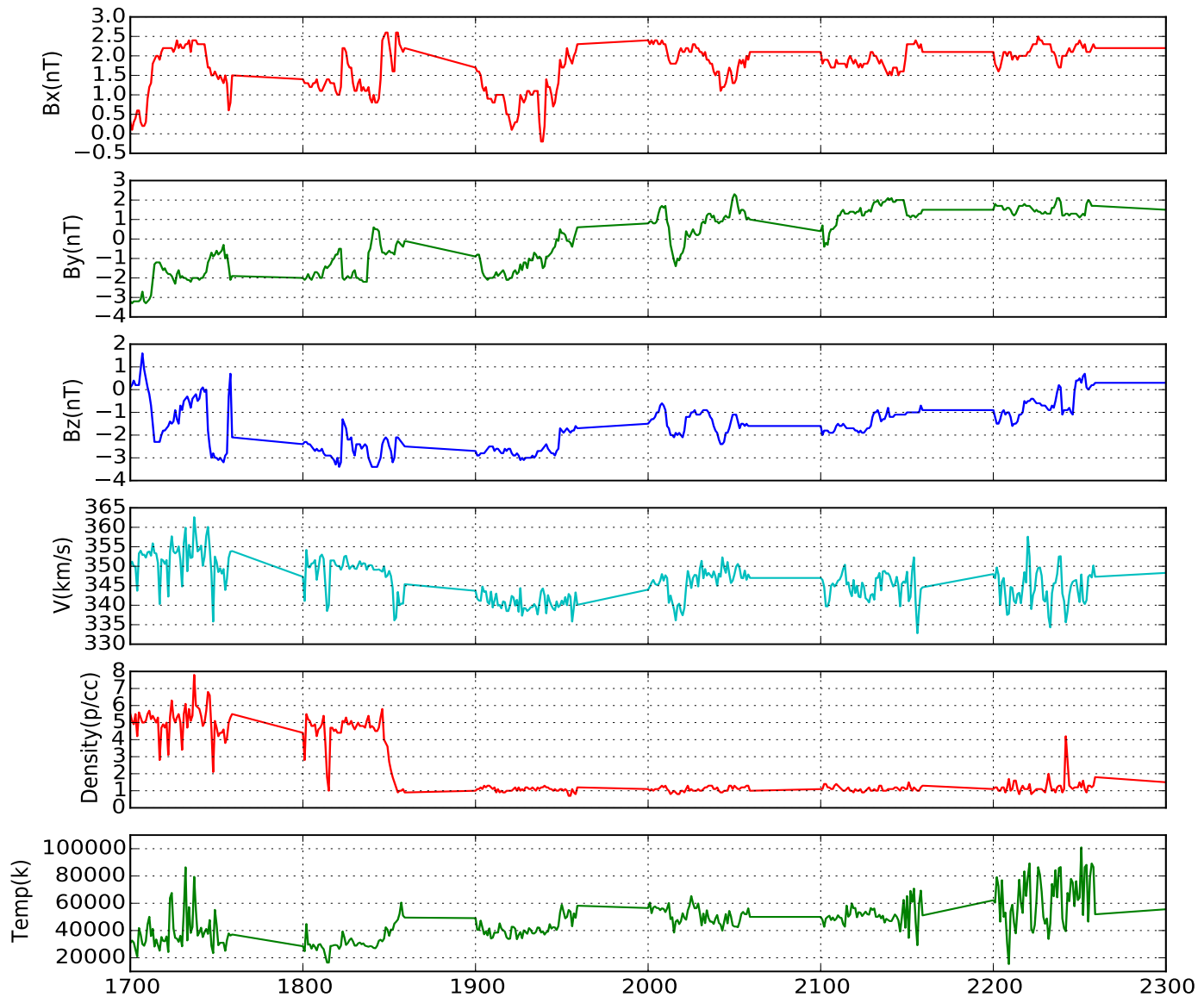




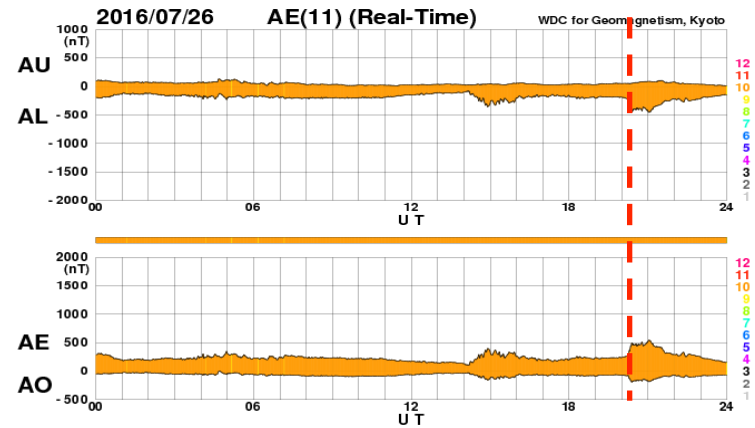
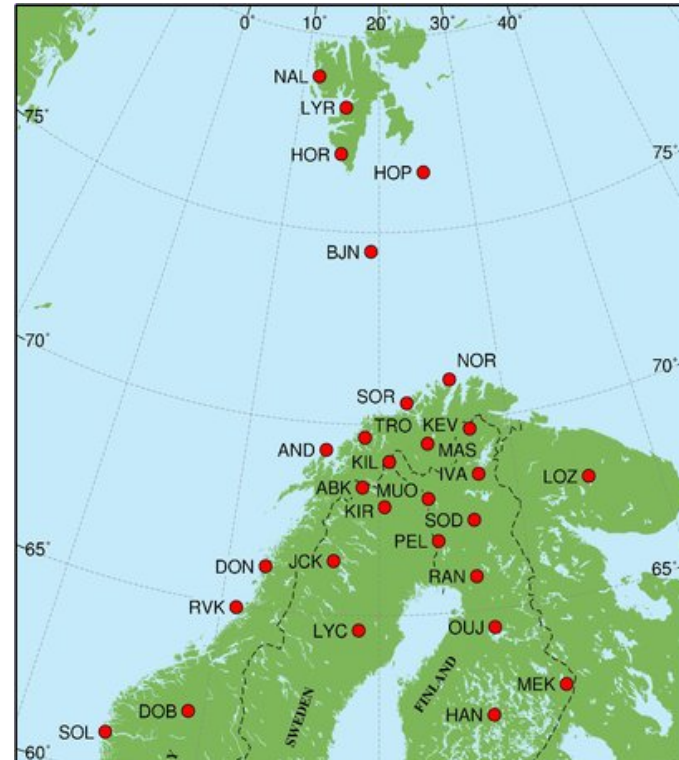
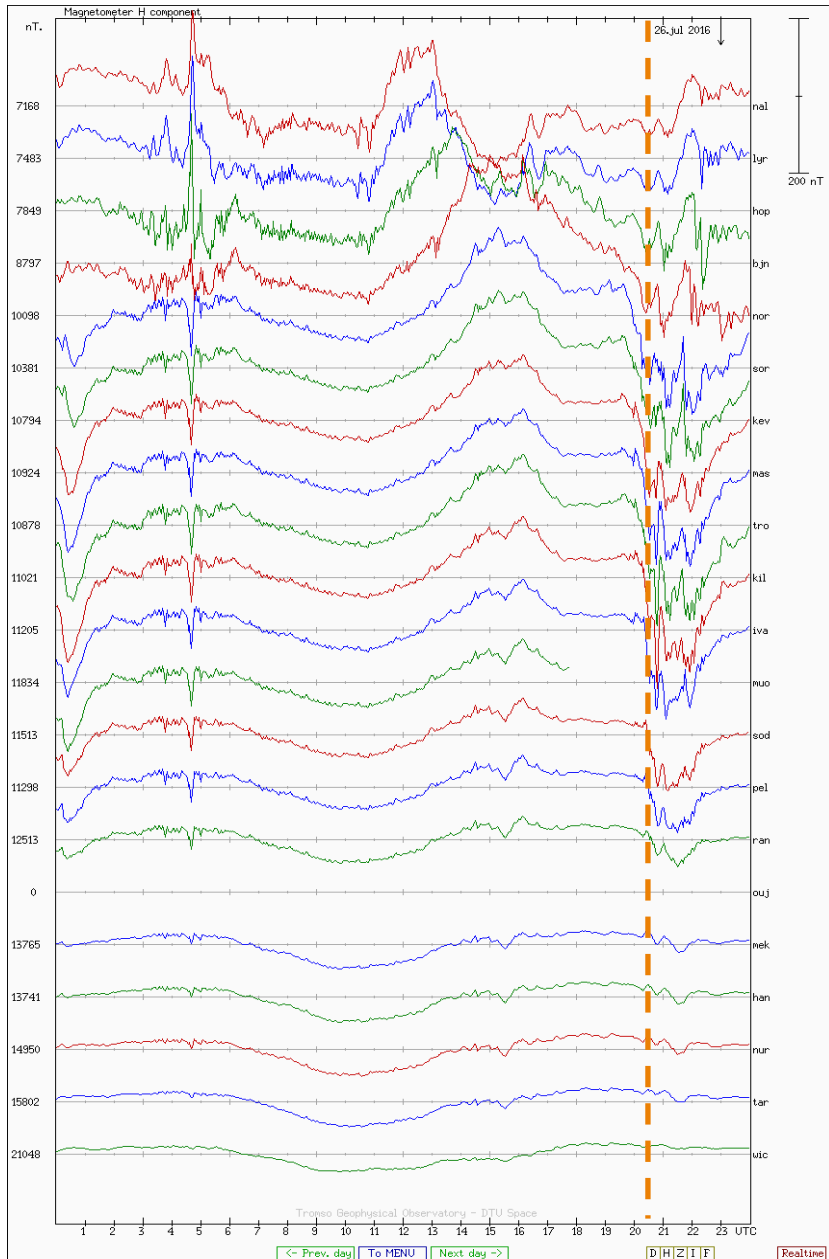
Study of Ion Upflow during Substorm using ESR 32 m (Folke)

Bashi Ferdousi, Biqiang Zhao, Derek McKay, Greg Myer, Timothy David
Group1, RADAR Summer School, Sodankyla, Finland
July 29th, 2016

Solar Wind Conditions from ACE on July 26th, 2016



H Component of Ground Magnetometer and AE Index



[Created at 2016-07-27 13:20UT]

Background to Ion Outflow

Ion outflow originates from dayside cusp and night side aurora region

[Yu and Ridley, 2013]

Auroral Bulk Upflow

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graph TD; A[Auroral Bulk Upflow] --> B[Type-I]; A --> C[Type-II];
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Type-I

- High convection electric field
- Strong Joule heating
- No auroral precipitation

Type-II

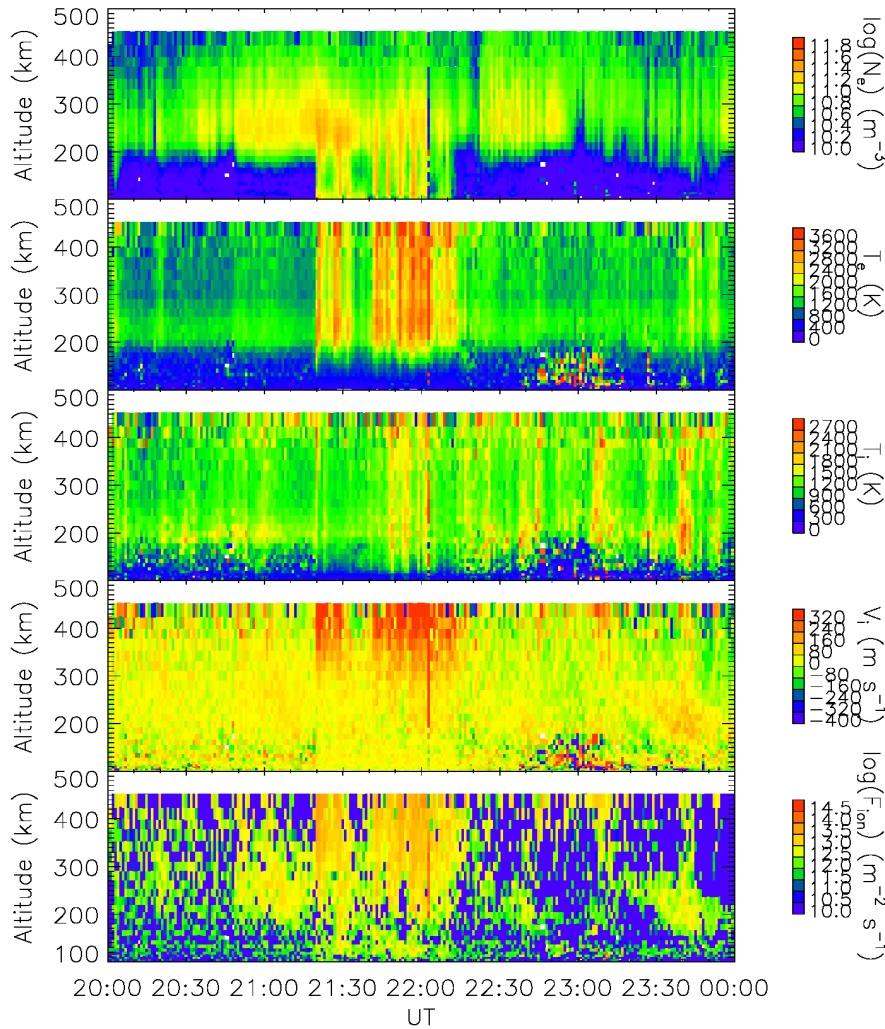
- $FAV > 500 \text{ ms}^{-1}$
- $F_{\text{ion}} > 10^{13} \text{ m}^{-2}\text{s}^{-1}$
- Strong T_e in the F-region
- No evidence of significant ion heating

Wahlund and Opgenoorth (1989) ; Wahlund et al. (1992)

Result: Comparison between 15/03/07 and 26/07/16 event

ESR Parameter Plot

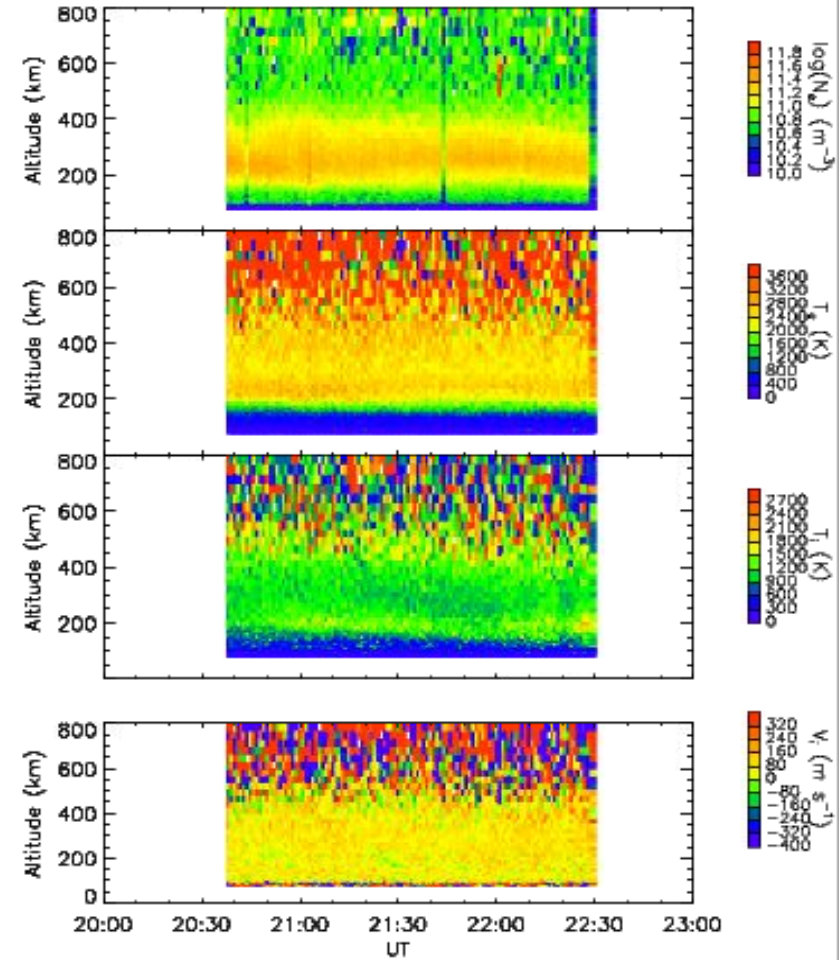
20070315



Our Observation

ESR Parameter Plot

20160726



Conclusions

NO ION UPFLOW OBSERVED

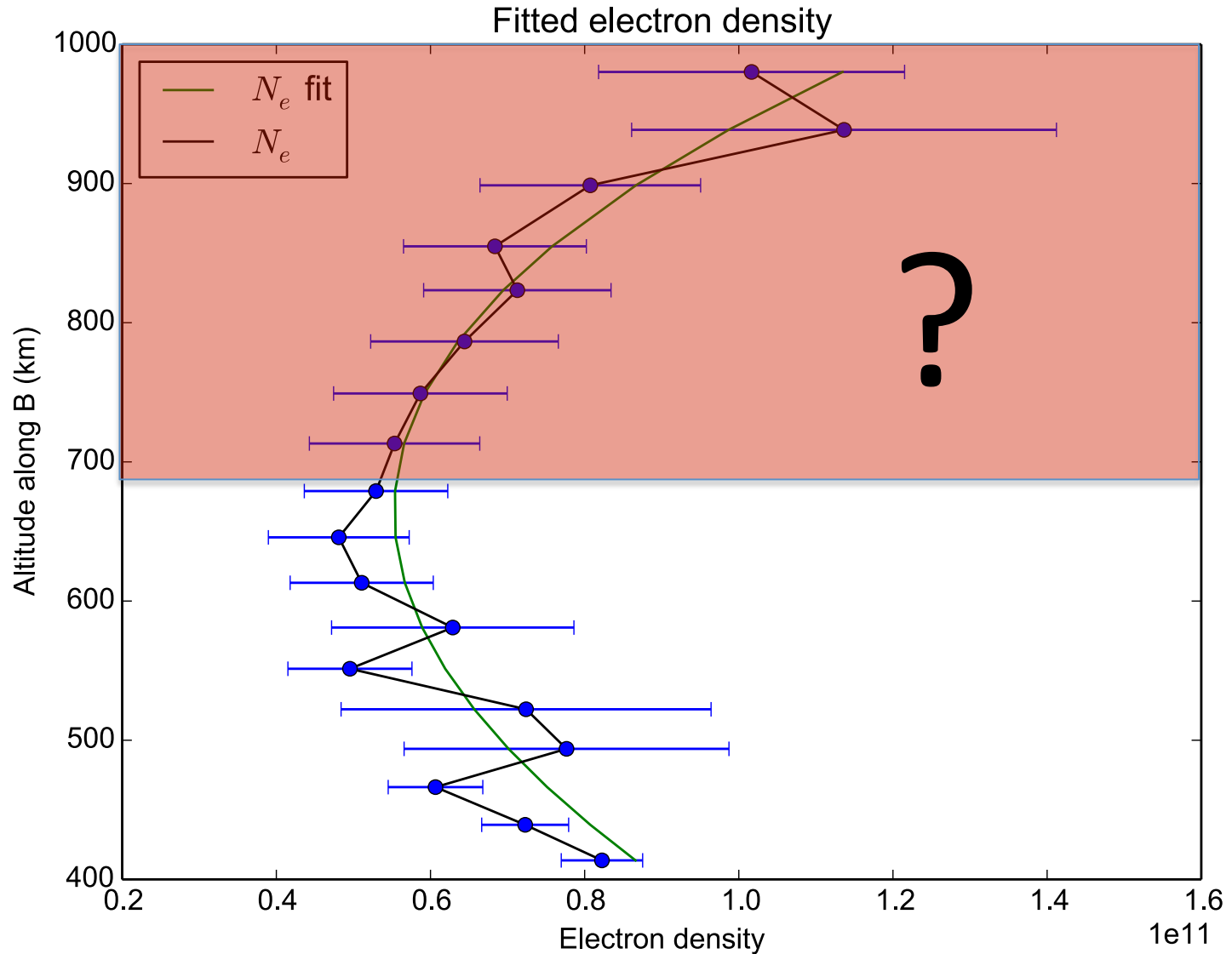
Thanks
Questions!

Problem Derivation

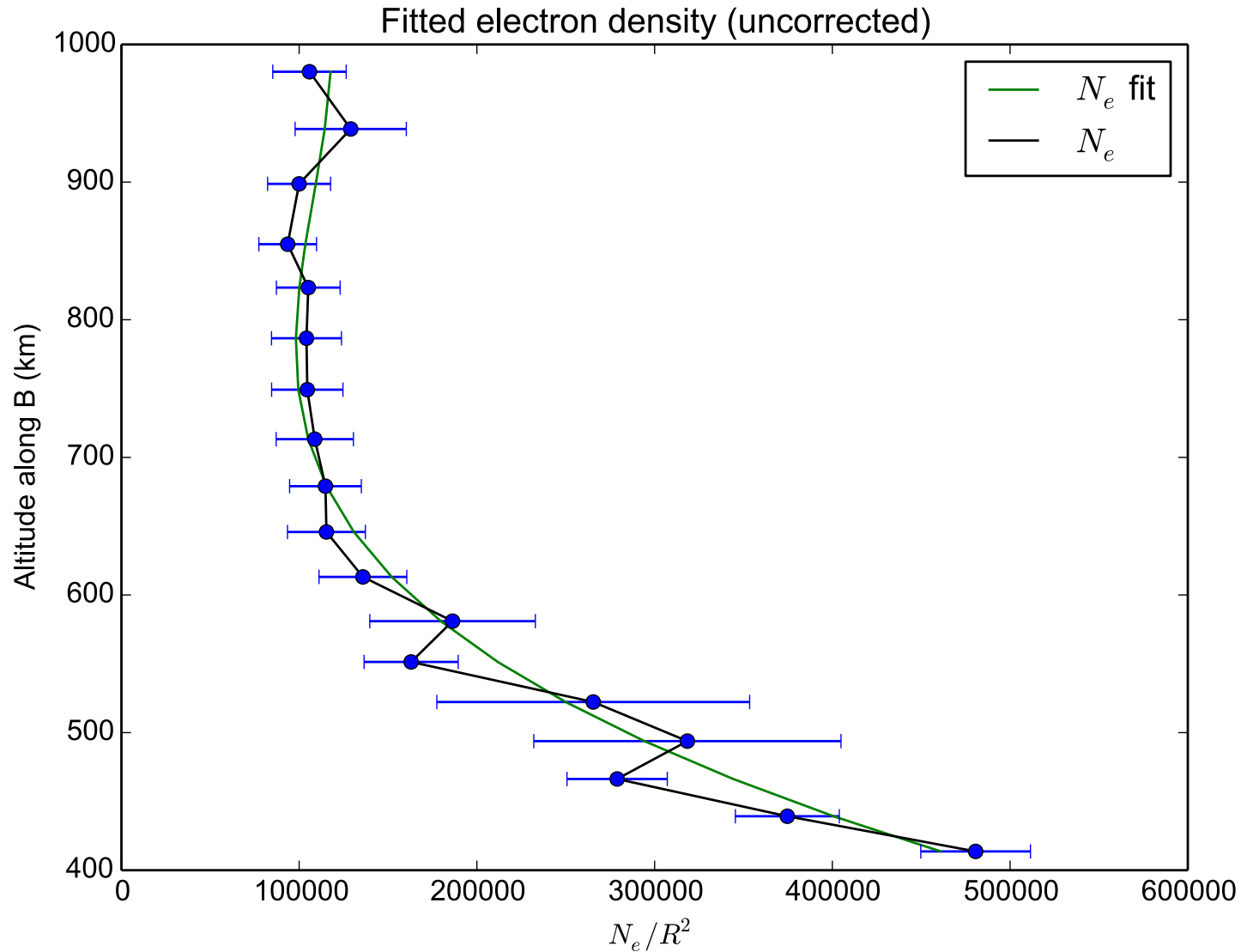
- Momentum Equation;
$$\rho \frac{d\vec{V}}{dt} = \rho \vec{g} - \nabla P + \rho_q \vec{E} + \vec{j} \times \vec{B}$$
- Ambipolar electric field;
$$E_{\parallel} = -\frac{1}{eN_e} \nabla_{\parallel} (N_e K_B T_e)$$
- State of Equilibrium;
$$0 = -m_i N_i g - \nabla_{\parallel} (N_i K_B T_i) + e N_i E_{\parallel}$$
$$= -m_i N_e g - \nabla_{\parallel} (N_e K_B T_i) - \nabla_{\parallel} (N_e K_B T_e)$$

$$\frac{\partial}{\partial s} (N_e K_B (T_i + T_e)) = -m_i g N_e$$

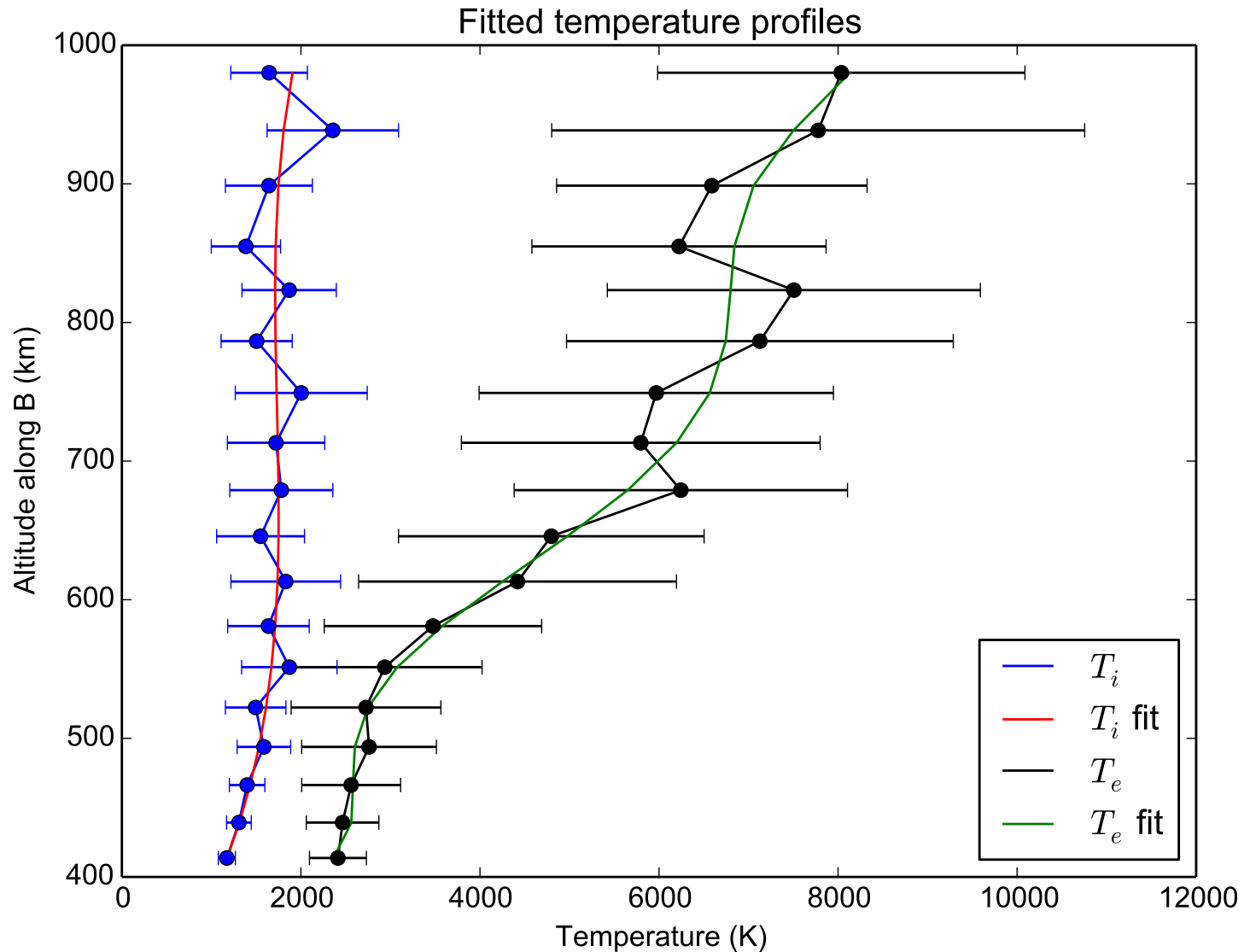
Polynomial Fitting for $N_e(s)$



Removal of R^2 Dependency

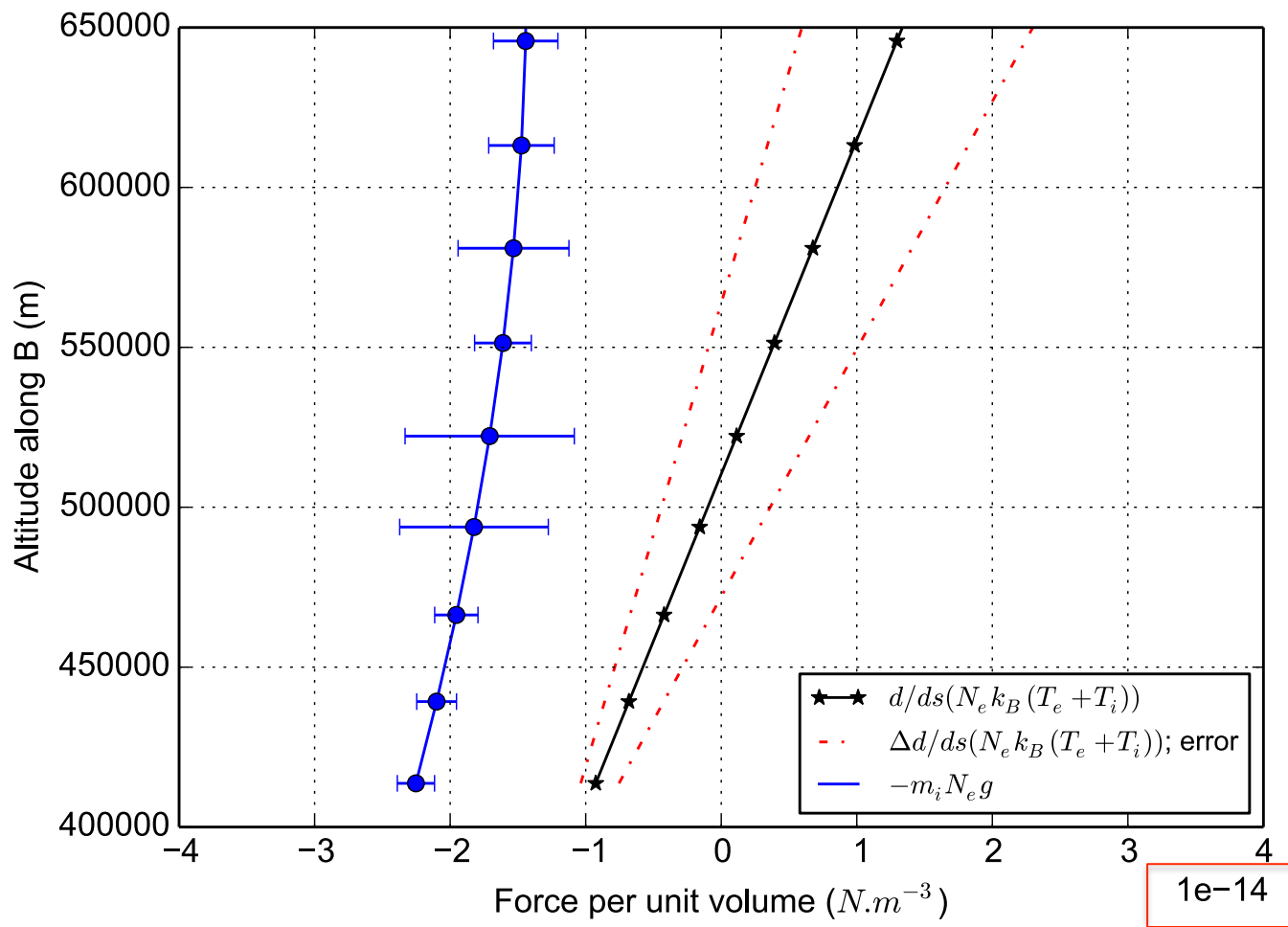


Polynomial Fitting of Temperatures

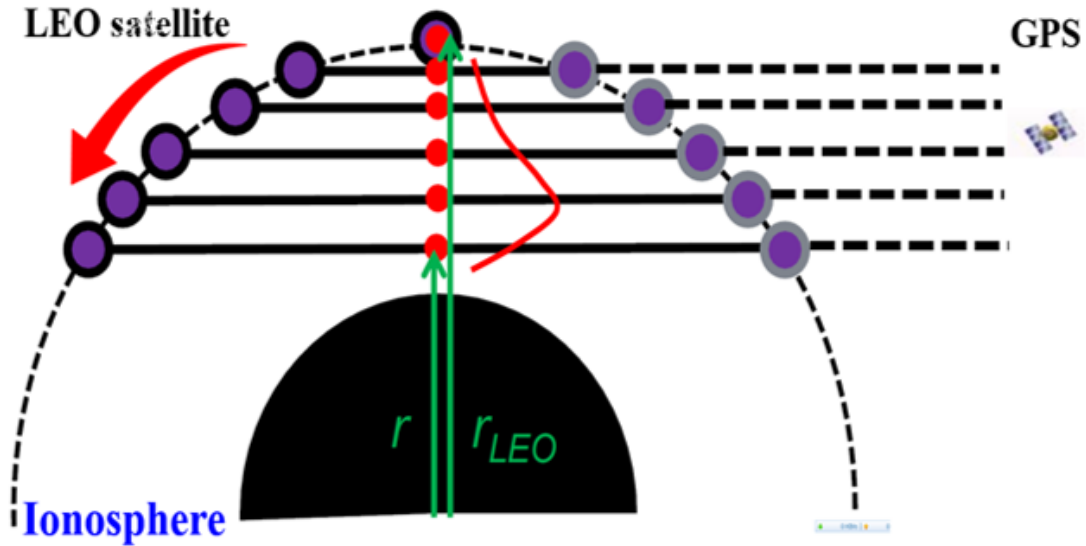
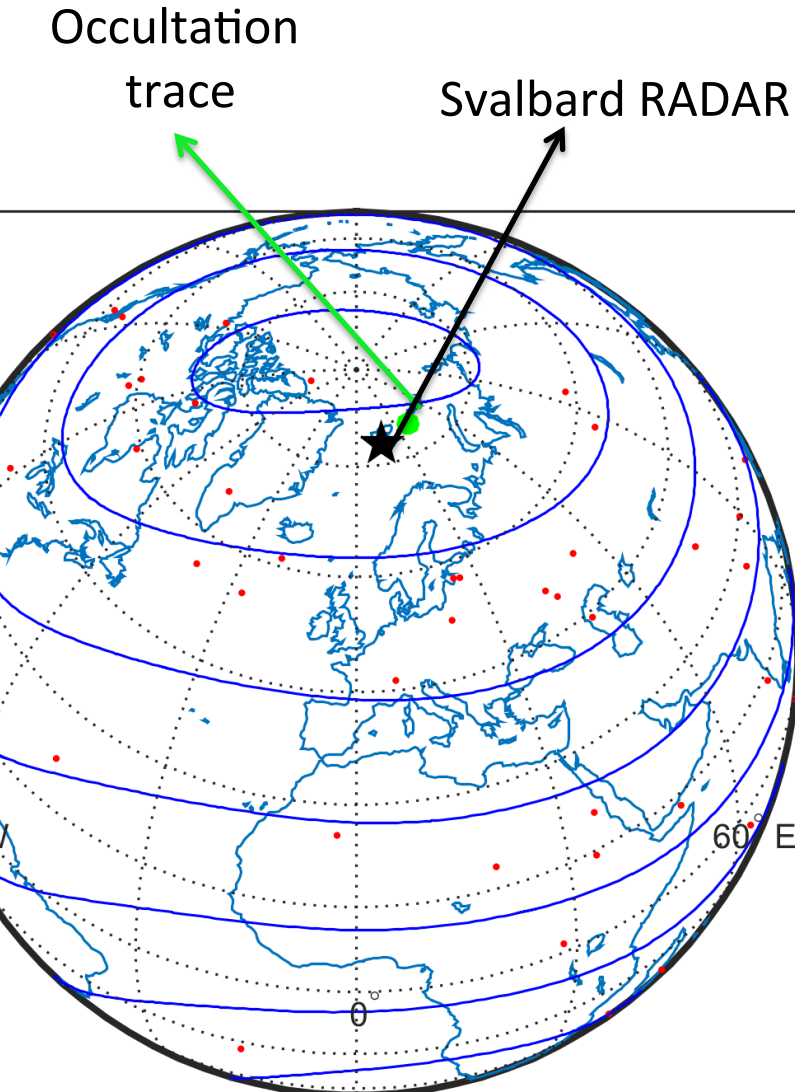


Comparison of Force Density Balance

$$\frac{\partial}{\partial s} (N_e K_B (T_i + T_e)) = -m_i g N_e$$



“COSMIC Background”

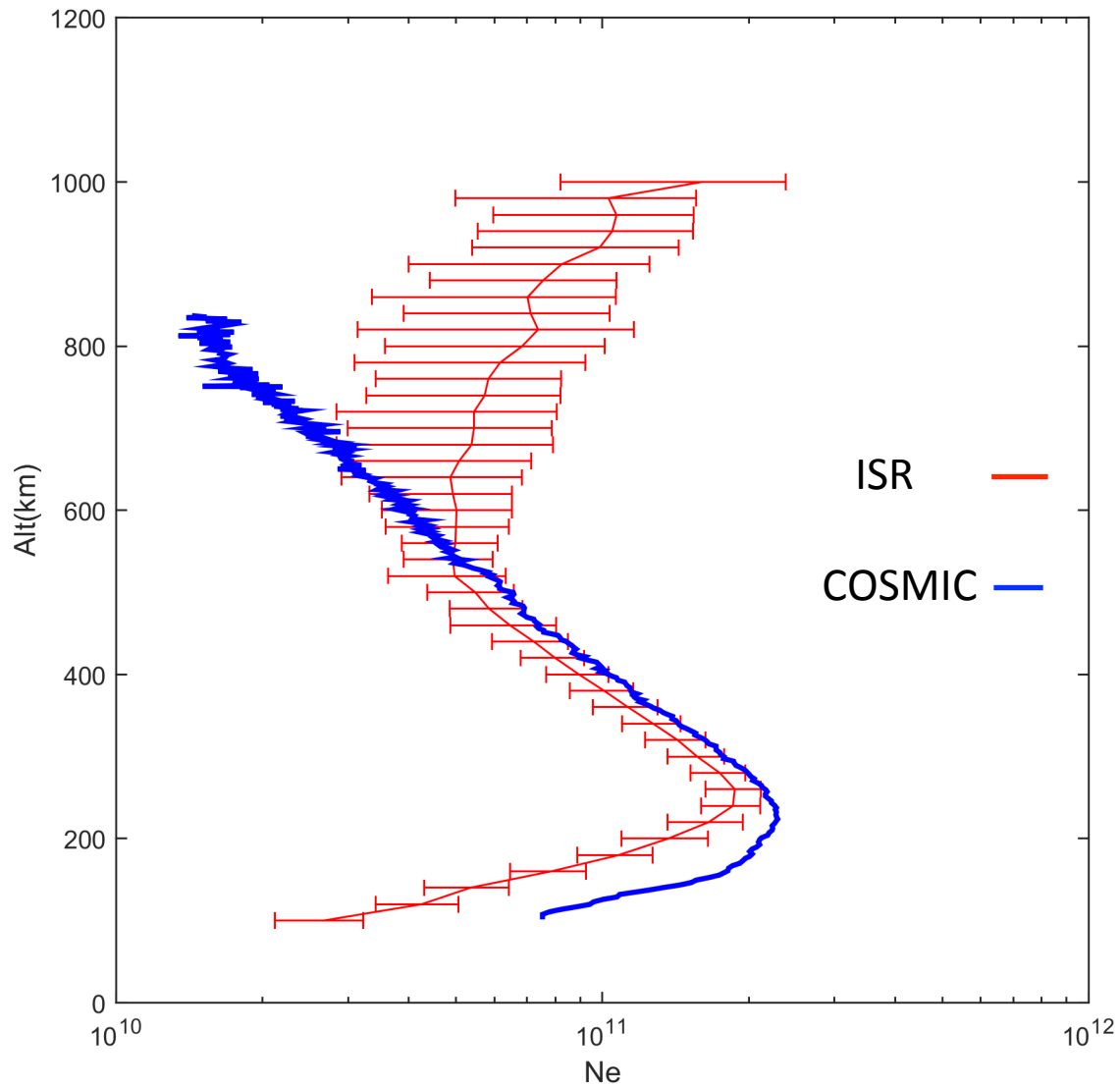


$$TEC_{OB}(r_P) = \left(\int_{r_P}^{r_{GPS}} + \int_{r_P}^{r_{LEO}} \right) \frac{r n_e(r)}{\sqrt{r^2 - r_P^2}} dr$$

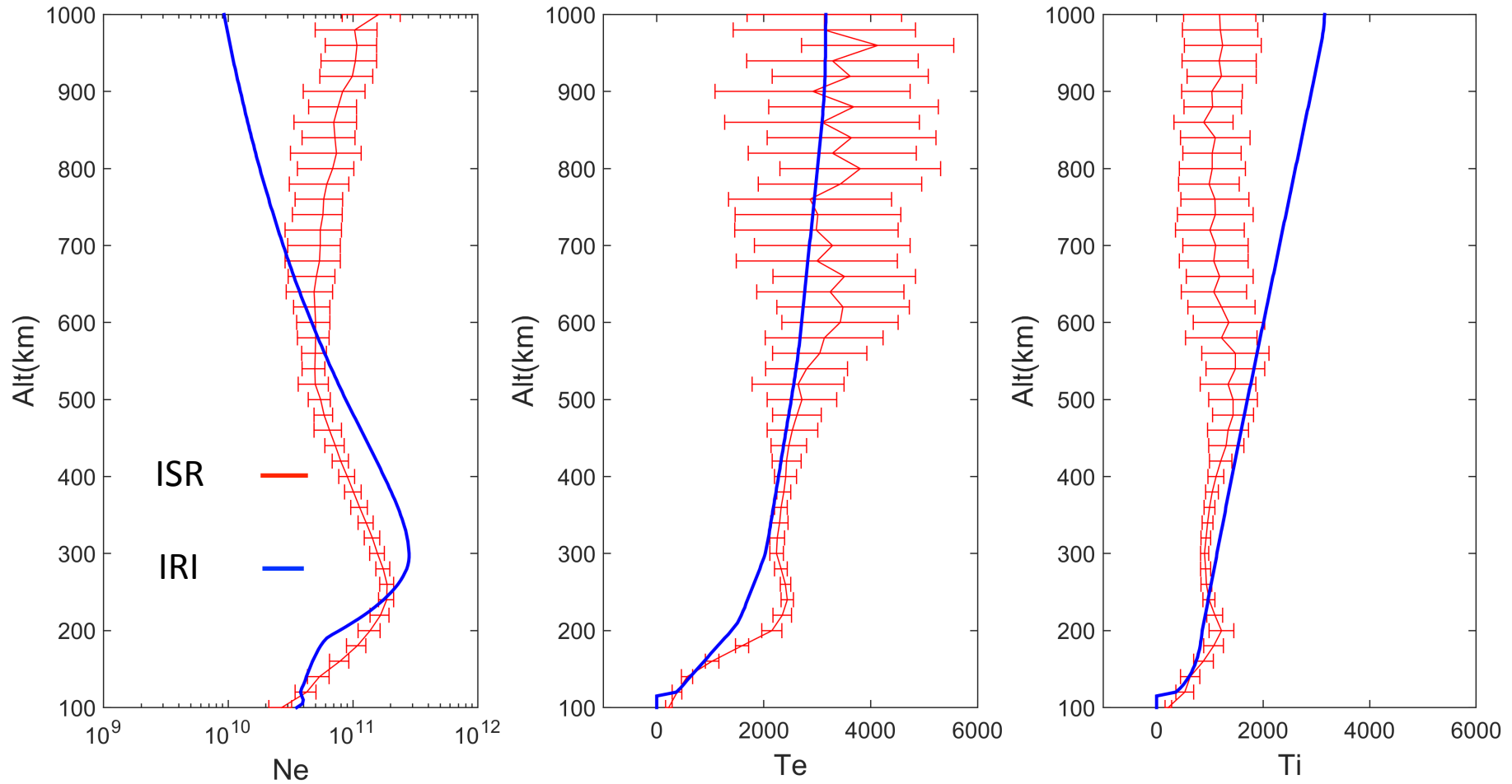


$$N_e(r) = -\frac{1}{\pi} \int_r^{r_{LEO}} \frac{dTEC_0(p)/dp}{\sqrt{p^2 - r^2}} dp$$

Comparison between ISR and COSMIC Measurements



Comparison of ISR Measurement and IRI Model



Conclusions

- ESR 32 m (Folke) has been used to study ion upflow.
- No ion upflow has been observed during on 26/07/16 between 20:30 to 22:30 UT.
- This allowed us to examine quiet ionosphere equilibrium : hydrostatic pressure and ambipolar electric field.
- Data was used to verify COSMIC measurement and IRI model.
- Given more time, another data set containing ion upflow would be analyzed to show comparisons.