

Advanced Modular Incoherent Scatter Radar

Roger H. Varney

¹Center for Geospace Studies
SRI International

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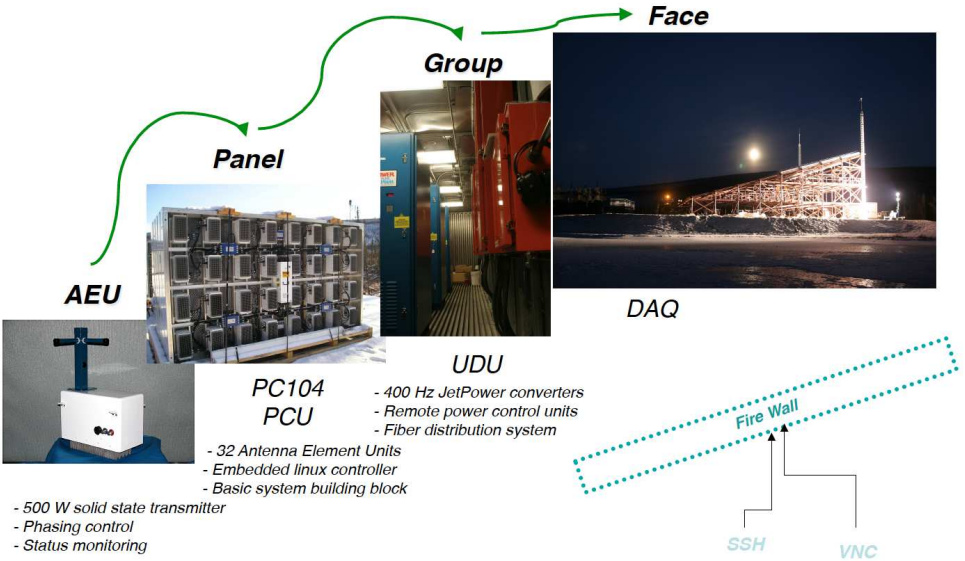
1 AMISR Technology

- Hardware
- Electronic Beam Steering Capabilities
- Design Considerations for ISR Science

2 AMISR Science

- PFISR Science
- RISR-N Science
- RISR-C First Results
- UMET-14 Science

AMISR Modular Design



Antenna Element Unit (AEU) Specifications

- **Distributed Solid State Power Amplifiers (SSPAs)**
- **430-450 MHz instantaneous bandwidth**
- **10% Maximum duty cycle**
- **Minimum PRF interval 500 usec**
- **Maximum pulsewidth 2 msec**
- **Passive cooling (no moving parts)**
- **400 Hz prime power**



- **Crossed dipoles, circular polarization on axis**
- **Balun built into the antenna support shaft**
- **Constant impedance over bandwidth and scan angle**
- **Spacing is hexagonal for efficiency**
- **Tx/Rx polarizations are opposite and fixed (not measurable)**

The AMISR UHF System

AMISR AEU = Tx/Rx Unit

- 500 W solid state transmitter
- Phasing control
- Status monitoring
- 4096 AEU's/AMISR radar face

Antenna Element Unit (AEU)



32/panel

AMISR Panel

- 32 Antenna Element Units arranged in hexagonal pattern
- 3.5 x 2 meters; 19.8 dBi / panel
- 16 kW peak power per panel
- Basic system building block for AMISR
- Embedded linux controller



Panel (with PCU)



Face



Utility Distribution Unit (UDU)



AMISR Control System (ACS)

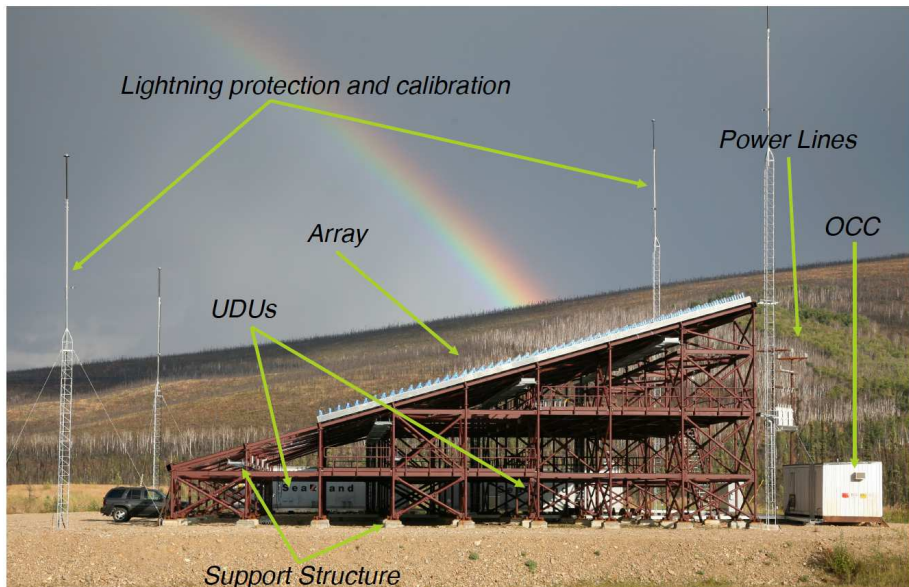
AMISR UDU

- 400 Hz JetPower converters
- Remote power control units
- Fiber distribution system

AMISR ACS

- Flexible transmit and receive system
- Completely remotely controlled
- Experiments run off a scheduler

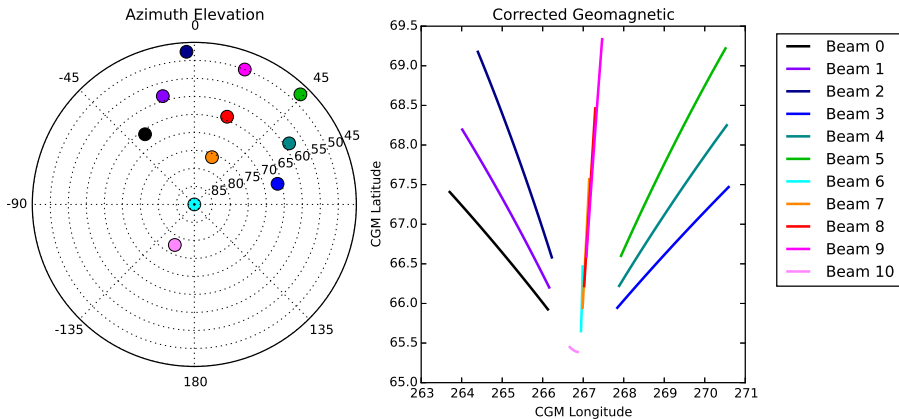
Poker Flat Incoherent Scatter Radar (PFISR)



AMISR Technical Specifications

- Peak Power: 2 MW
- Max RF Duty: 10%
- Pulse Length: 1 μ sec - 2 msec
- TX Frequency: 430-450 MHz
- Antenna Gain: ~ 43 dBi
- Antenna Aperture: ~ 715 m²
- Beam Width: $\sim 1.1^\circ$
- System temperature: ~ 120 K
- Steering: Pulse to pulse over $\sim \pm 25^\circ$
- Max system power consumption: ~ 700 KW
- Max operations: continuous, depending on power availability
- Unattended operations
- Data volume ~ 6 TB/year at Poker Flat
- No moving parts on the antenna
- Environment: -40° C to $+35^\circ$ C
- Altitude coverage: ~ 60 km to ?? km (depending on Ne)
- Minimum measurable electron densities: $\sim 1e9$ m⁻³
- Typical time resolution:
 - E region $< \sim 3$ min,
 - F region $< \sim 1$ min,
 - ~ 10 look directions and typical ionospheric conditions - many caveats apply!
- Typical range resolution: 600 meters to 72 km (mode dependent, can be extended)
- Plasma parameters: Ne, Te, Ti, Vi, v_{in} , composition
- Derived parameters: E, J, J-E, J-E', Un, σ_p , σ_H

Electronic Beam Steering



- Time to steer beam is $\sim 400 \mu\text{s}$. Less than typical IPP (1-10 ms).
- Beam steering happens pulse-to-pulse.

Mechanical Steering vs Electronic Steering

Mechanical Steering Experiment

- Steer to position 1
- Send many pulses for 1 min
- Steer to position 2
- Send many pulses for 1 min
- Steer to position 3
- Send many pulses for 1 min
- Steer to back to position 1

Electronic Steering Experiment

- Pulse in beam 1
- Pulse in beam 2
- Pulse in beam 3
- ⋮
- Pulse in beam N
- Pulse in beam 1 again

Advantages of Electronic Steering:

- Each beam is revisited once every N_{TIPP} .
- Can average data at any multiple of N_{TIPP} . Incoherent integration time is adjustable after the fact.
- Data being combined is nearly simultaneous.
- No time lost steering between pulses.

Statistical Considerations with Pulse Steering

$$\frac{\delta\hat{S}}{S} = \frac{1}{\sqrt{K}} \left(1 + \frac{1}{S/N} \right)$$

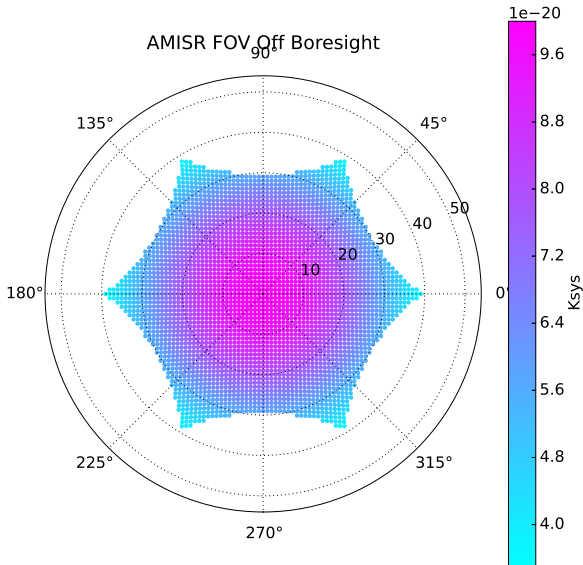
- Larger number of beams \Rightarrow fewer pulses per beam
- If I can comfortably integrate for 1 min using 7 beam positions \Rightarrow I would need to integrate 6 min to get the same data quality using 42 beams.

Multiple frequency channels can help statistics. Example: RISR-N “ImagingLP” mode for imaging F-region polar cap patches.

- 51 beam positions
- Long pulses on 3-frequency channels
- In each IPP 2 frequencies Tx, 3rd collects noise/cal samples
- Same statistics as a single frequency experiment on 17 beams

Limitations of Phased Array Beam Steering

- FOV limited by grating lobe limit $\sim 30^\circ - 40^\circ$
- Antenna gain decreases with steering angle off of boresight
- Antenna works best within $\sim 25^\circ$ off of boresight



Sizing an AMISR for ISR

Statistical Accuracy of ISR Measurements:

$$\frac{\delta \hat{S}}{S} = \frac{1}{\sqrt{K}} \left(1 + \frac{1}{S/N} \right) \approx \frac{1}{\sqrt{K}} \frac{1}{S/N}$$

Soft Target Radar Equation:

For an active phased array

$$\frac{S}{N} \propto P_{Tx} \frac{G}{4\pi R^2} \eta V_s \frac{A_{eff}}{4\pi R^2}$$

$$P_{Tx} \propto \text{Panels}$$

$$A_{eff} \propto \text{Panels}$$

$$G \sim \frac{4\pi}{\Omega} \quad V_s \sim R^2 \Omega$$

$$\frac{S}{N} \propto (\text{Panels})^2$$

$$K \propto (\text{Panels})^4$$

$$\frac{S}{N} \propto P_{Tx} \frac{1}{4\pi R^2} \frac{4\pi}{\Omega} \eta R^2 \Omega \frac{A_{eff}}{4\pi R^2}$$

$$\frac{S}{N} \propto \frac{1}{4\pi R^2} P_{Tx} A_{eff} \eta$$

1 min integration with 128 panels \Rightarrow
16 min integration with 64 panels

Existing AMSIR Fields of View



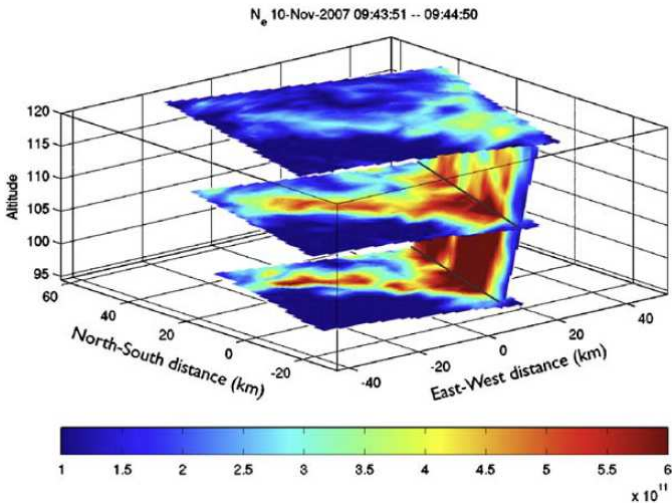
PFISR Field of View at Multiple Altitudes



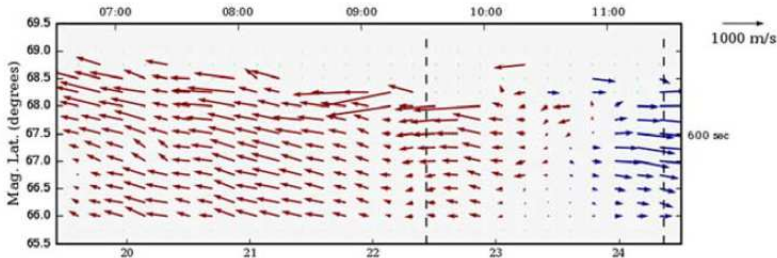
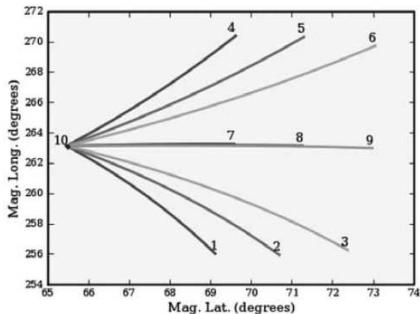
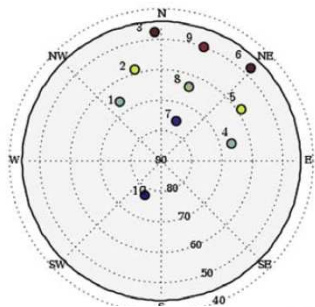
PFISR Overview

- 128 Panels
- Science operations began in 2007
- 65.13° geographic latitude (a few degrees south of Tromsø)
- 65.39° geomagnetic latitude (Auroral Region)
- Common Tx Frequencies 449.3, 449.6, 449.8 MHz
- Powered by grid
- Funded by the US National Science Foundation (NSF)
- Operated by SRI International

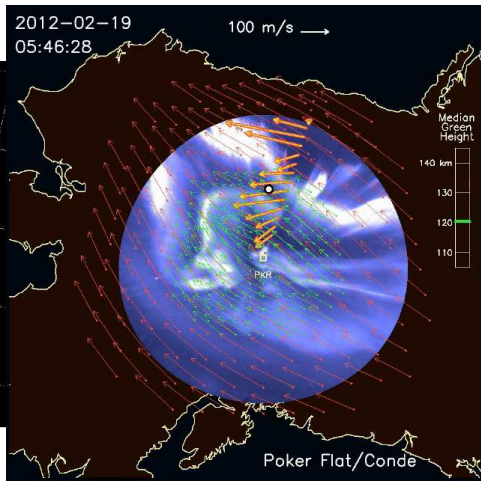
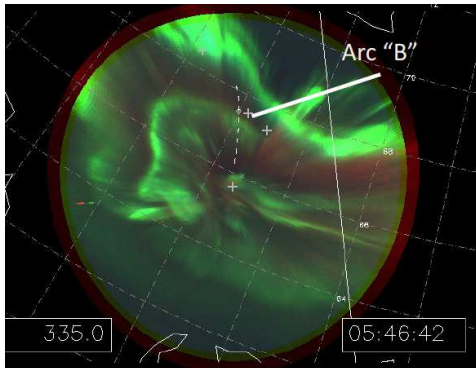
Imaging Auroral Structure [Semeter et al. (2009)]



Vector Electric Fields [Heinselman and Nicolls (2008)]



MICA Sounding Rocket Support

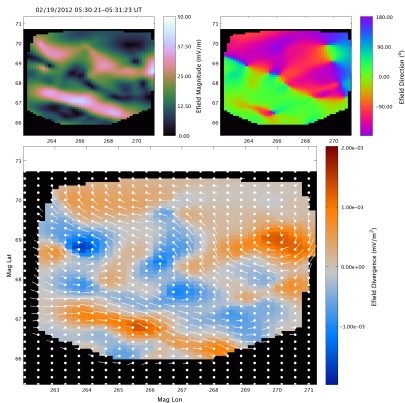


Lynch et al. (2015) *JGR*

Imaging Electrodynamics: 2-D \mathbf{E} -Estimation

Assumptions:

- \mathbf{E} maps along equipotential field lines
- $\nabla \times \mathbf{E} = 0 \Rightarrow \mathbf{E} = -\nabla\Phi$
- \mathbf{E} is “smooth” in that it minimizes a curvature measure G

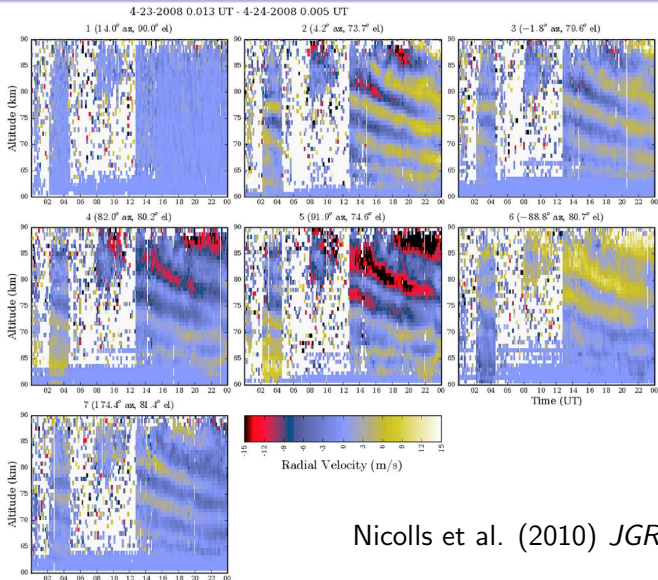


Constrained optimization problem using Lagrange multipliers:

$$\mathcal{L} = \|\Phi\|_G^2 + \lambda^\dagger (\tilde{\mathbf{v}}'_{los} - \mathbf{e} - M\Phi) + \Omega (\|\mathbf{e}\|_{C^{-1}}^2 - N + 1)$$

Nicolls et al. (2014) *Radio Sci.*

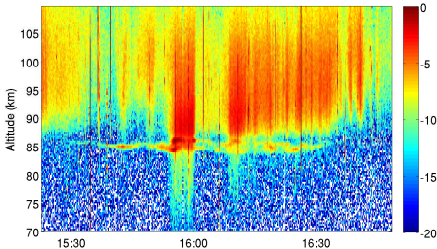
Mesospheric Winds and Gravity Waves



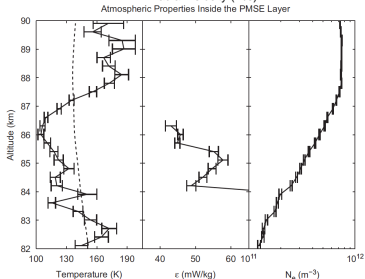
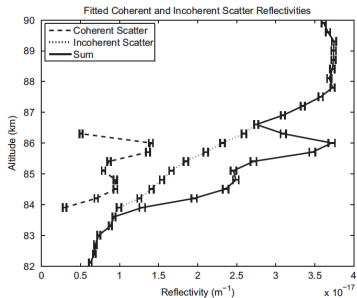
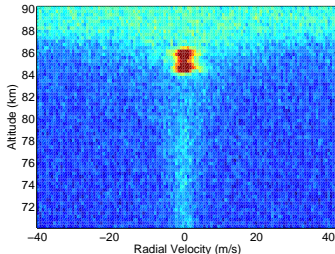
Nicolls et al. (2010) *JGR*

Polar Mesospheric Summer Echoes

SNR on 8/10/2008



Spectra



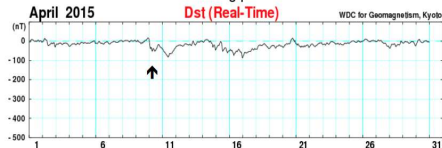
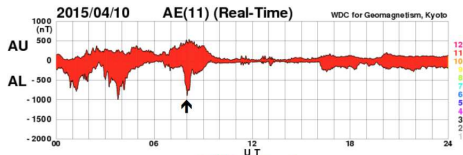
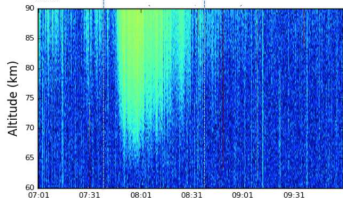
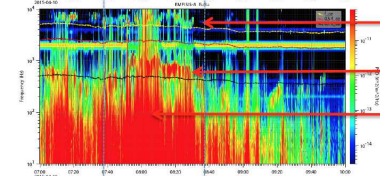
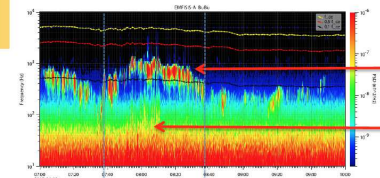
Varney et al. (2011) *JASTP*

Imaging PMSE

Movie

Nicolls et al. (2007) *GRL*

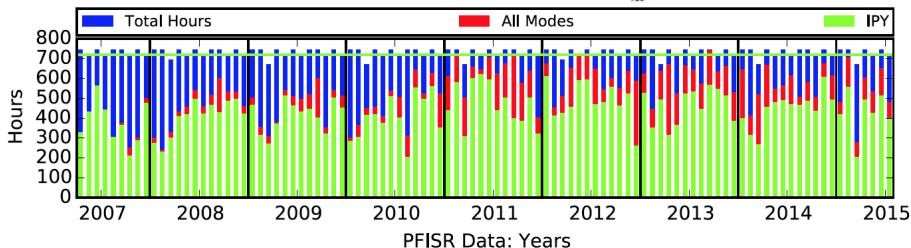
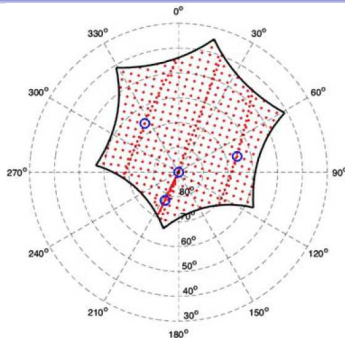
RBSP event #1 – 10 April 2015



[Created at 2015-05-08 15:05:11]

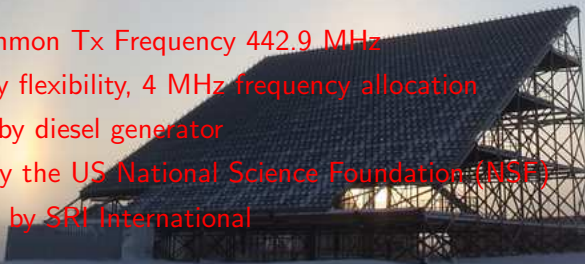
PFISR IPY Mode (Continuous Operations)

- 1% duty cycle
- 4 beams, including up-B
- Alternating code (E-region), Long Pulse (F-region)
- 5 min integration and fitting:
 - N_e, T_e, T_i, V_{LOS}
 - Vector electric field
 - E-region neutral wind

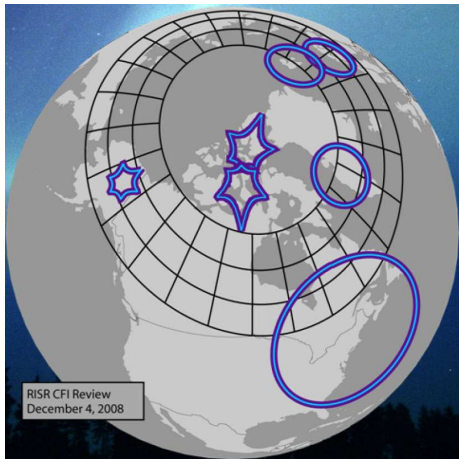
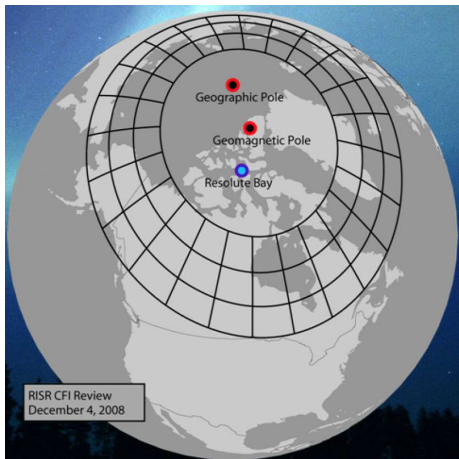


RISR-N Overview

- 121 Panels
- Science operations began in 2009
- 74.73° geographic latitude (not as high as Svalbard)
- 82.77° geomagnetic latitude (deep in the polar cap; highest of any ISR)
- Most common Tx Frequency 442.9 MHz
- Frequency flexibility, 4 MHz frequency allocation
- Powered by diesel generator
- Funded by the US National Science Foundation (NSF)
- Operated by SRI International

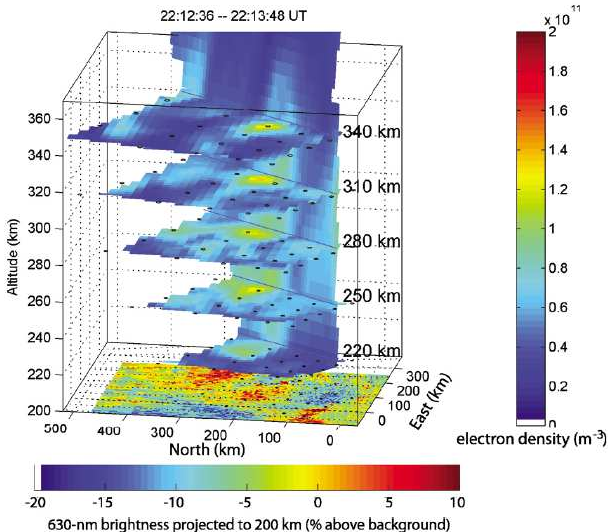


Unique Location of Resolute Bay



Figures courtesy Eric Donovan

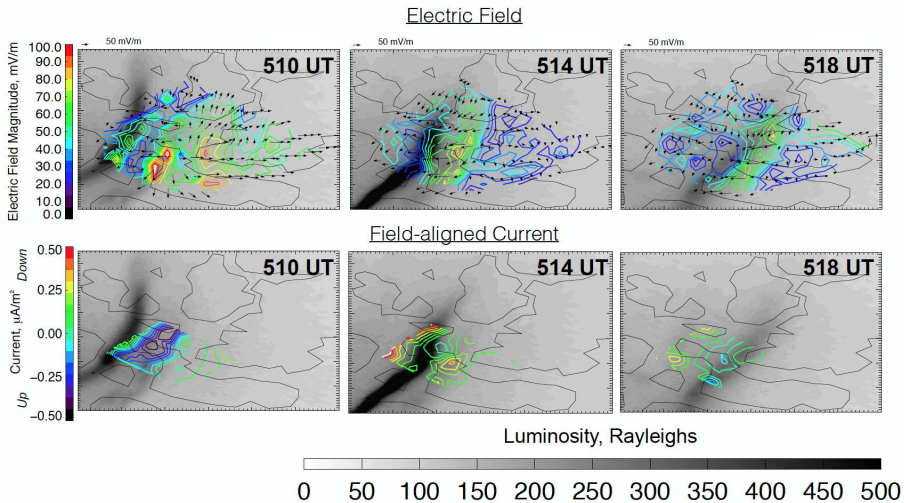
Polar Cap Patch Imaging



RISR-N
Volumetric
Images of
Polar Cap
Patches

Dahlgren et
al., 2012,
GRL.

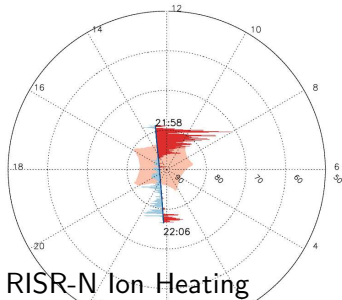
Electrodynamics of Polar Cap Arcs



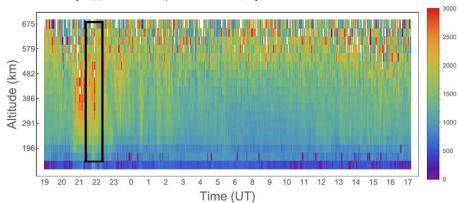
Perry et al. (2015) *JGR*

e-POP, SuperDARN, and RISR-N Observations

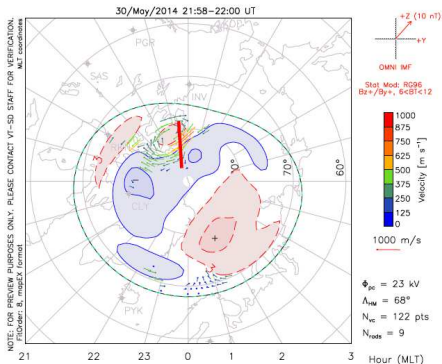
e-POP Upflow



RISR-N Ion Heating



SuperDARN Convection

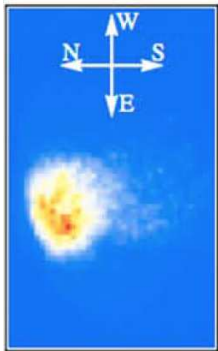


Shen et al. (2016) *JGR*

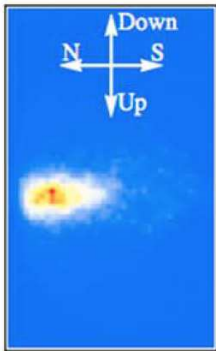
Intense Frictional Heating and Ion Anisotropy

SWARM Ion Distributions

Horizontal



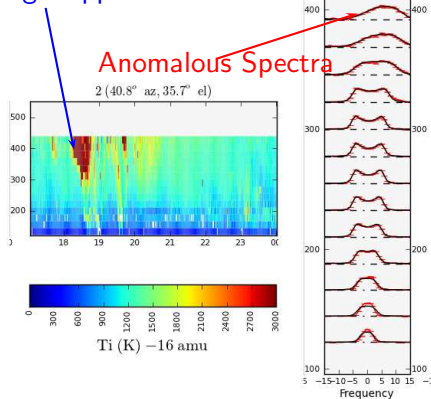
Vertical



Archer et al. (2015)

RISR-N Low-elevation Measurements

High Apparent T_i

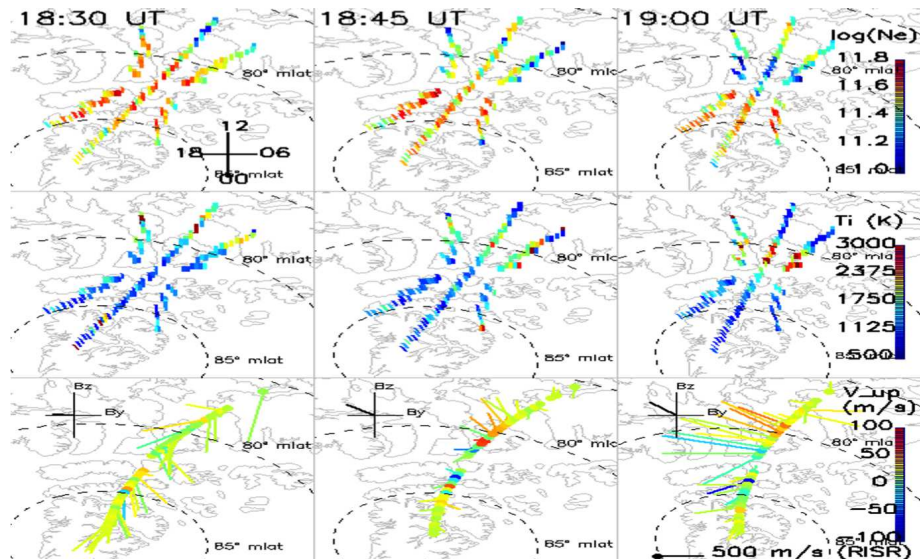


F-region anisotropy is common during intense frictional heating!

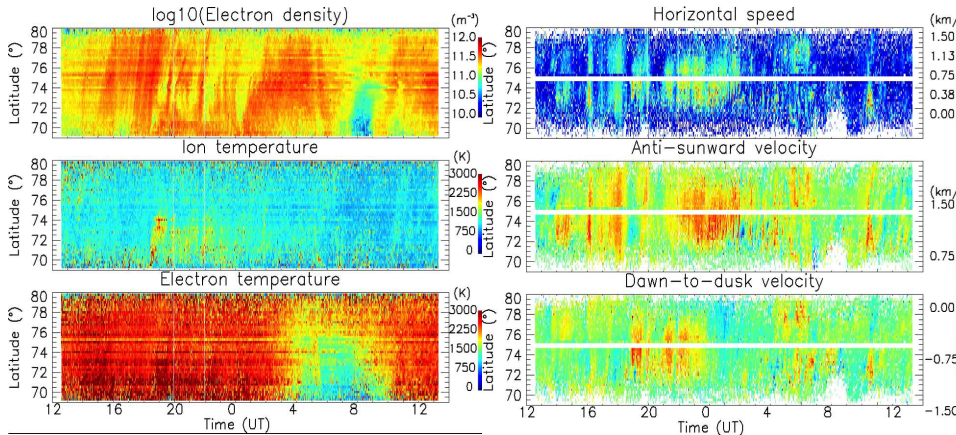
RISR-C Overview

- 121 panels (identical to RISR-N)
- Science operations began in 2016
- Pulses Synchronized to RISR-N. Both radars can operate simultaneously without interference
- Funded by the Canadian Foundation for Innovation (CFI)
- Operated by the University of Calgary

RISR-C and RISR-N Coordinated Observations



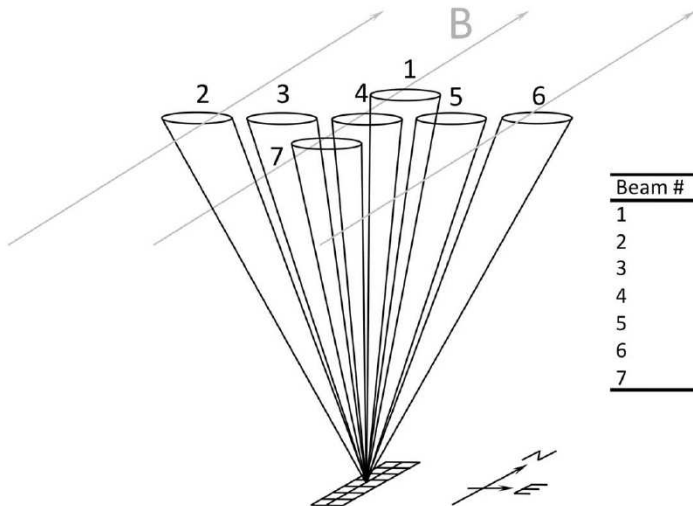
RISR “Keograms” (from Rob Gillies)



UMET-14 Overview

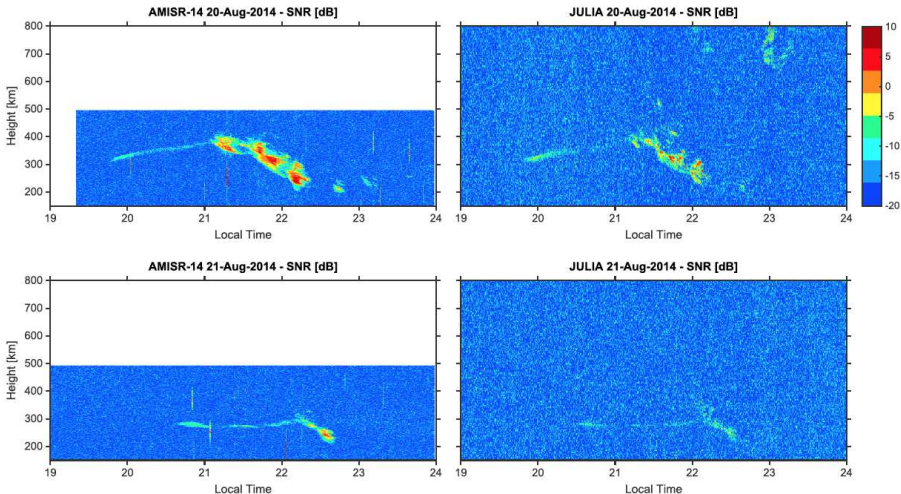
- 
- 14 panels
 - Deployed in August 2014
 - Long skinny array \Rightarrow beam narrow in N-S direction, wide in E-W direction \perp to \mathbf{B}
 - Operated by the Universidad Metropolitana in Puerto Rico (PI: Juan Arratia)

Beam Steering in Plane \perp to **B**



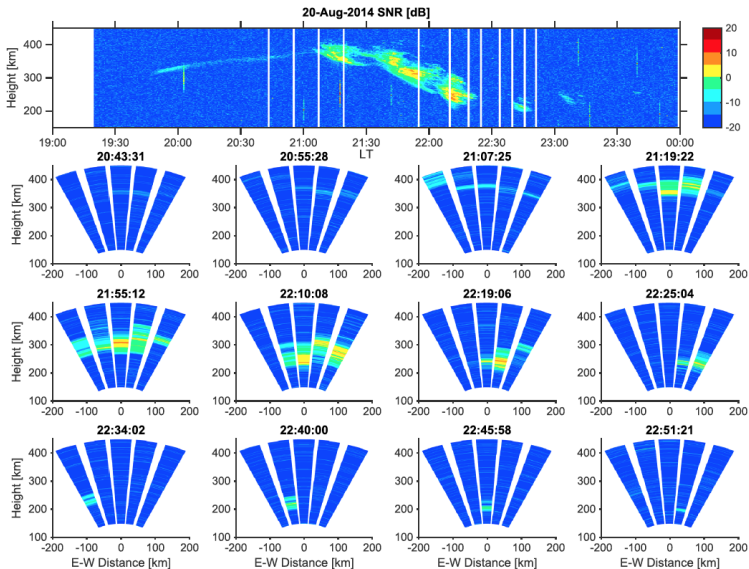
Beam #	Az	Elev
1	0°	85.2°
2	-90°	69.2°
3	-90°	80.4°
4	0°	90.0°
5	90°	80.4°
6	90°	69.2°
7	180°	85.2°

Spread-F Scatter at 33 cm vs 3 m Scales

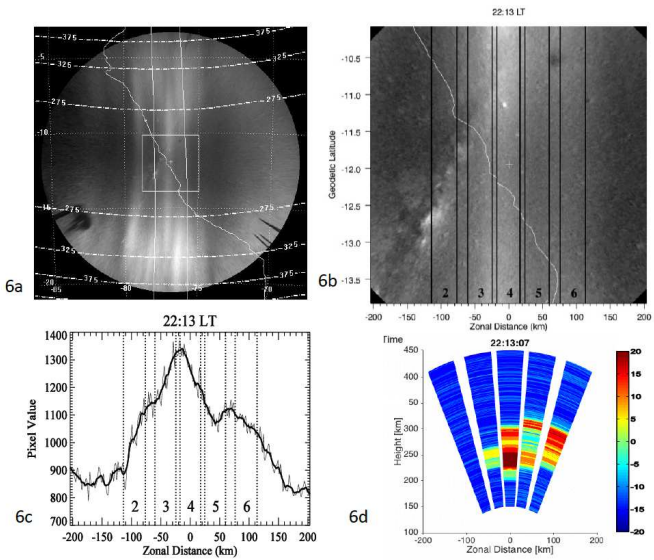


Rodrigues et al. (2015) *GRL*

Resolving Space-Time Ambiguity in Spread-F

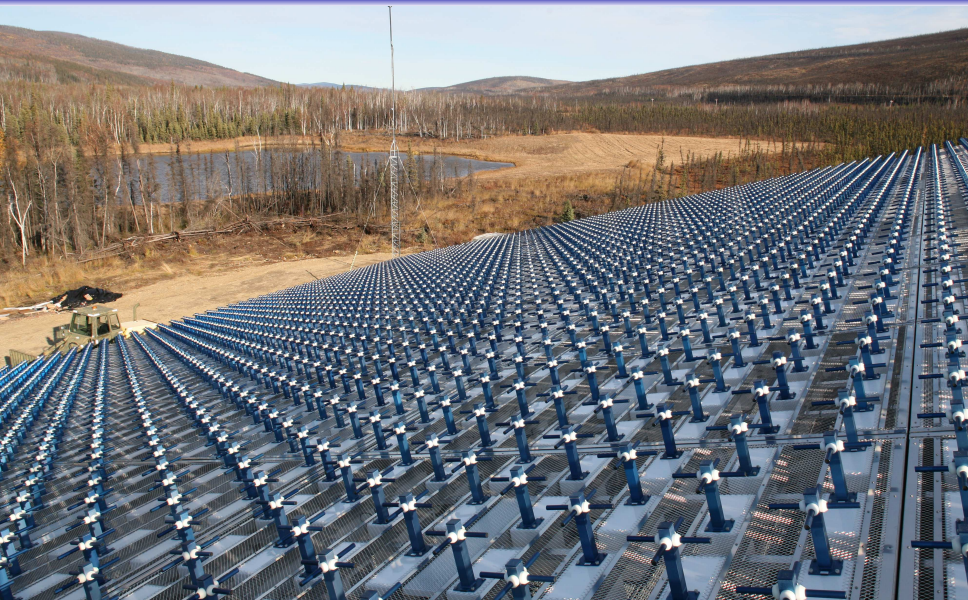


Location of Scatter within Airglow Depletions



Hickey et al. (2015) *JGR*

Future of AMISR Science?



Future of AMISR Science?

