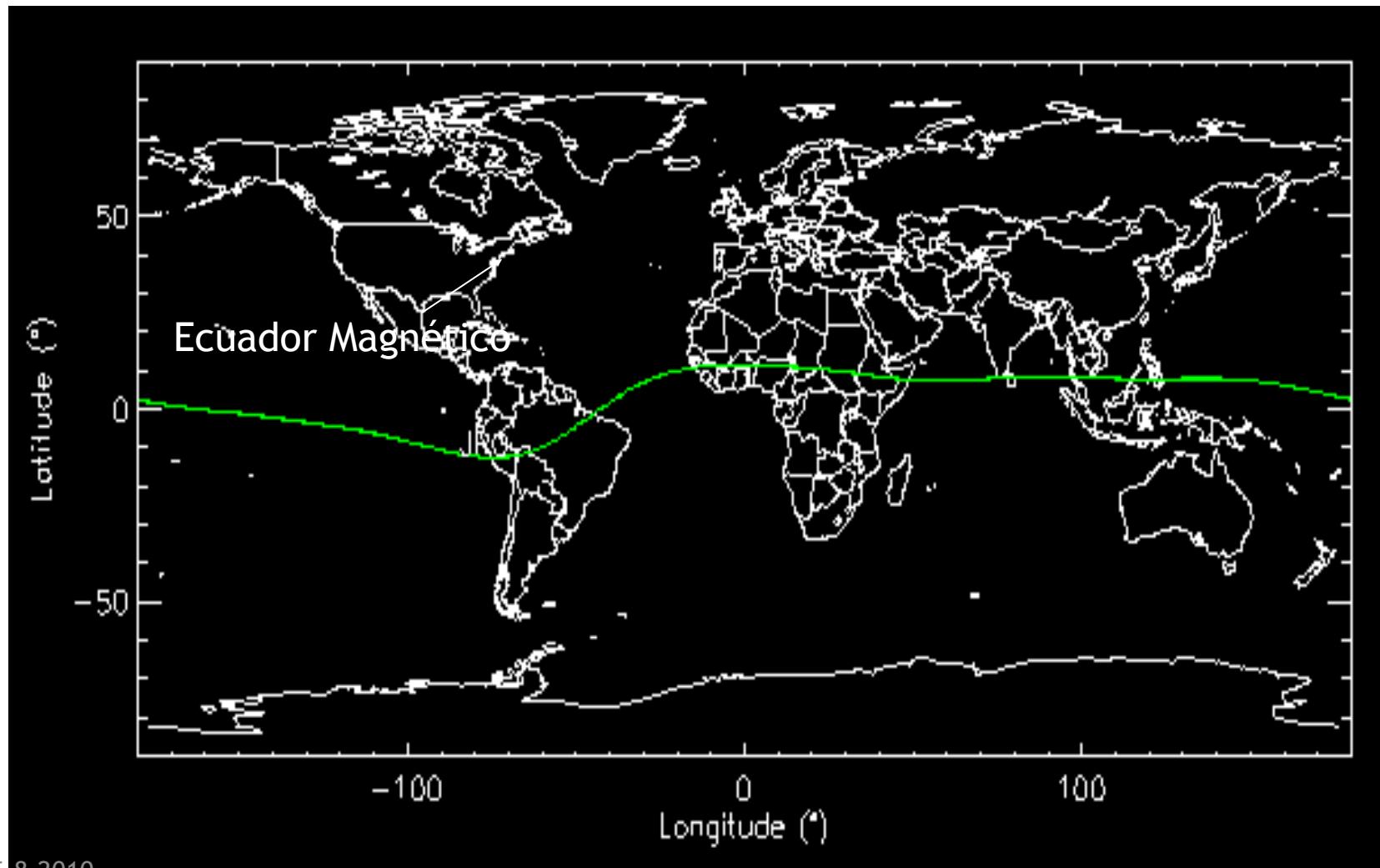




## Ken Bowles

- Performed successfully the first Incoherent scatter experiment in Urbana, Illinois, 1958
- Proposed and build the the JRO, 1960-1962

# Locus of Magnetic Equator



Carabayllo District

Ventanilla  
Callao

JRO

San Juan de Lurigancho

de Porres

Callao

Lima

Santa Anita

Ate

La Molina

Santiago de Surco



Image © 2015 DigitalGlobe

© 2015 Google

Image © 2015 CNES / Astrium

Data SIO, NOAA, US Navy, NGA, GEBCO

Lima and JRO Location



Bob Cohen and Ken Bowles checking the site for interference



Aerial look of the site chosen for construction

# Huayco and cable chair







# H. Ochs, Hector Cabada, Gerardo Vera. Antonio Arevalo and the golden dipole





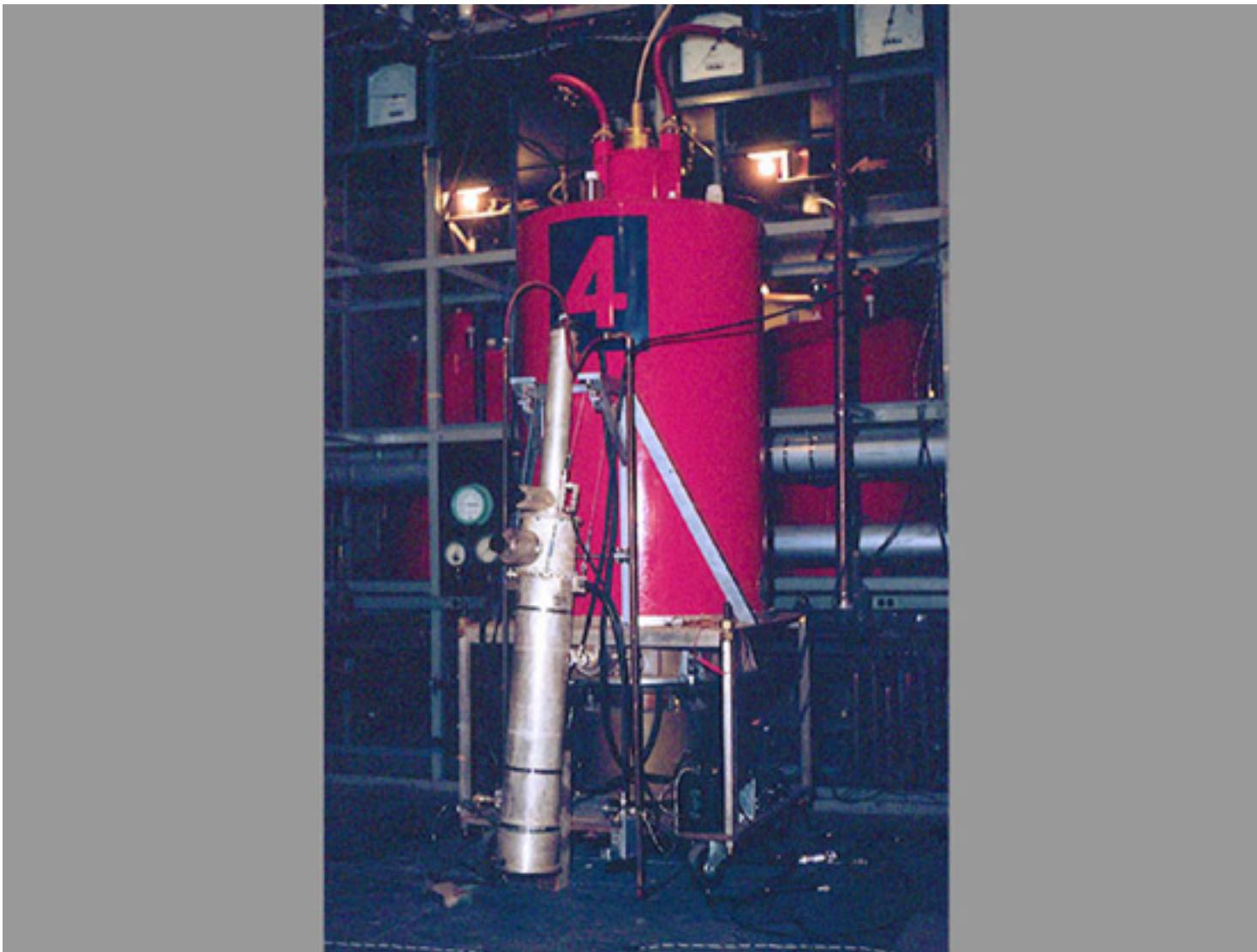
Aluminum dipoles still shinning





Ken bowles positioning a dipole with his bear hands.  
Young companion (Granger Morgan) has an easier job

# Power Amplifier: Driver and Final One of Four units



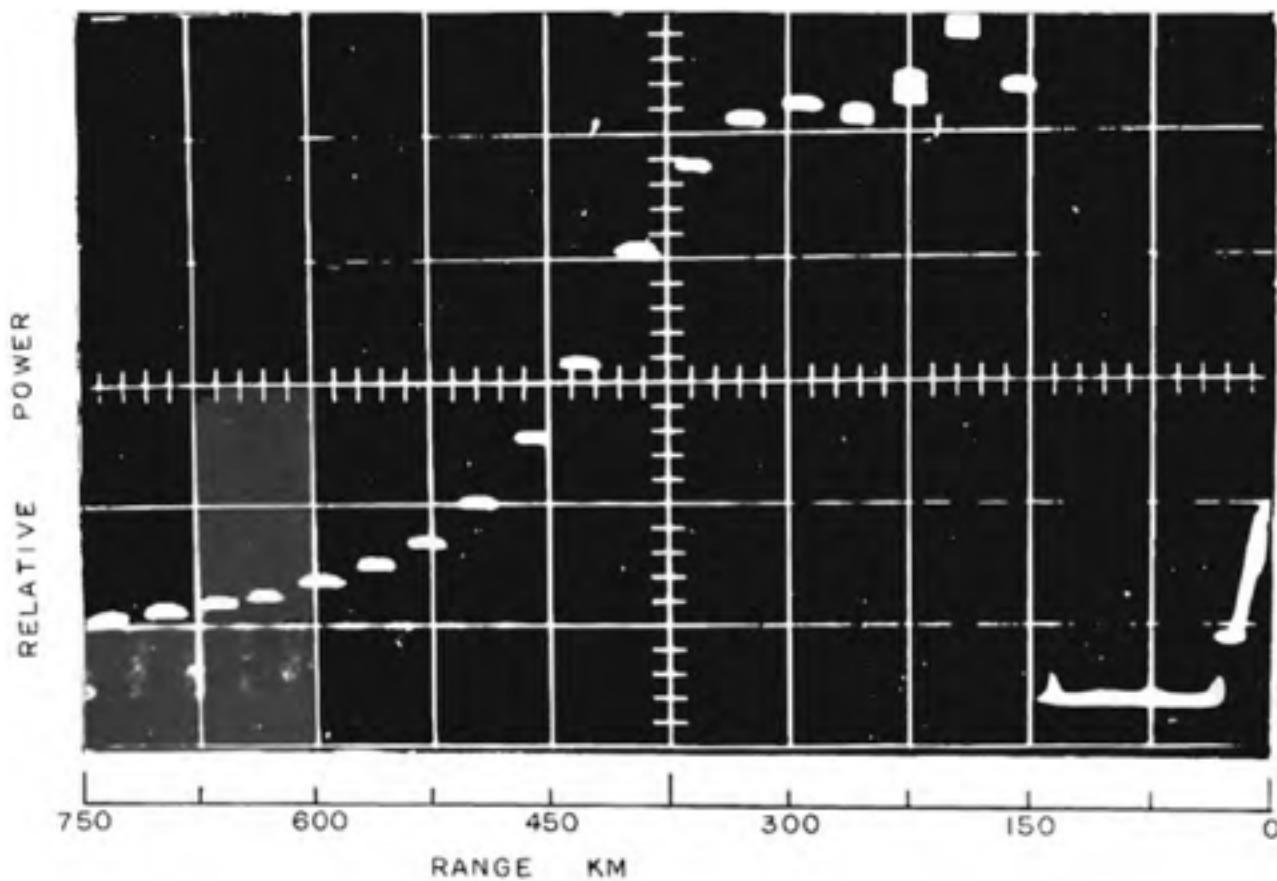


Figure 6 - Representative A'scope record with integration

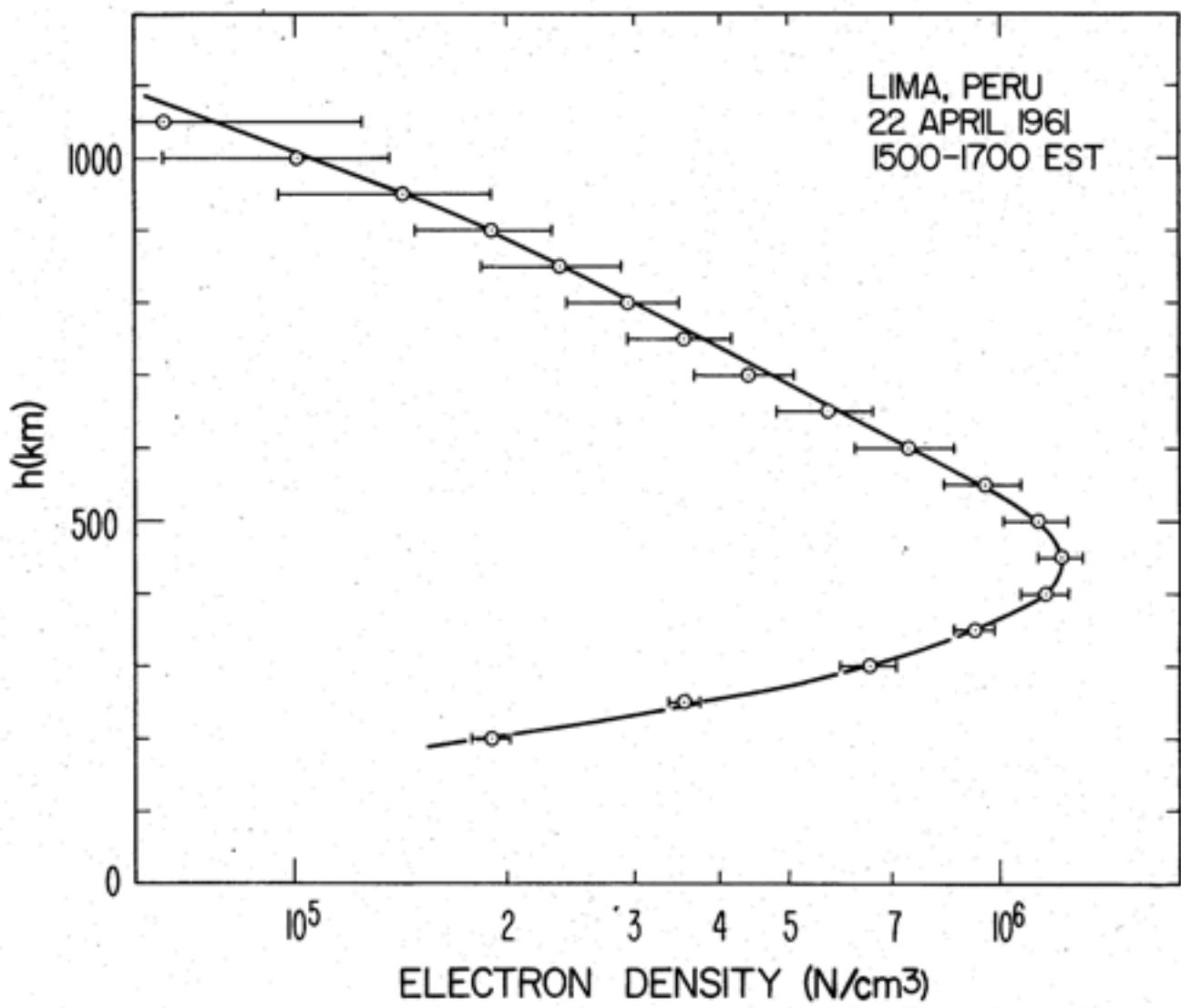
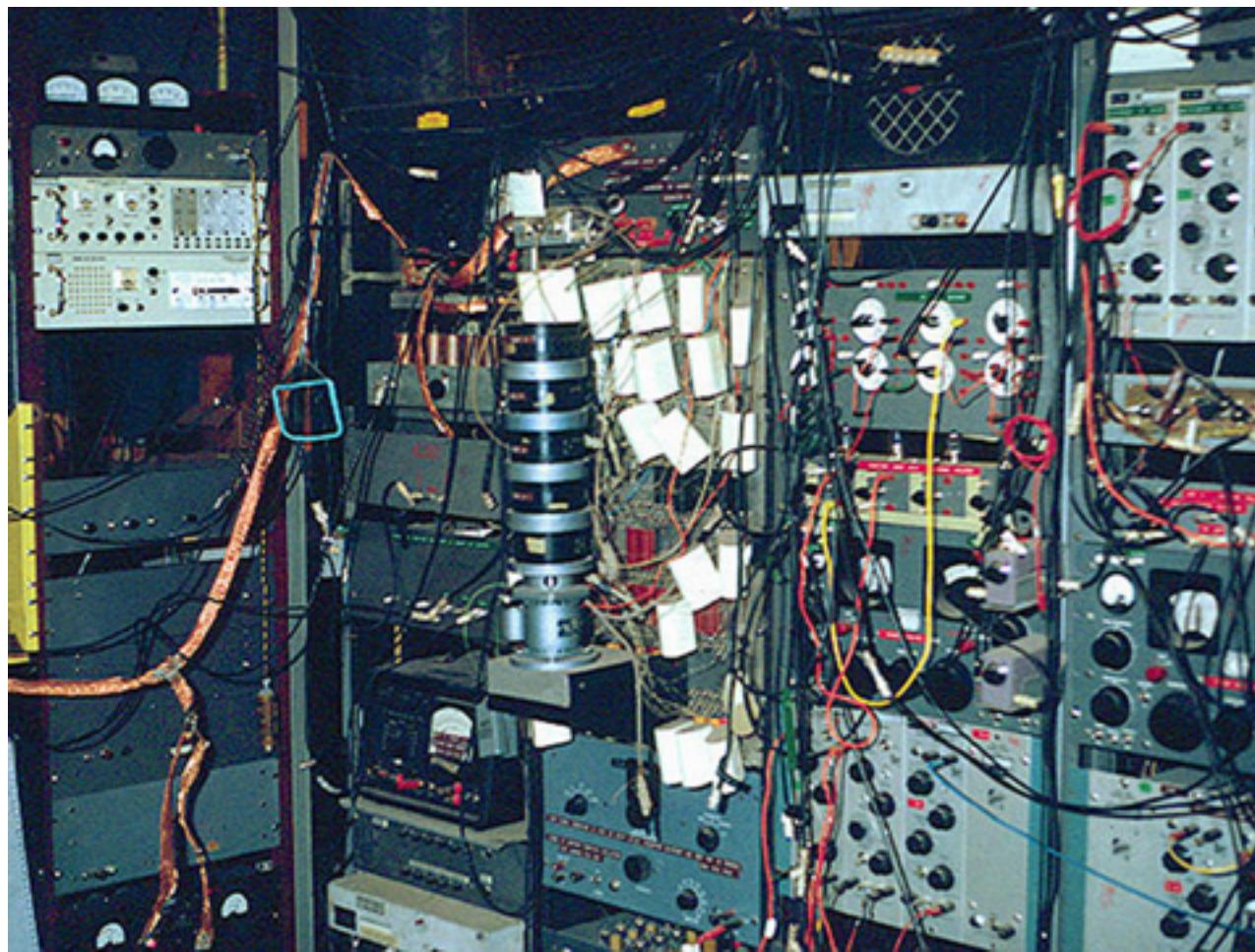


Figure 7

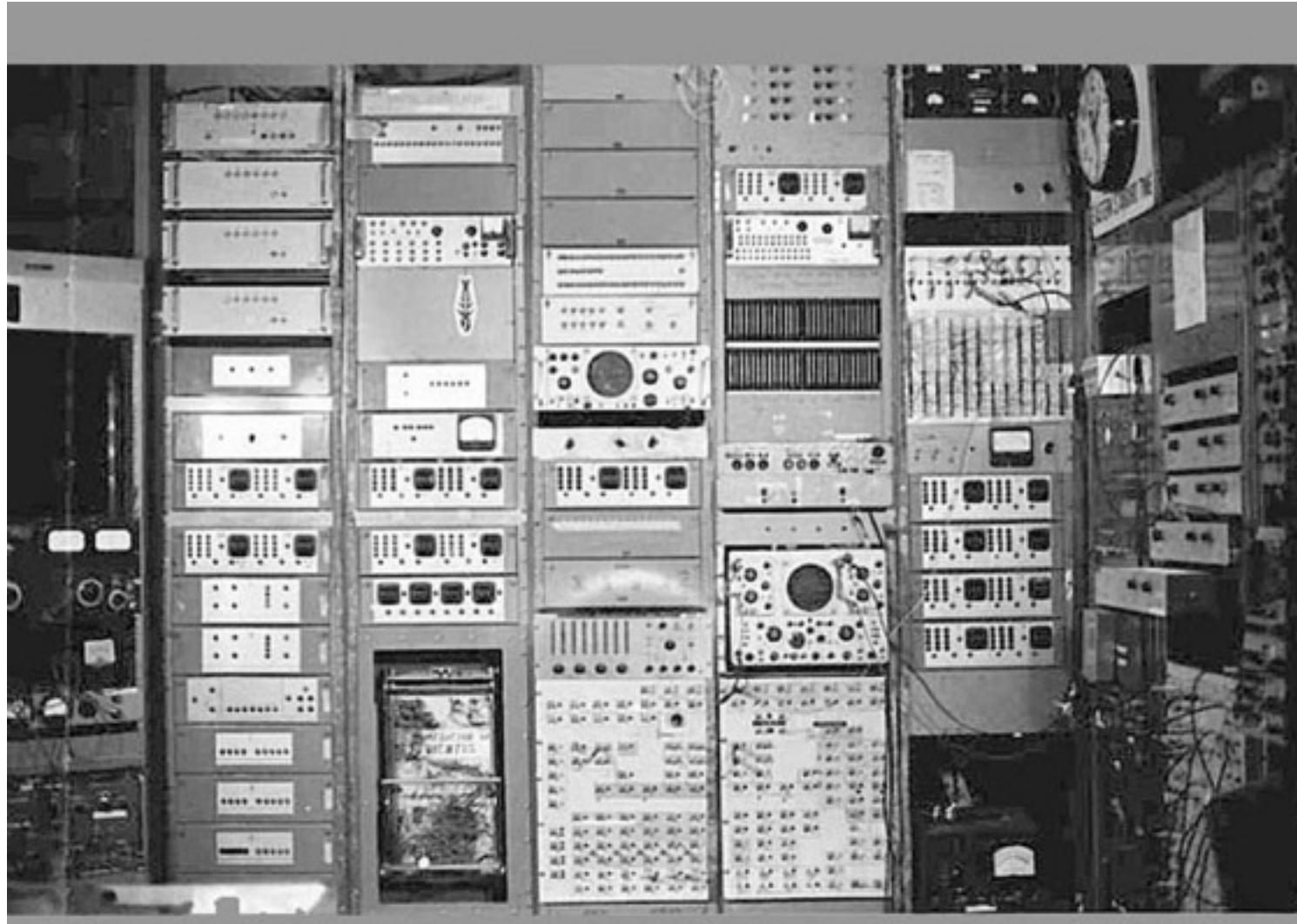
# Receiver rack, 1960's, note mercury switch

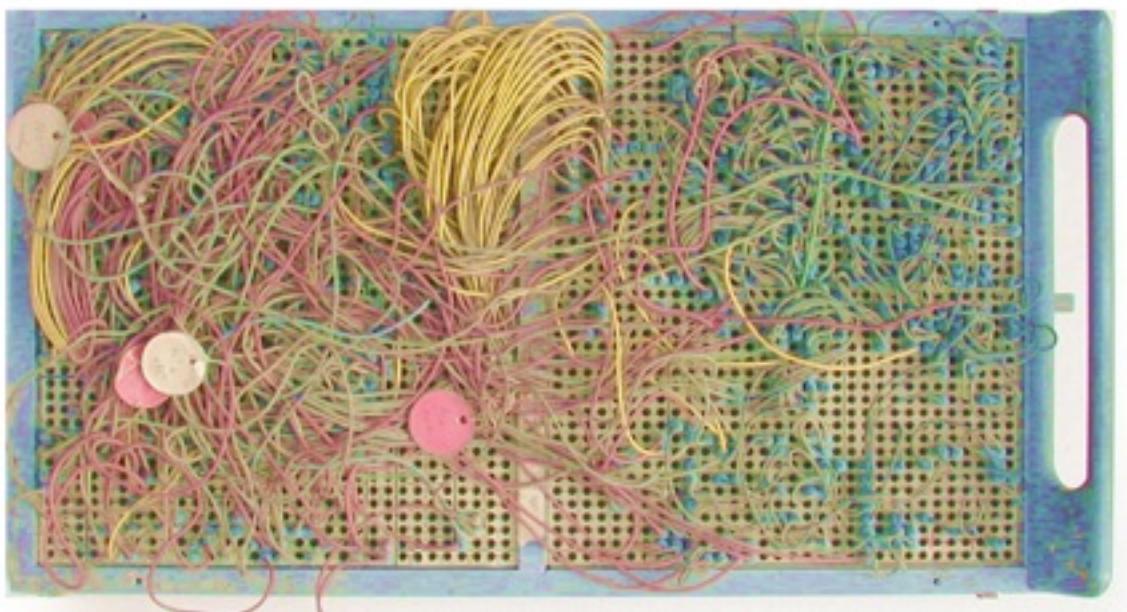
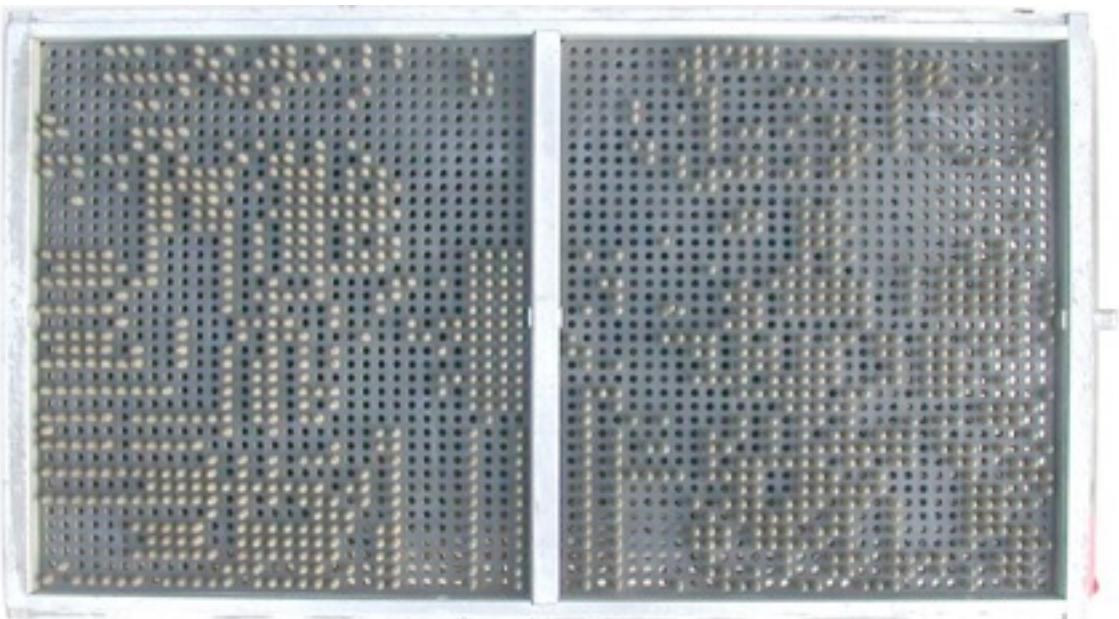


# Mercury jet commutator



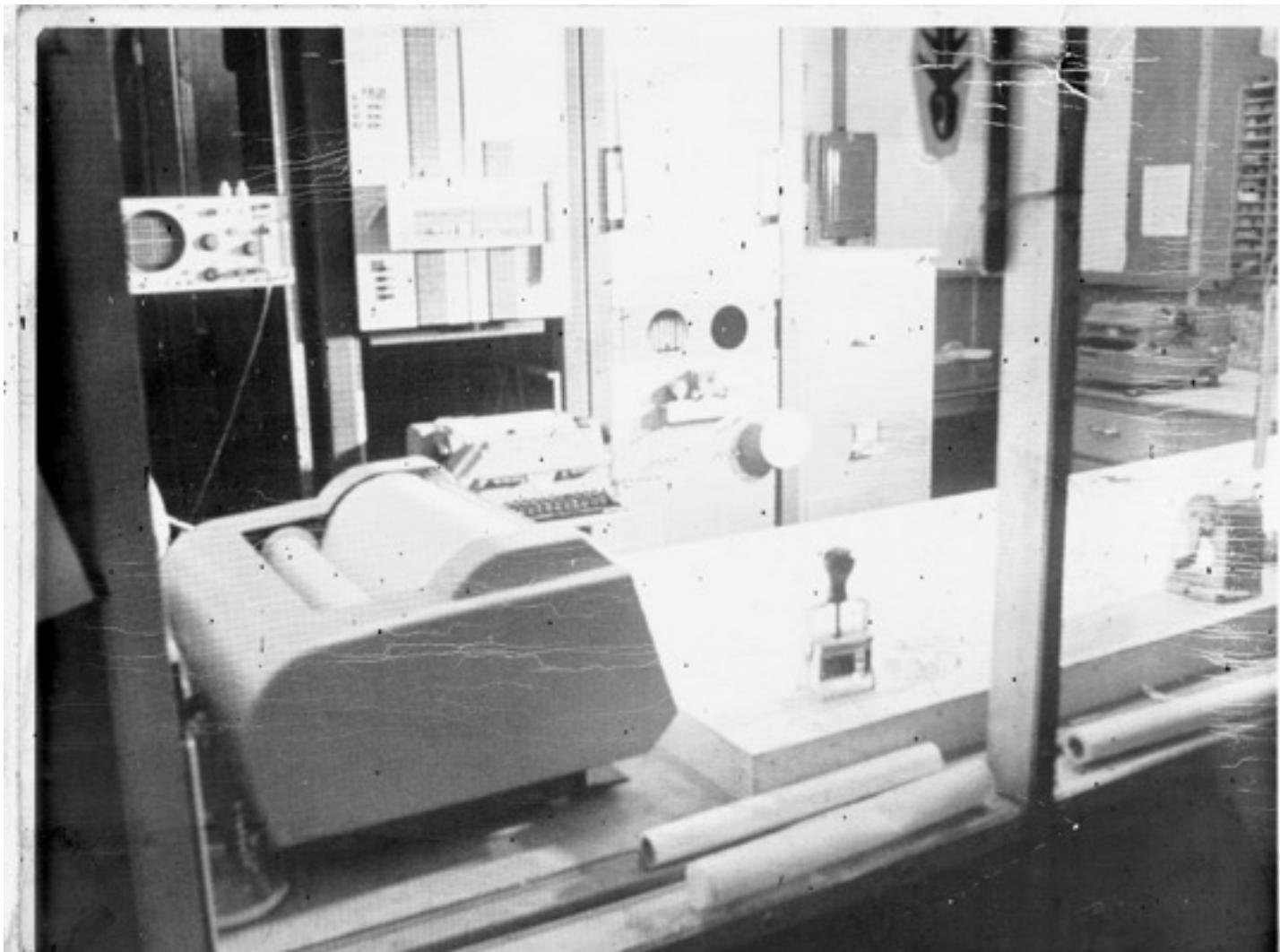
# The correlator





Plug in board

# PB250



## pb Packard Bell Computer

## PG 250 DRIVING SHEET

PROGRAM NAME db scalePAGE 6 of 5

PROGRAM NO.

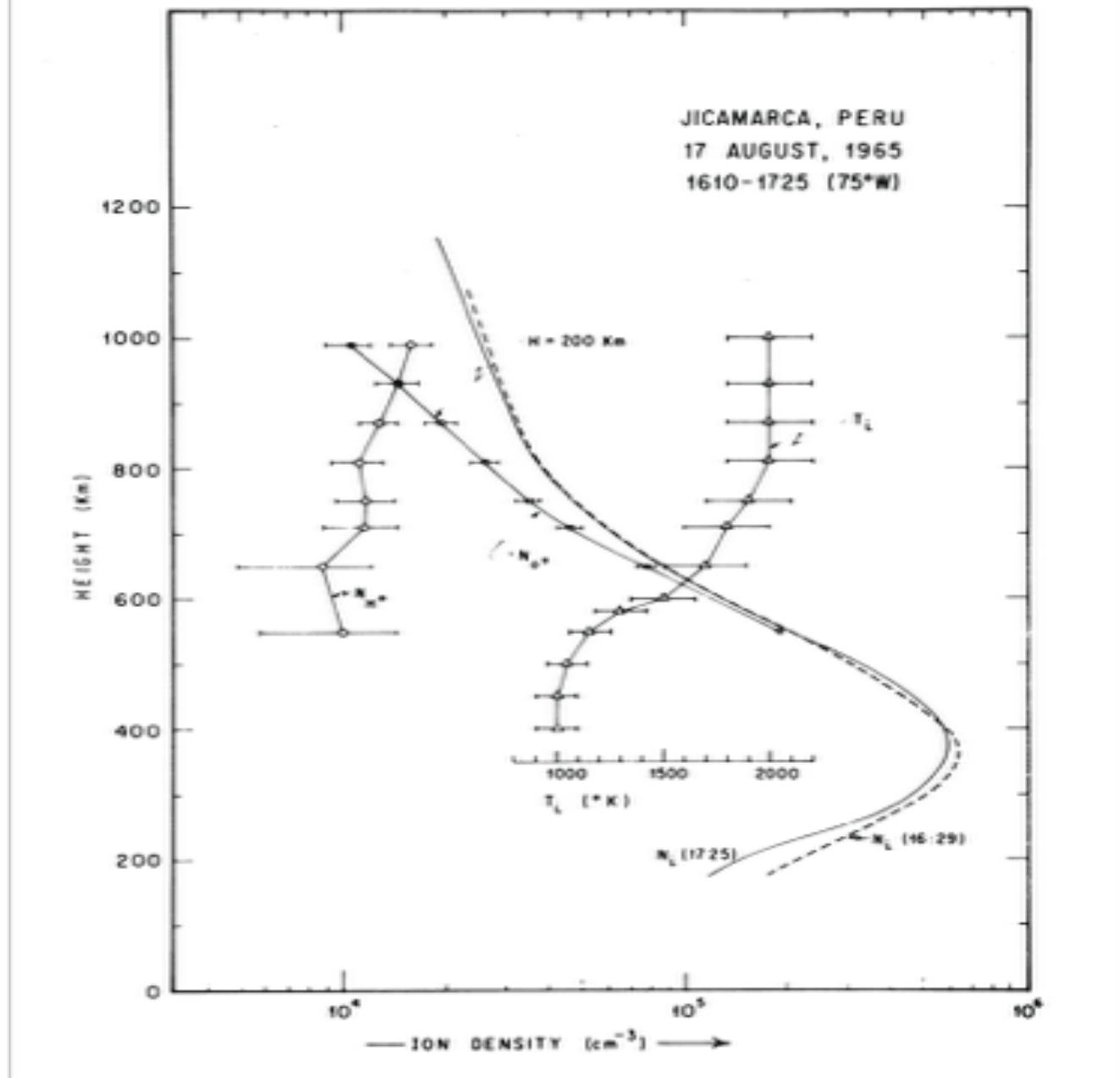
SPECIAL I.D.

RE

DATE

LINE <u>11</u>	INSTRUCTION					NOTES					
	SECTOR	S	OP.	LINE	I	SYMBOLIC OF CODE	LOCATION M	A REGISTER	B REGISTER	C REGISTER	INDEX REG.
140	141	5	04	11	;	LOC					
141	+000	004	00			count					
142	362	60	31	;		MOC -%					
143	144	5	15	11	;	348					
144	+000	000	01			count					
145	147	35	11	;		TAN					
146	140	5	37	11	;	TRU					
147	150	5	05	11	;	LDA					
150	+000	000	04			count half length of tab					
151	152	5	04	11	;	LOC					
152	+000	004	00			count					
153	353	60	21	;		MOC +%					
154	155	5	15	11	;	348					
155	+000	000	01			count					
156	160	35	11	;		TAN					
157	151	5	37	11	;	TRU					
160	073	05	04	;		LDA					
161	208	35	11	;		TAN					
162	072	15	04	;		SUB					
163	164	5	15	11	;	SUB					
164	+000	000	01			count					
165	172	35	11	;		TAN					
166	167	5	04	11	;	LOC					
167	+000	004	00			count					
170	370	60	24	;		MOC +4					
171	163	5	37	11	;	TRU					
172	160	05	11	;		LDA					
173	174	5	14	11	;	A00					
174	001	00	00			count into increment					
175	160	11	11	;		STA					
176	162	5	04	11	;	ZER					
177	200	5	14	11	;	A00					





FARLEY, McCLURE, STERLING, AND GREEN

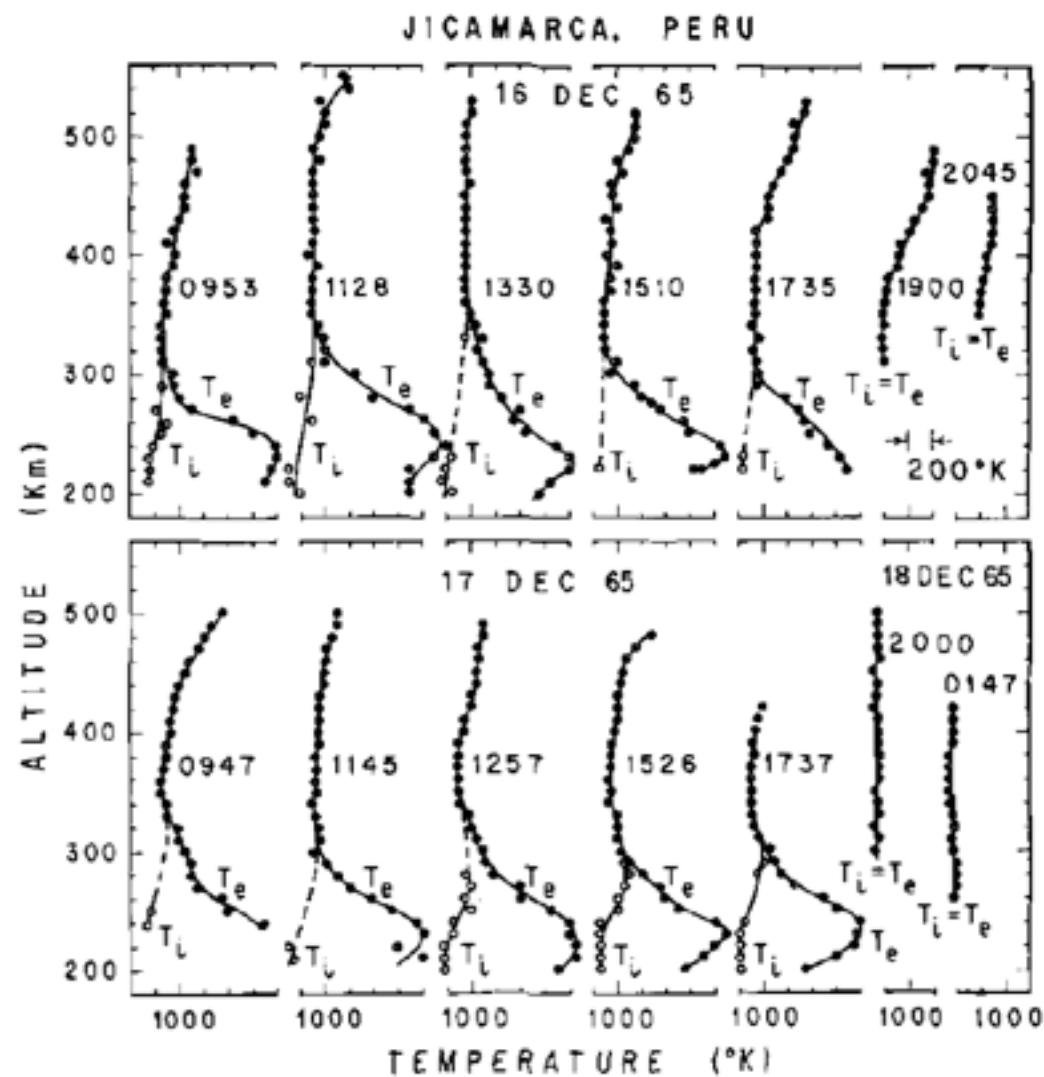


Fig. 5. A series of summer profiles of electron and ion temperatures.

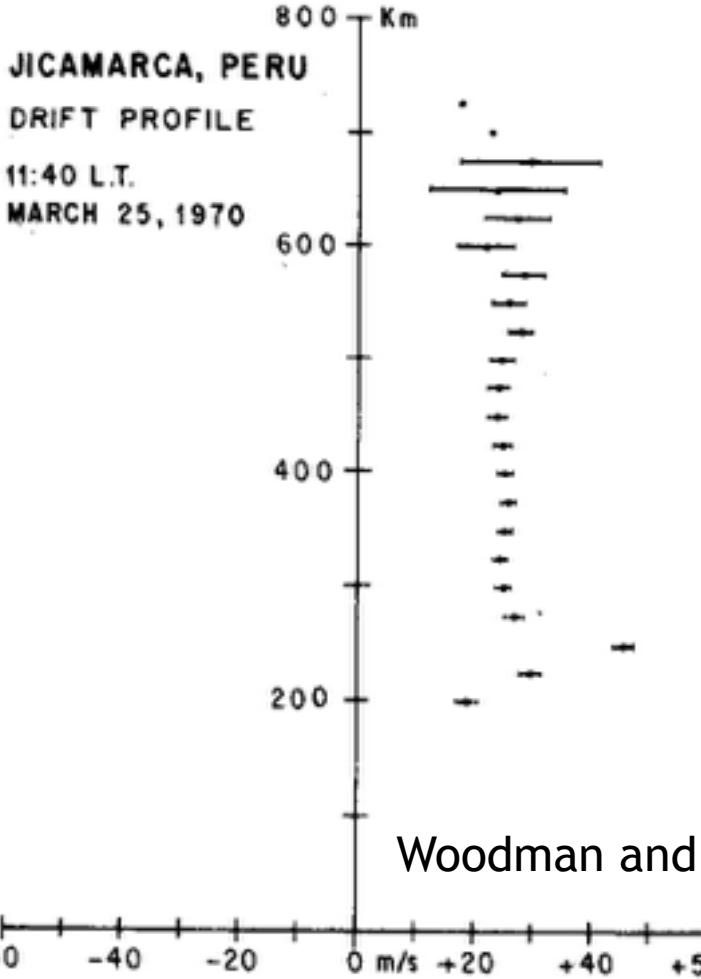


Fig. 1. Typical vertical velocity profile record obtained 'on line' at Jicamarca with 10 min of integration. The three lowest points are contaminated by strong electrojet echoes received through a side lobe of the antenna.

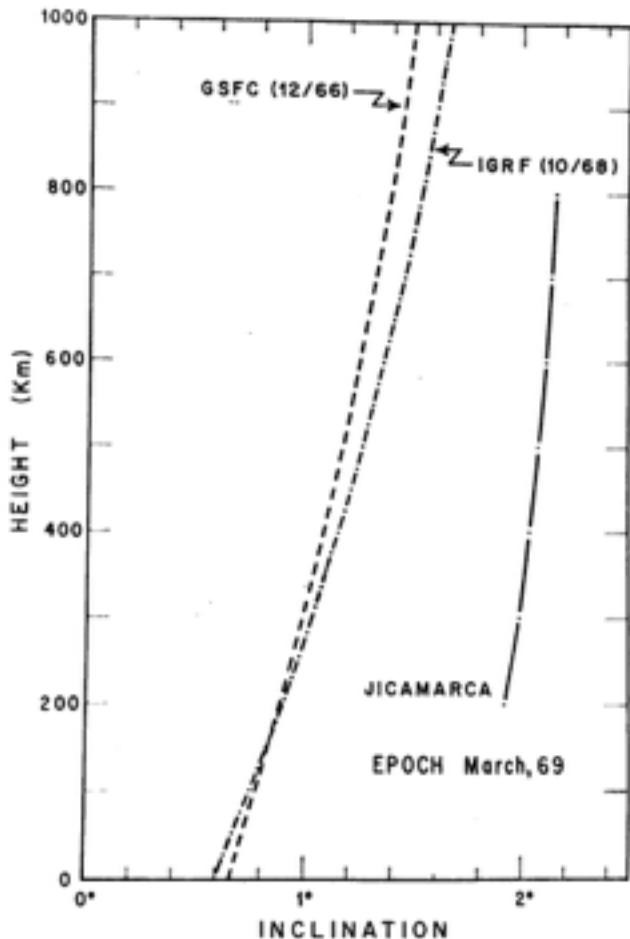
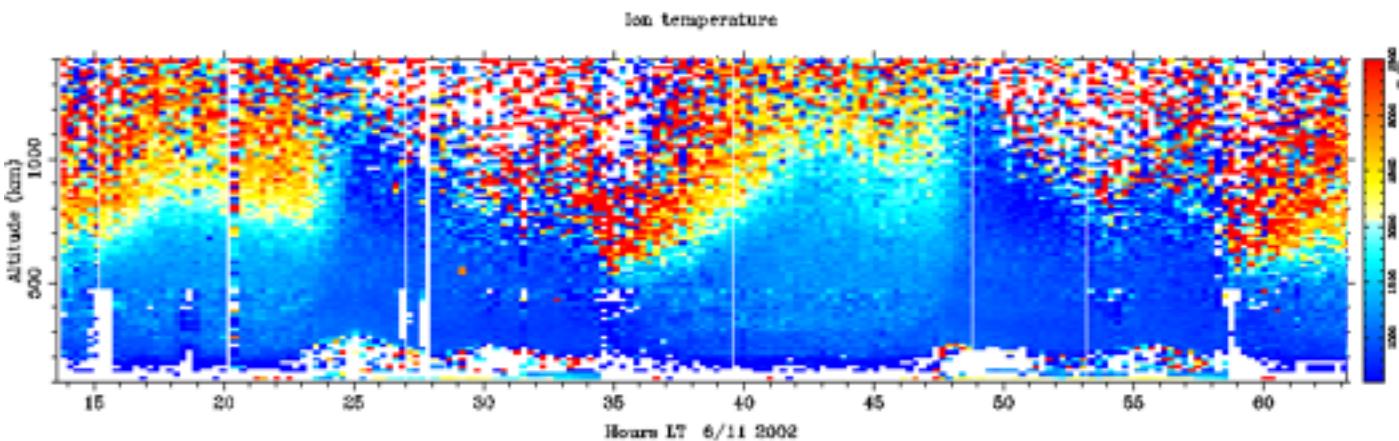
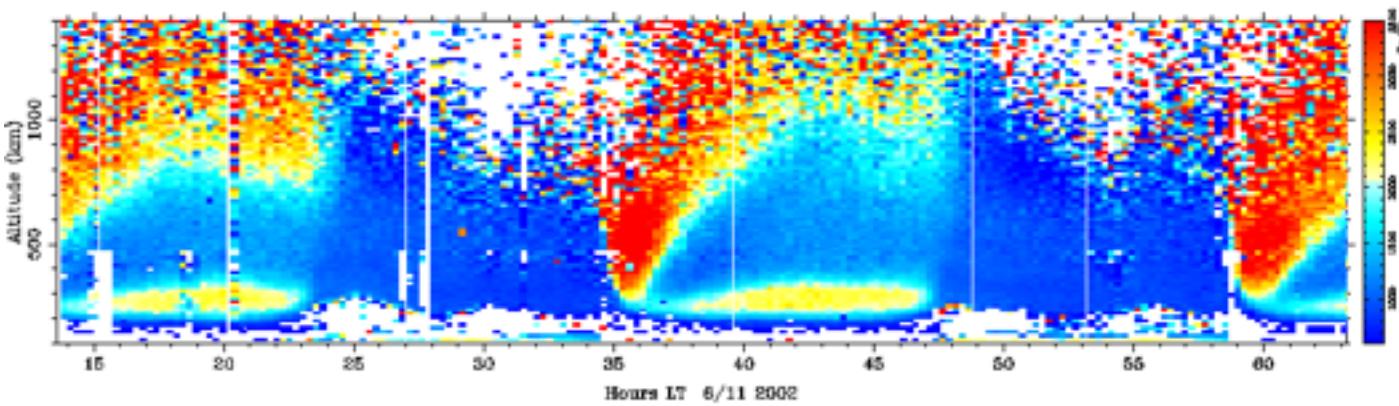
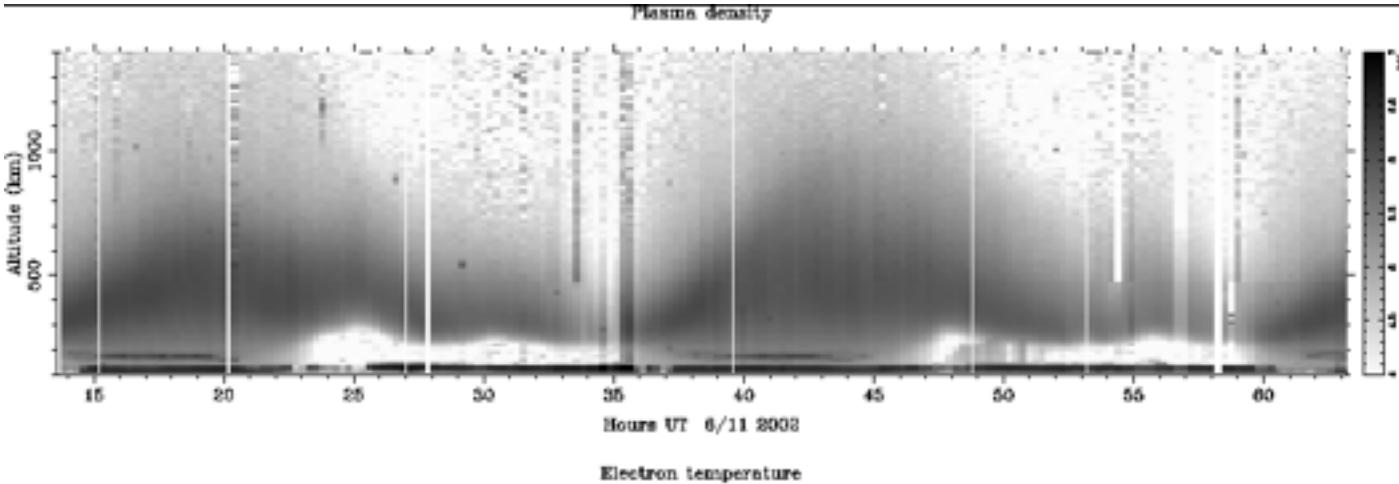


Fig. 6. Inclination of the magnetic field along Jicamarca vertical ( $-11.95^{\circ}$  latitude  $76^{\circ}52'20''$  longitude). The interrupted solid line corresponds to values determined experimentally at Jicamarca; the other two correspond to two of the latest earth magnetic field models, GSFC 12/66 and IGRF 10/68.

Ne  
Te  
Ti



Jicamarca, 1967, Elvis Presely?





En 1969, se transfiere el ROJ al IGP

- Some other contributions:
  - Spread F
  - *Electrojet Irregularities*
  - Other radars
  - + Rocket Campains

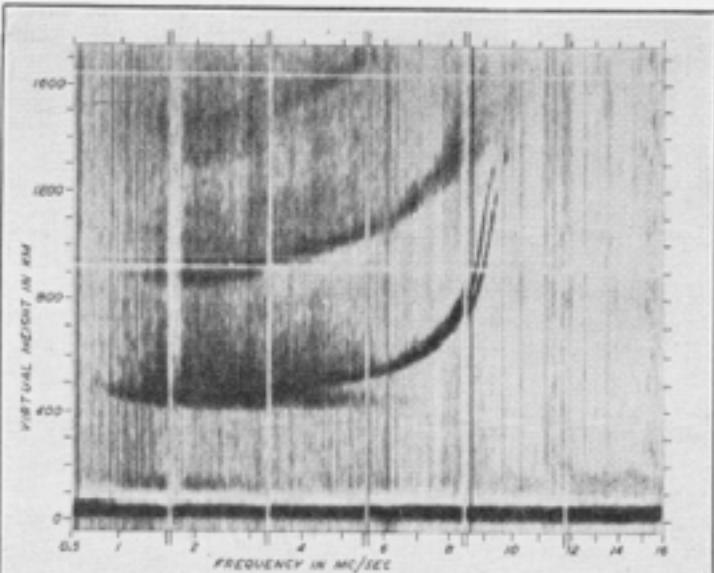


FIG. 1—RECORD SHOWING REGULAR AND DIFFUSE F-REGION ECHOES, HUANCAYO MAGNETIC OBSERVATORY, FEBRUARY 14, 1938, 20°15'-20°30', 75° WEST MERIDIAN TIME

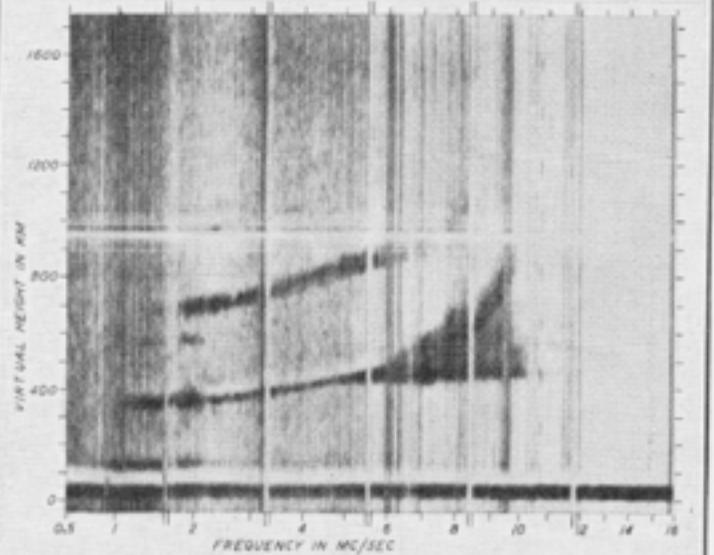
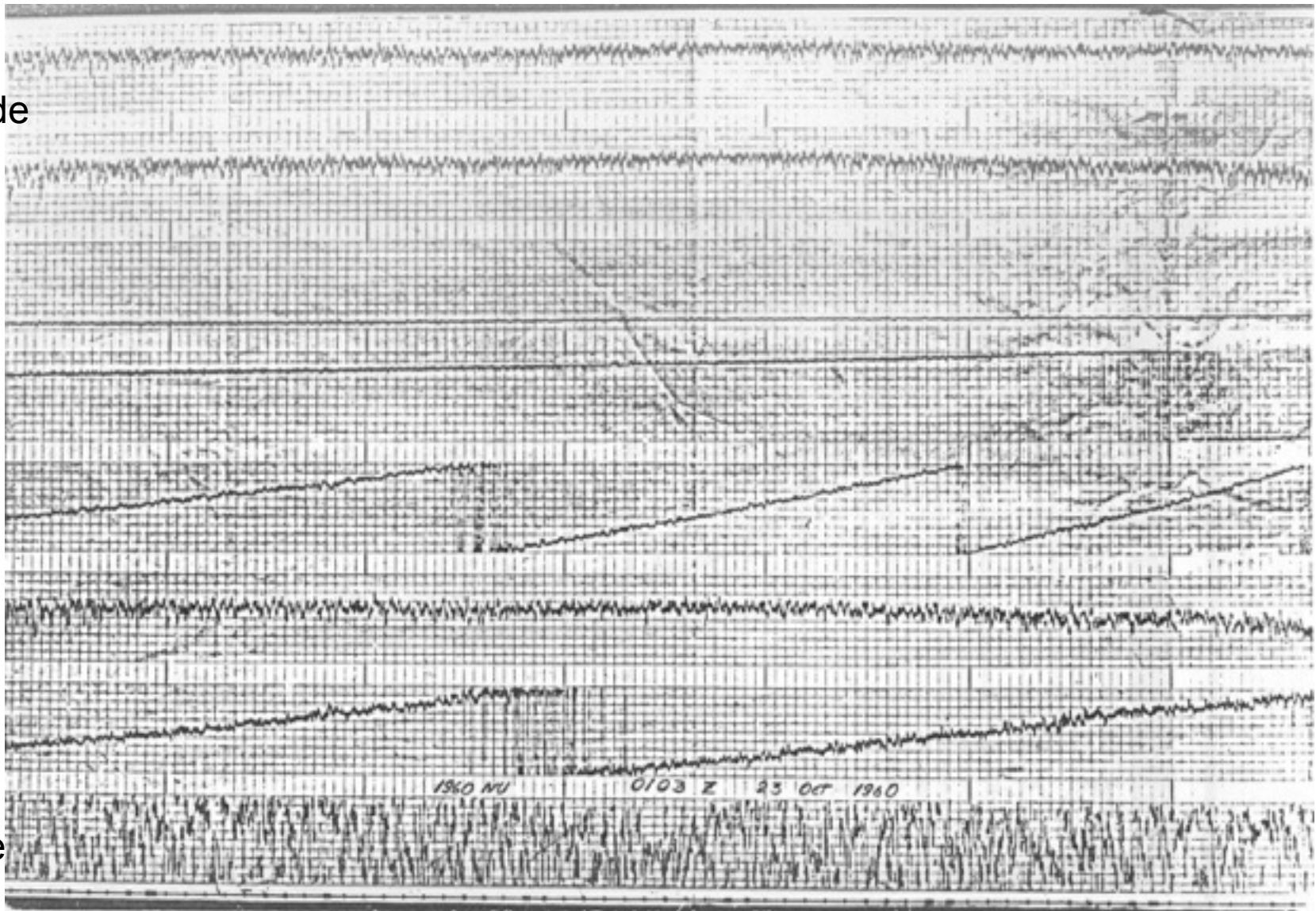


FIG. 2—RECORD SHOWING REGULAR AND DIFFUSE F-REGION ECHOES, HUANCAYO MAGNETIC OBSERVATORY, FEBRUARY 23, 1938, 21°30'-21°45', 75° WEST MERIDIAN HOURS

Amplitude

NS Fine

EW Fine



Satellite phase scintillation, Woodman. 1960

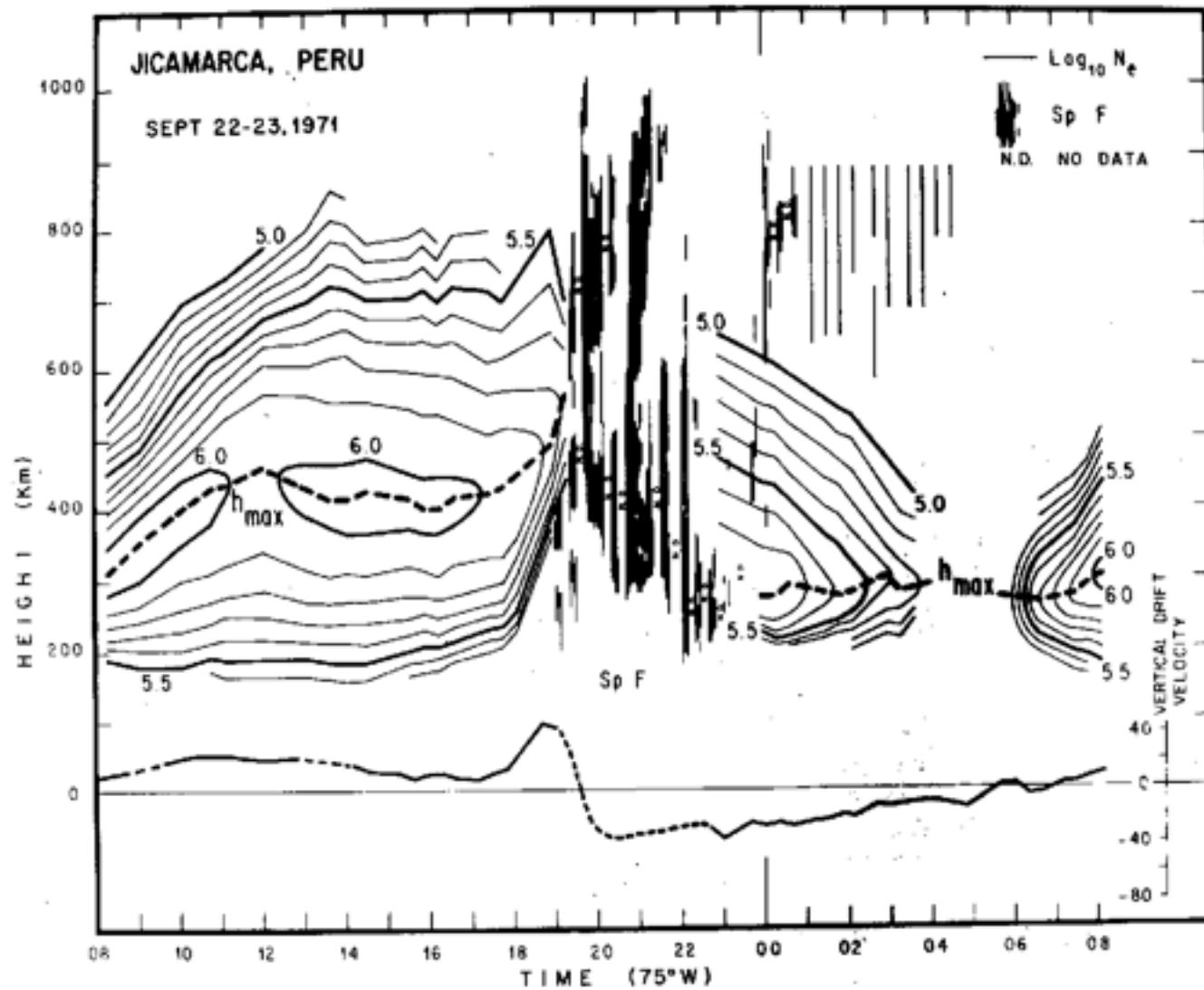


Fig. 2. Altitude range and time of occurrence of spread *F* echoes and their relation to the electron density, electron density gradients, and vertical drifts at *F* region heights.

After Woodman and La Hoz, 1976.  
Illustrates Farley et al, 1970, conclusions

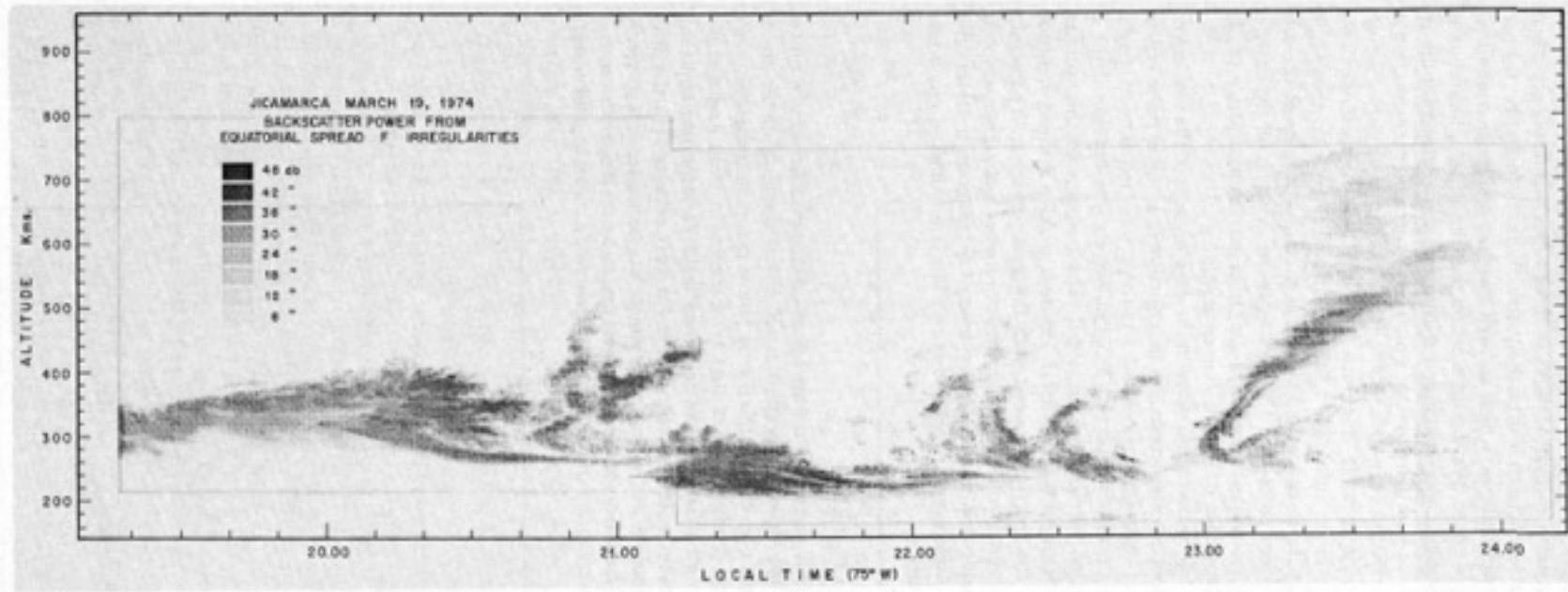
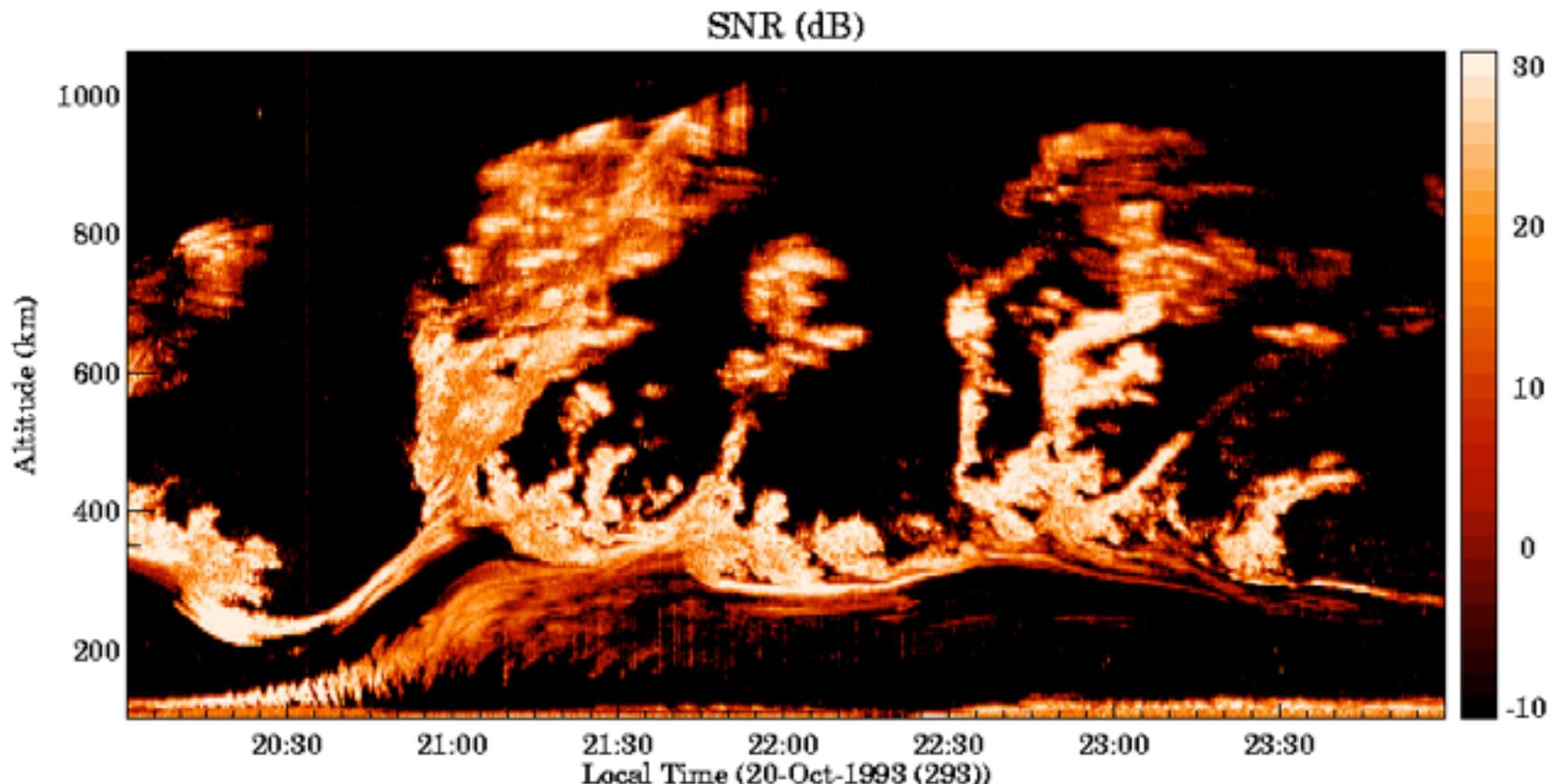


Fig. 3c

Woodman and La Hoz, 1976

# ESF echoes

(from *Woodman and Chau [2001]*)



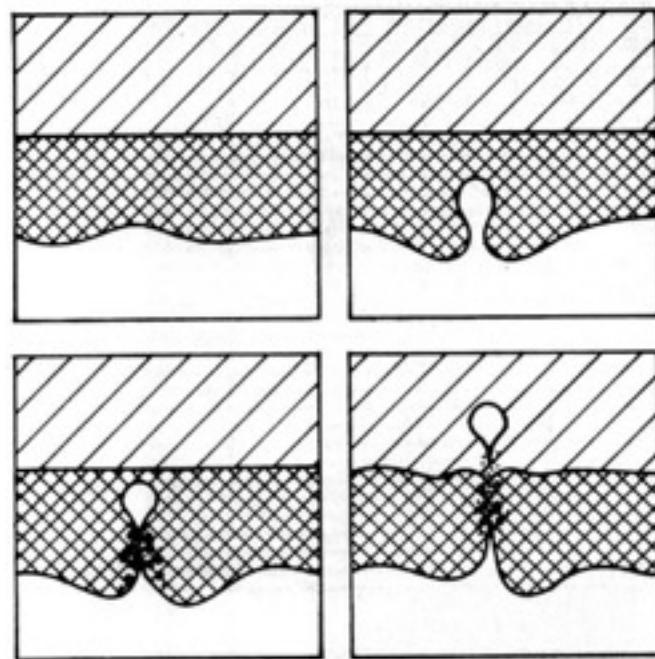
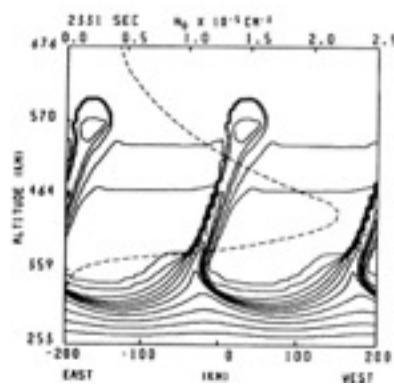
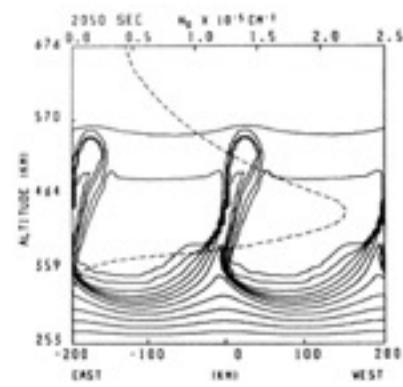
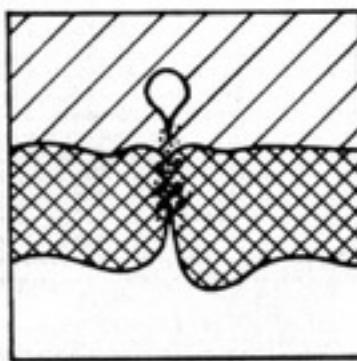
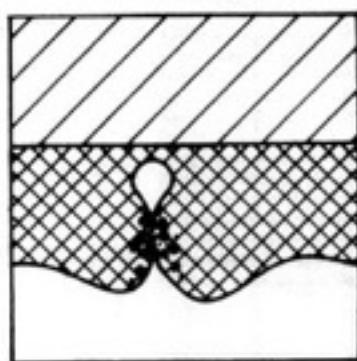
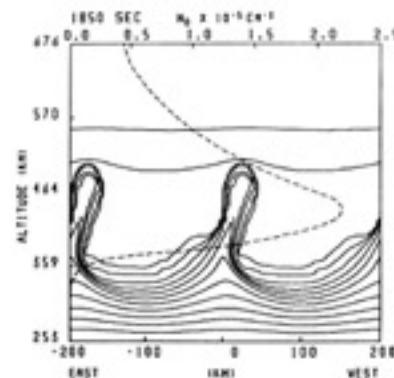
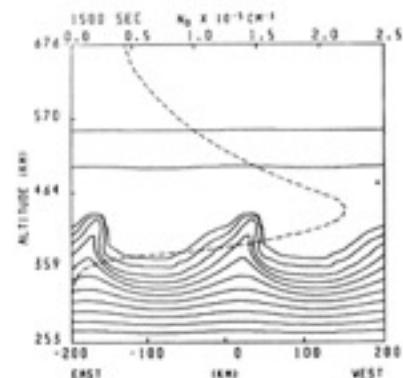
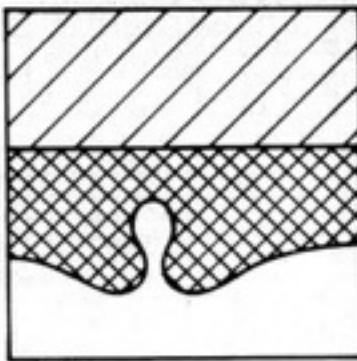
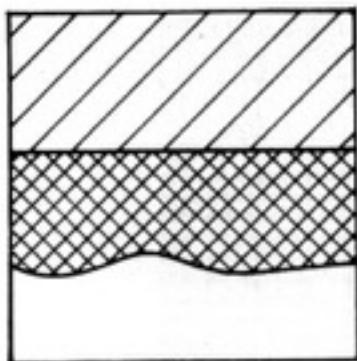


Fig. 9. Schematic representation of a three-density model of the ionosphere showing the formation of a bubble of low electron density and its propagation to the gravitationally stable top. The middle fluid is heavier than the top, and the top fluid heavier than the bottom.

Woodman and La Hoz, 1976



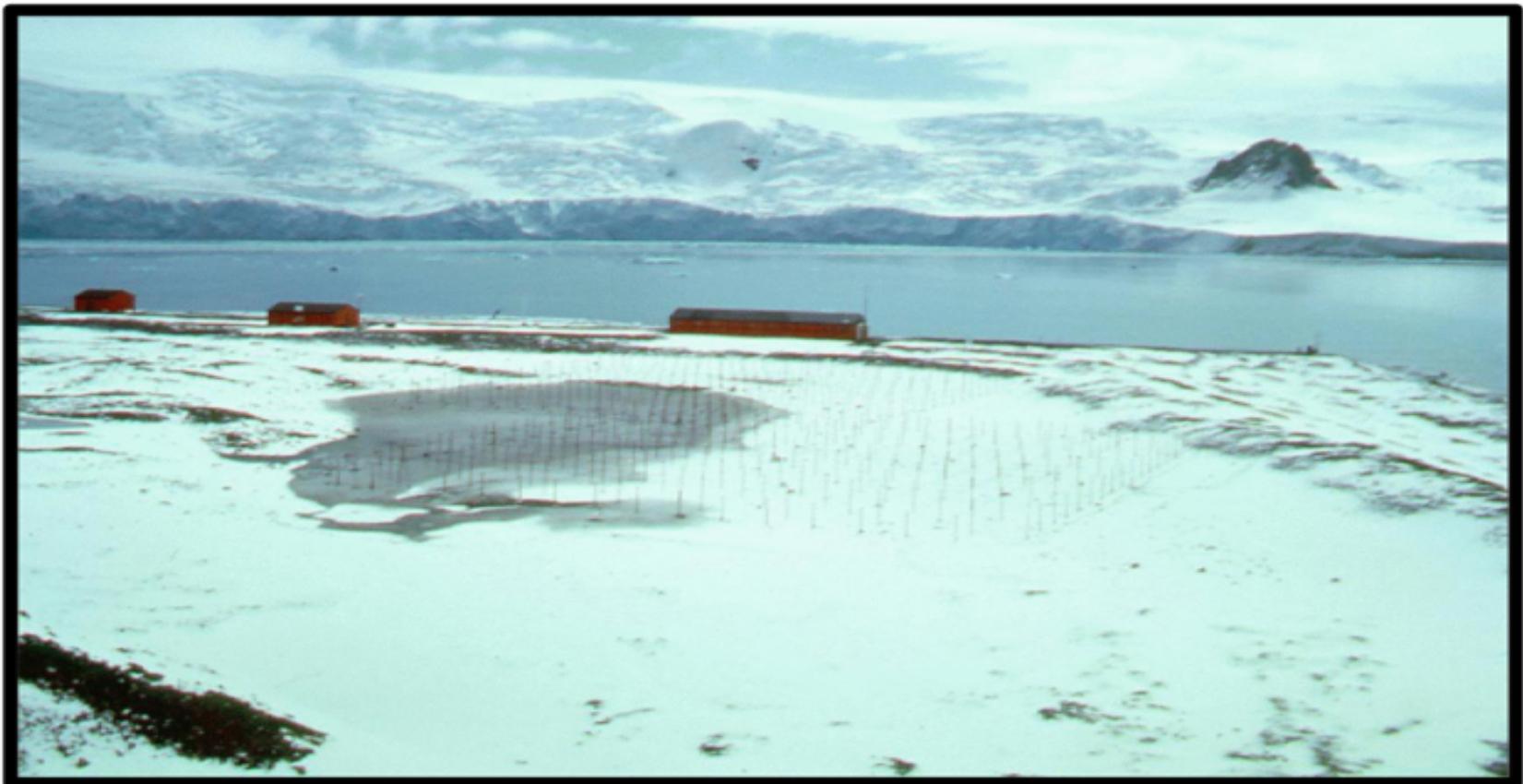
# Piura ST

# EOS

Transactions, American Geophysical Union  
Vol. 71 No. 50 December 11, 1990



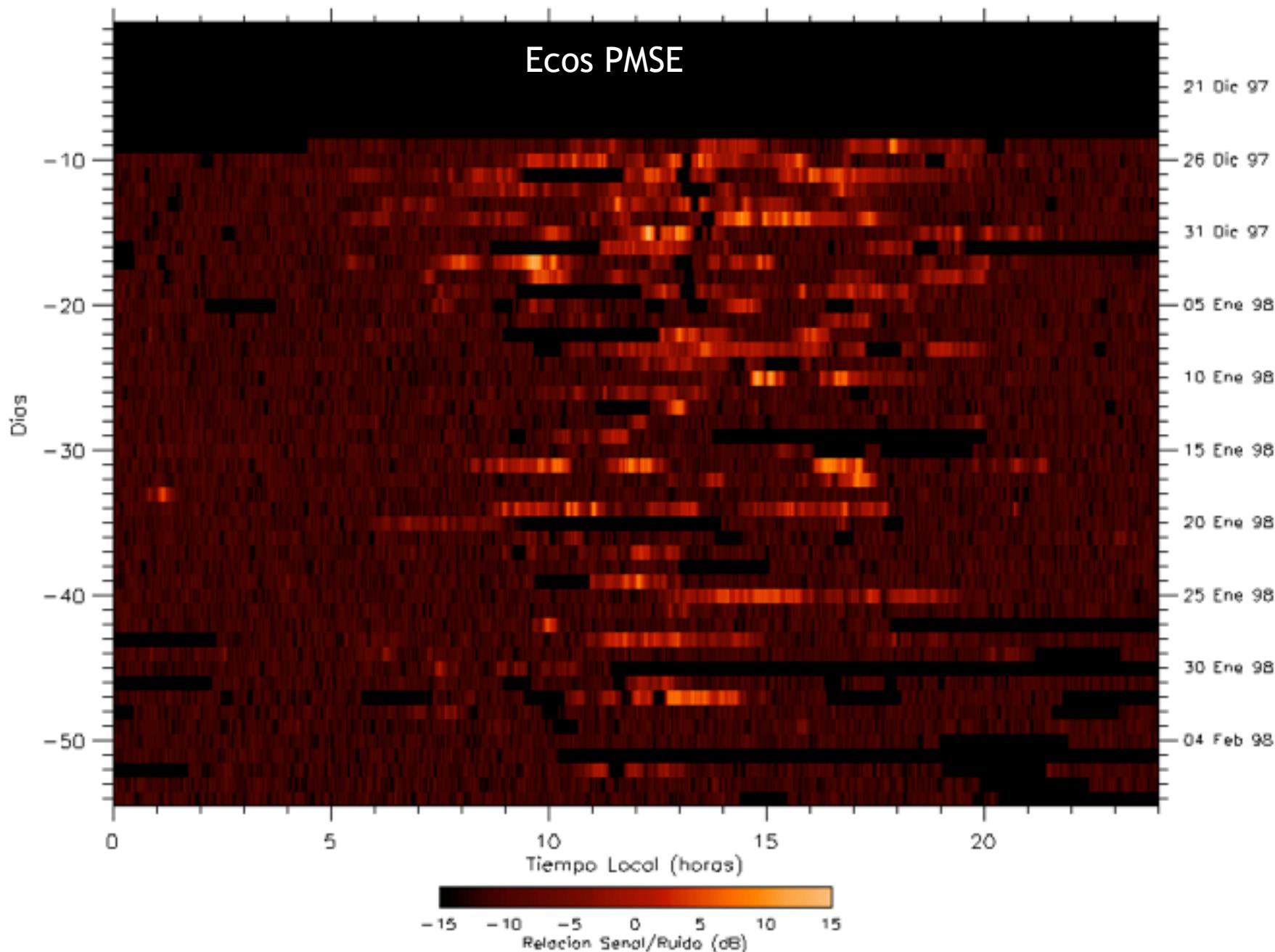
# RADAR ANTARTICO “MST”

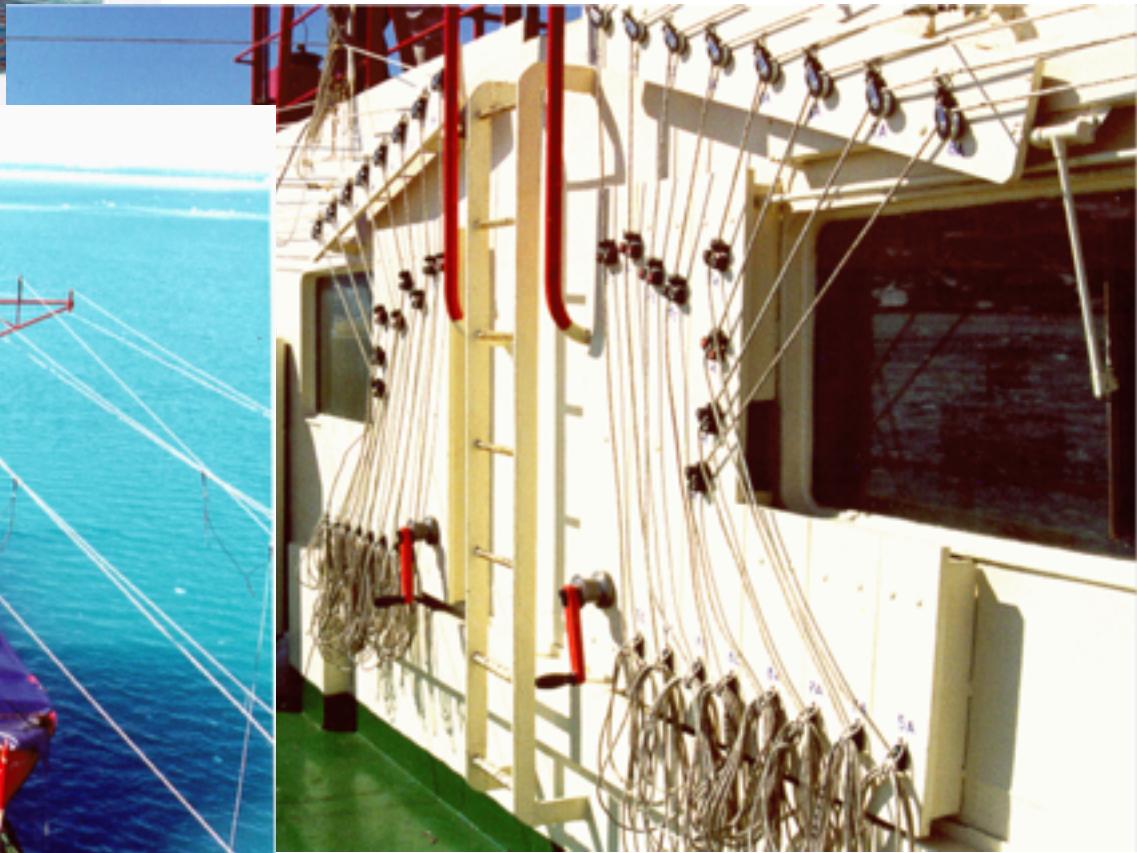


IX EXPEDICION CIENTIFICA  
*Estación Científica “MACHU PICCHU”*

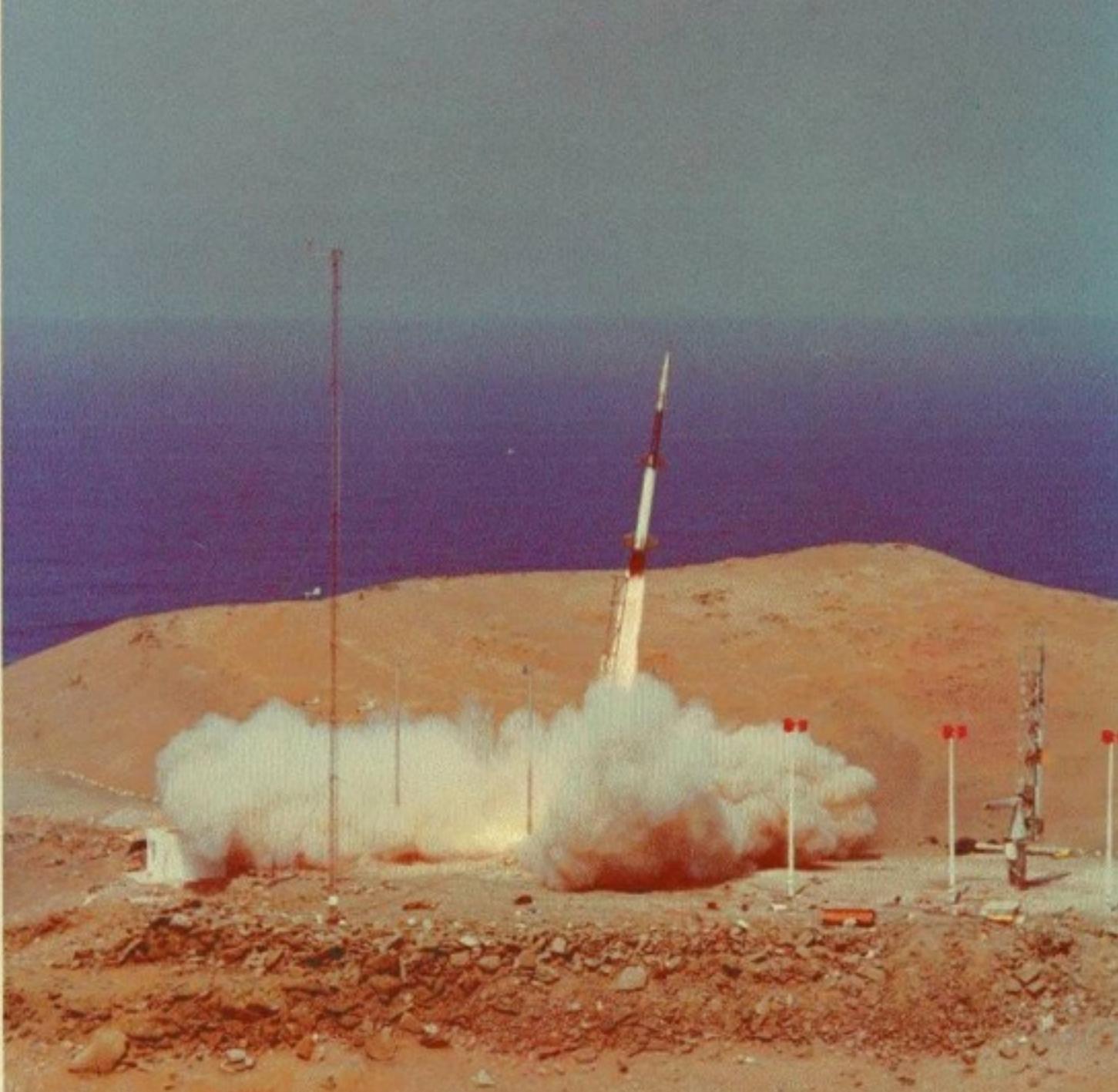
## Potencias Maximas – Beam Vertical

Ecos PMSE









Thank you