

# Introduction to Incoherent Scatter

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*MIT Haystack Observatory*

With credit and thanks to **Anja Strømme**,  
Craig Heinselmann, Phil Erickson, Bill  
Rideout, Josh Semeter, Juha Vierinen

And my advisor: William E. Gordon

# Outline

Distribution of ISRs across the world

Introduction to ISR principles and Incoherent Scatter Spectrum

# Global Network of Incoherent Scatter Radars



***Can Measure Physical Properties of the Space Environment  
as a function of altitude:***

**electron density, electron temperature, ion temperature, plasma velocity**

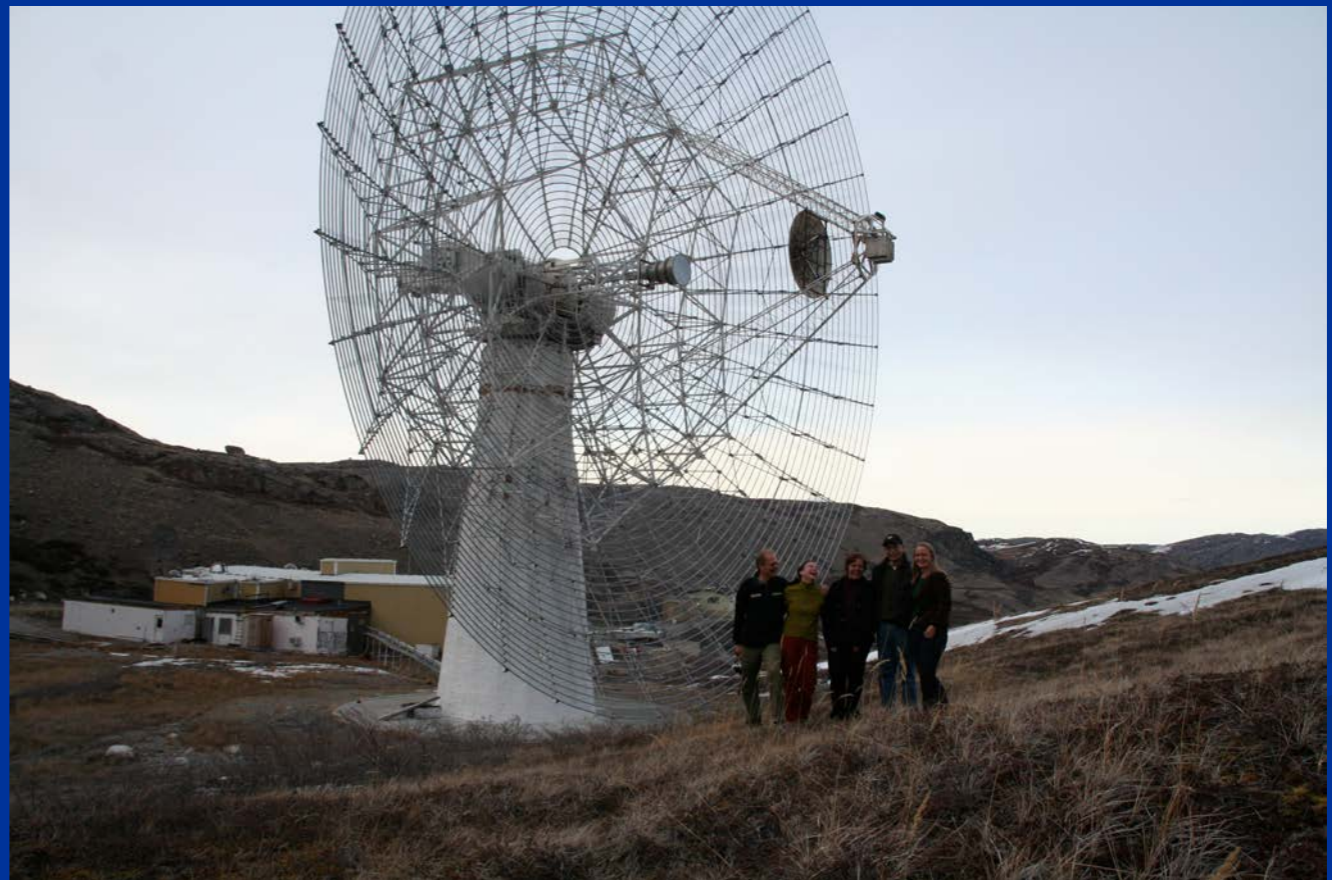
***Can Infer:***

**electric field strength, conductivity, current, neutral air temperature, wind**

**speed**



## High-Latitude Incoherent Scatter Radars



# PFISR (Poker Flat Incoherent Scatter Radar) and RISR-N (Resolute Bay Incoherent Scatter Radar)



# Mid-Latitude Incoherent Scatter Radars



# Low-Latitude Incoherent Scatter Radars



# IS Radar Remote Sensing Capabilities

Parameters sensed:

## Basic

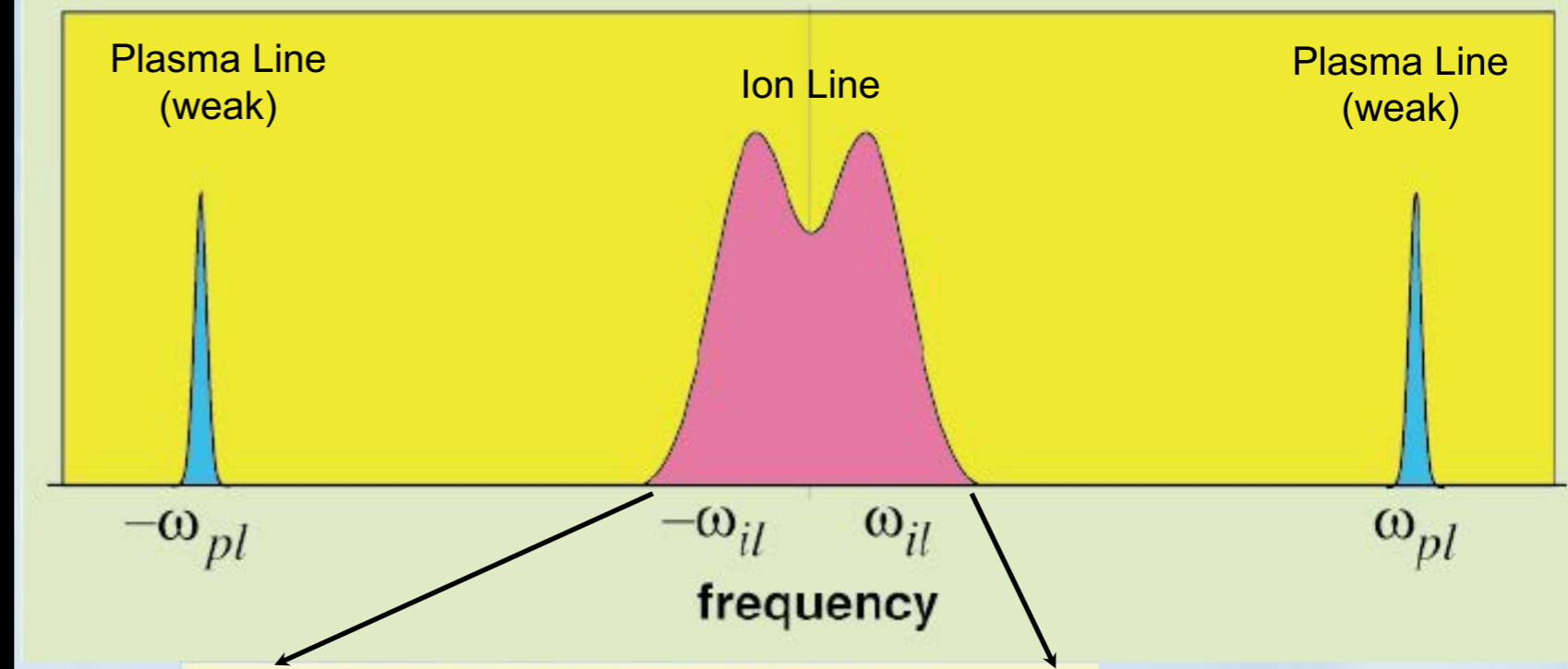
- Electron density
- Electron temperature
- Ion temperature
- Ion composition
- LOS Velocity

## Derived

- Neutral winds
- Neutral temperature
- Vector velocity

## More limited

- Ion-neutral collisions (E region)
- Background mag field (equator)
- Regularized binned/gridded data
- Etc....

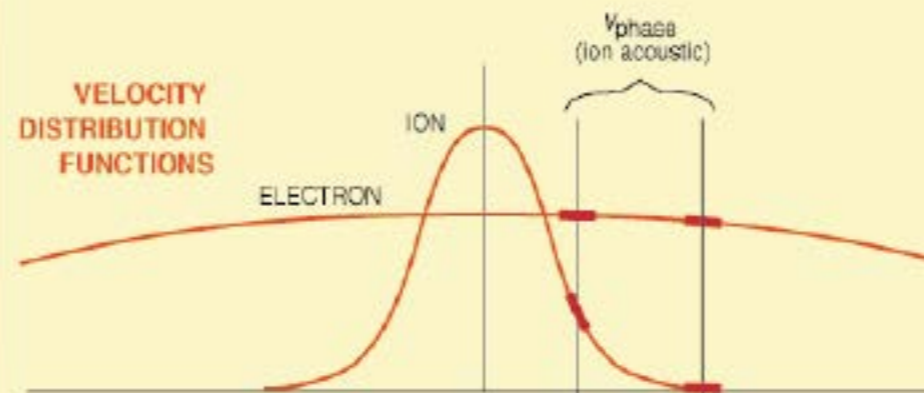


## THE EFFECT OF LANDAU DAMPING ON THE INCOHERENT SCATTER ION LINE SPECTRUM

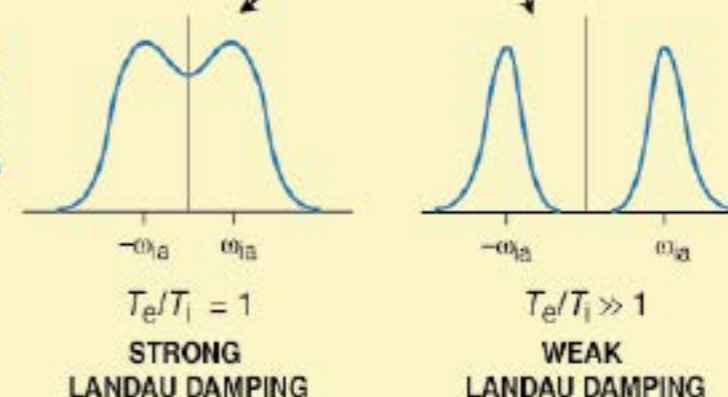
ION-ACOUSTIC DISPERSION EQUATION

$$\omega_{ia} = k v_{\text{phase}} = k \left( \frac{T_e + 3T_i}{m_i} \right)^{1/2}$$

VELOCITY DISTRIBUTION FUNCTIONS



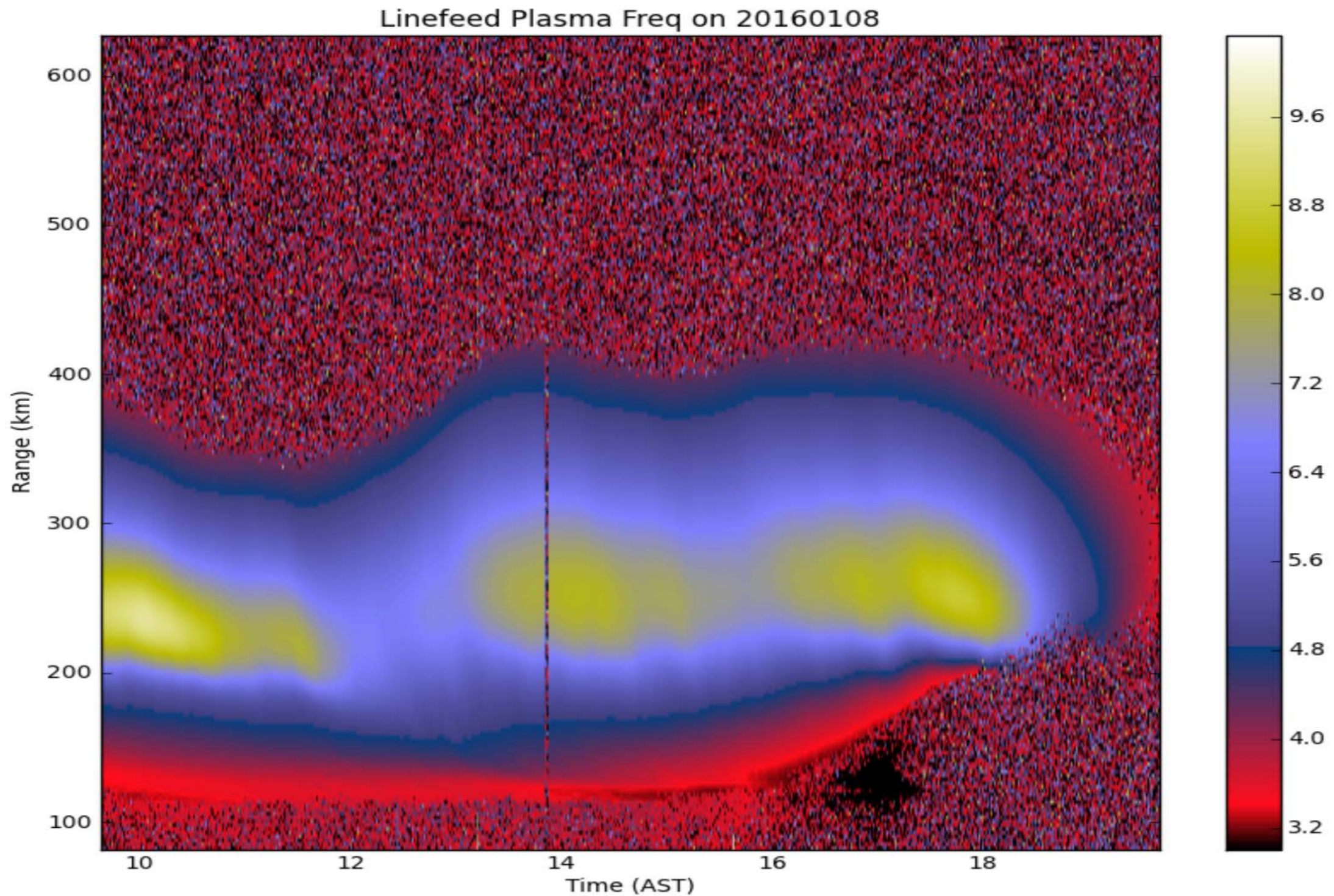
INCOHERENT SCATTER ION LINE SPECTRA



Example:  
"Ion Line"  
Sensitivity to  
Plasma  
Temperatures



# Arecibo's Summary Plasma Line Data



# Incoherent Scatter Radar

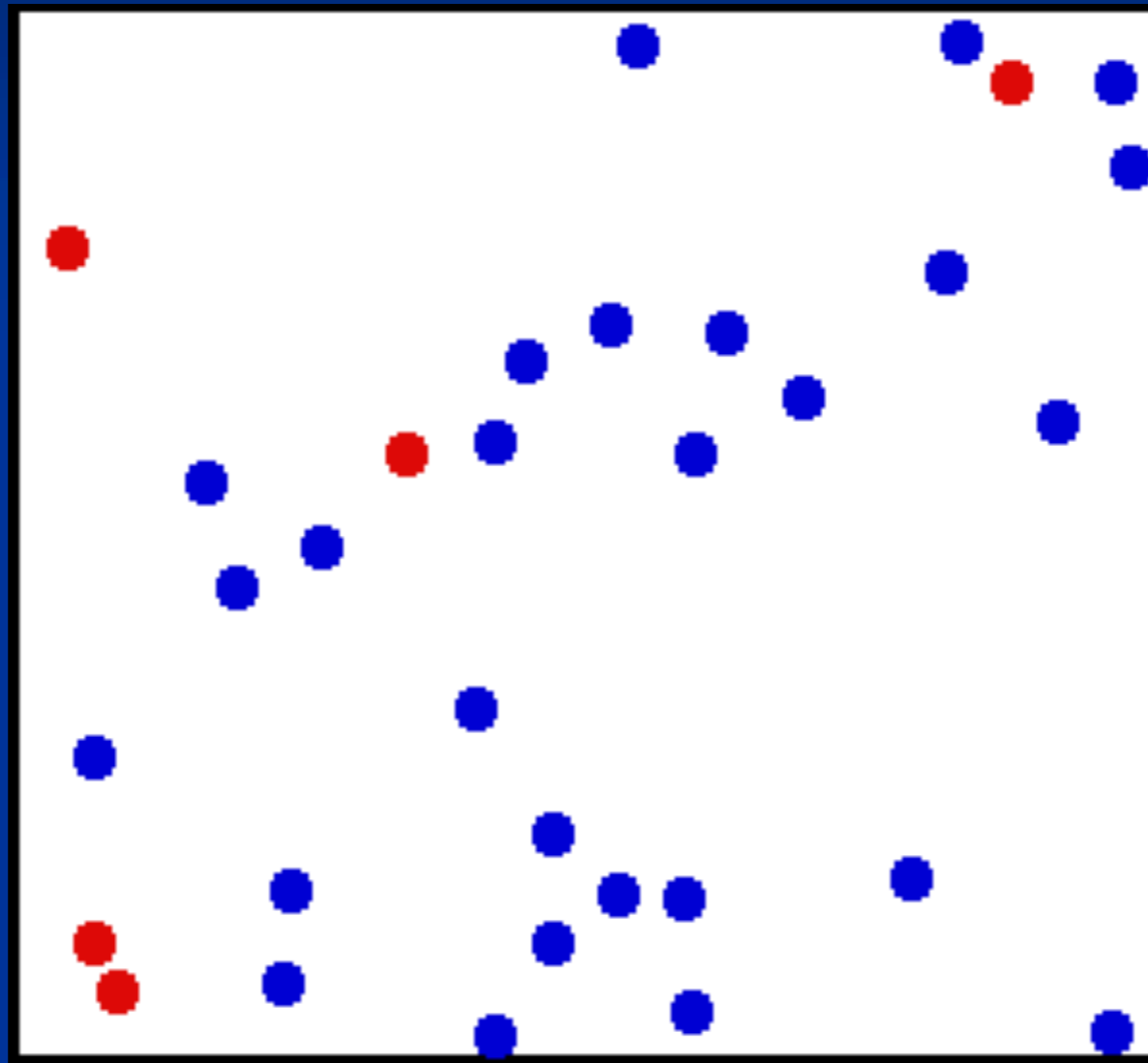
- Radar
- Scatter
- Incoherent

# Radar

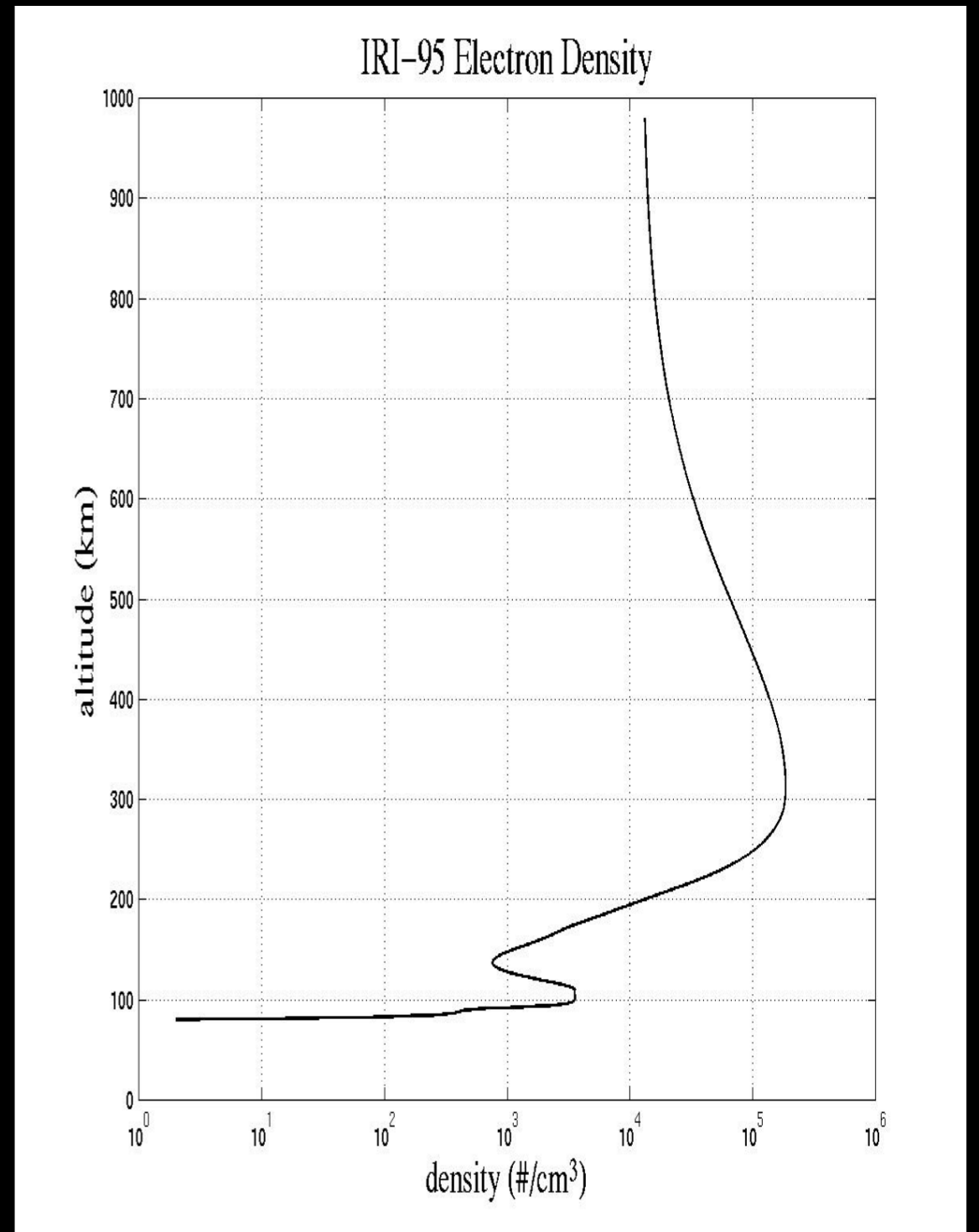
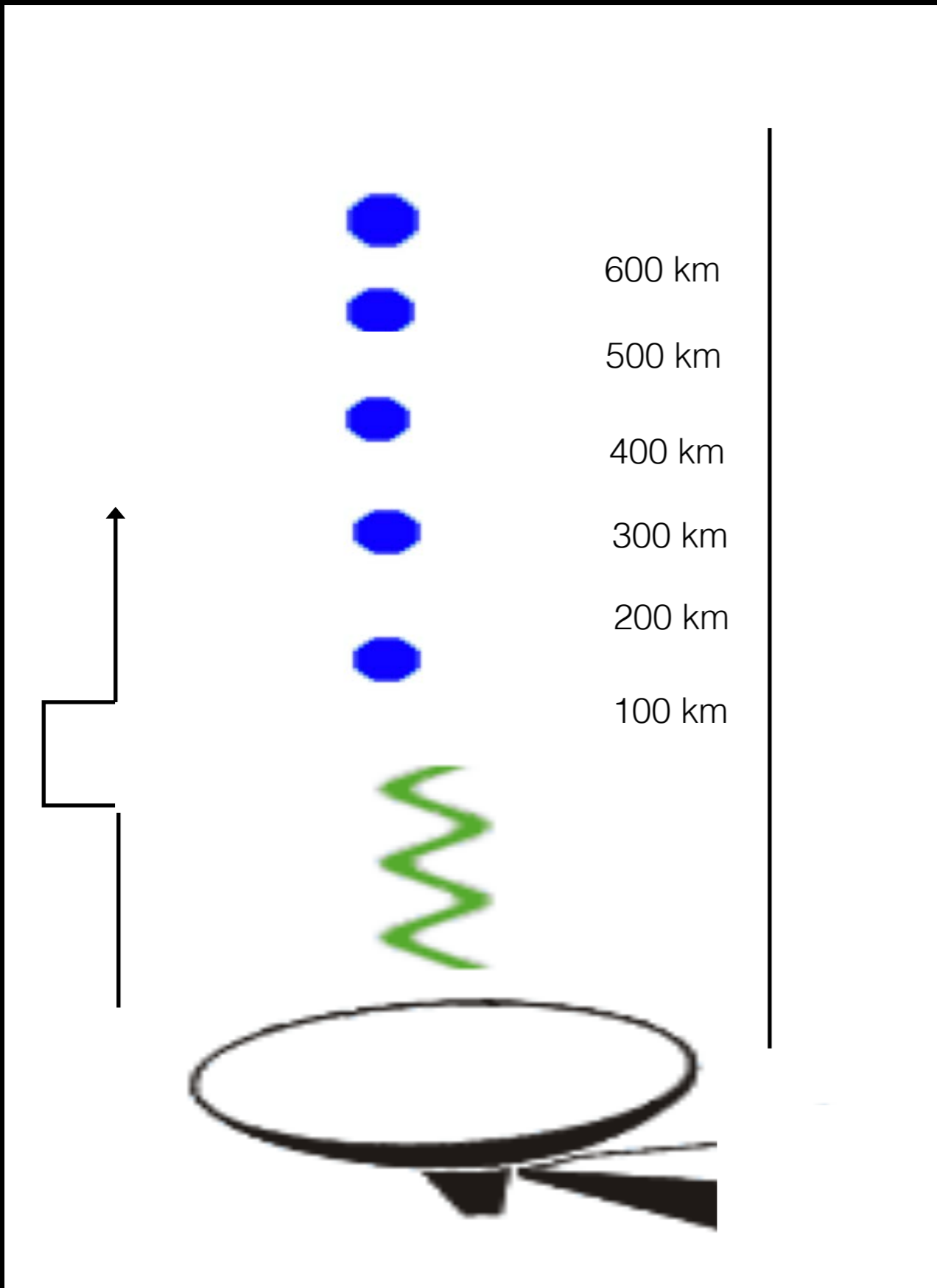
- **RADAR (Radio Detection And Ranging)**
- A technique for detecting and studying remote targets by transmitting a radio wave in the direction of the target and observing the reflection of the wave.
- **Radar** is an object detection system which uses radio waves to determine the range, altitude, direction, or speed of objects.  
(Wikipedia)

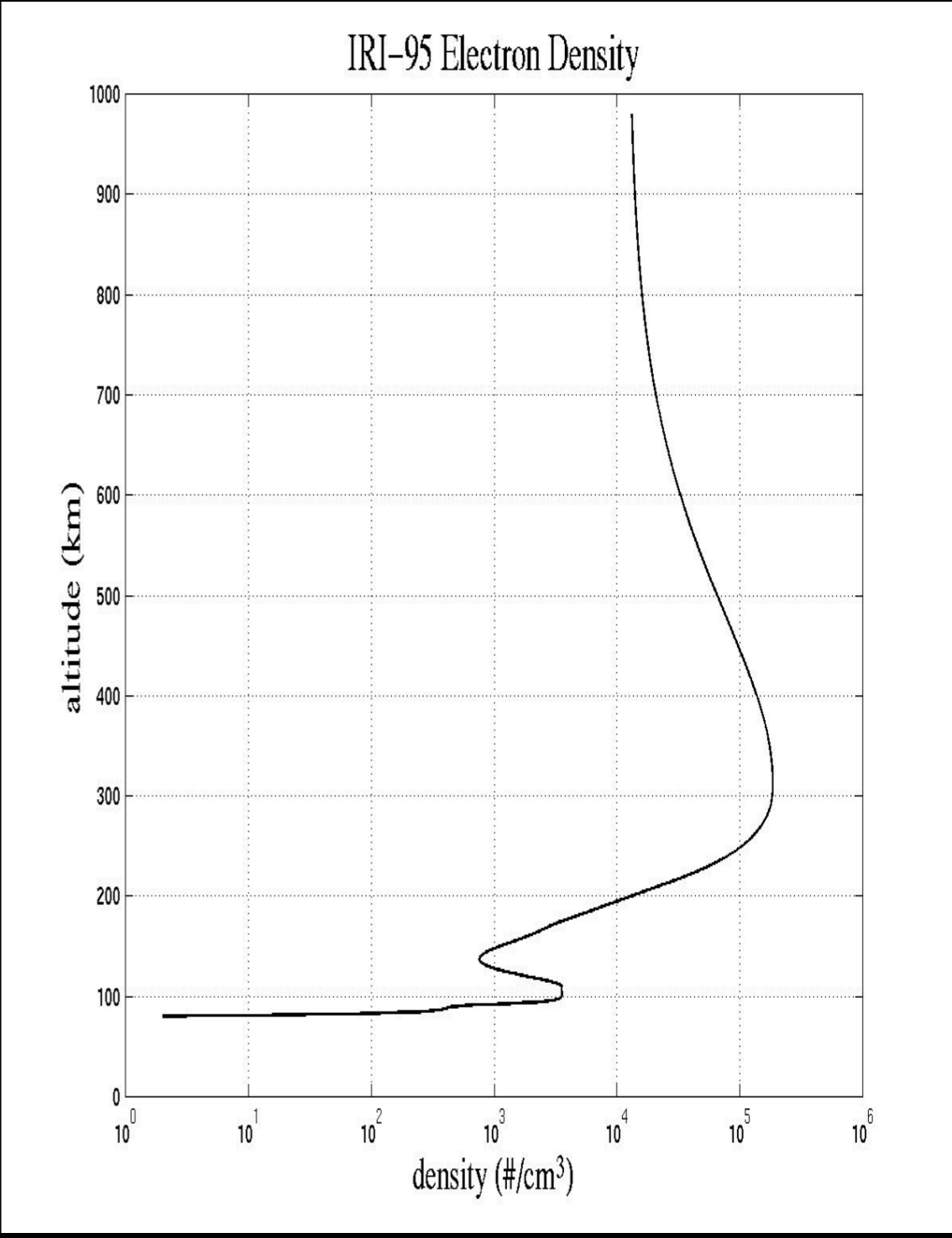
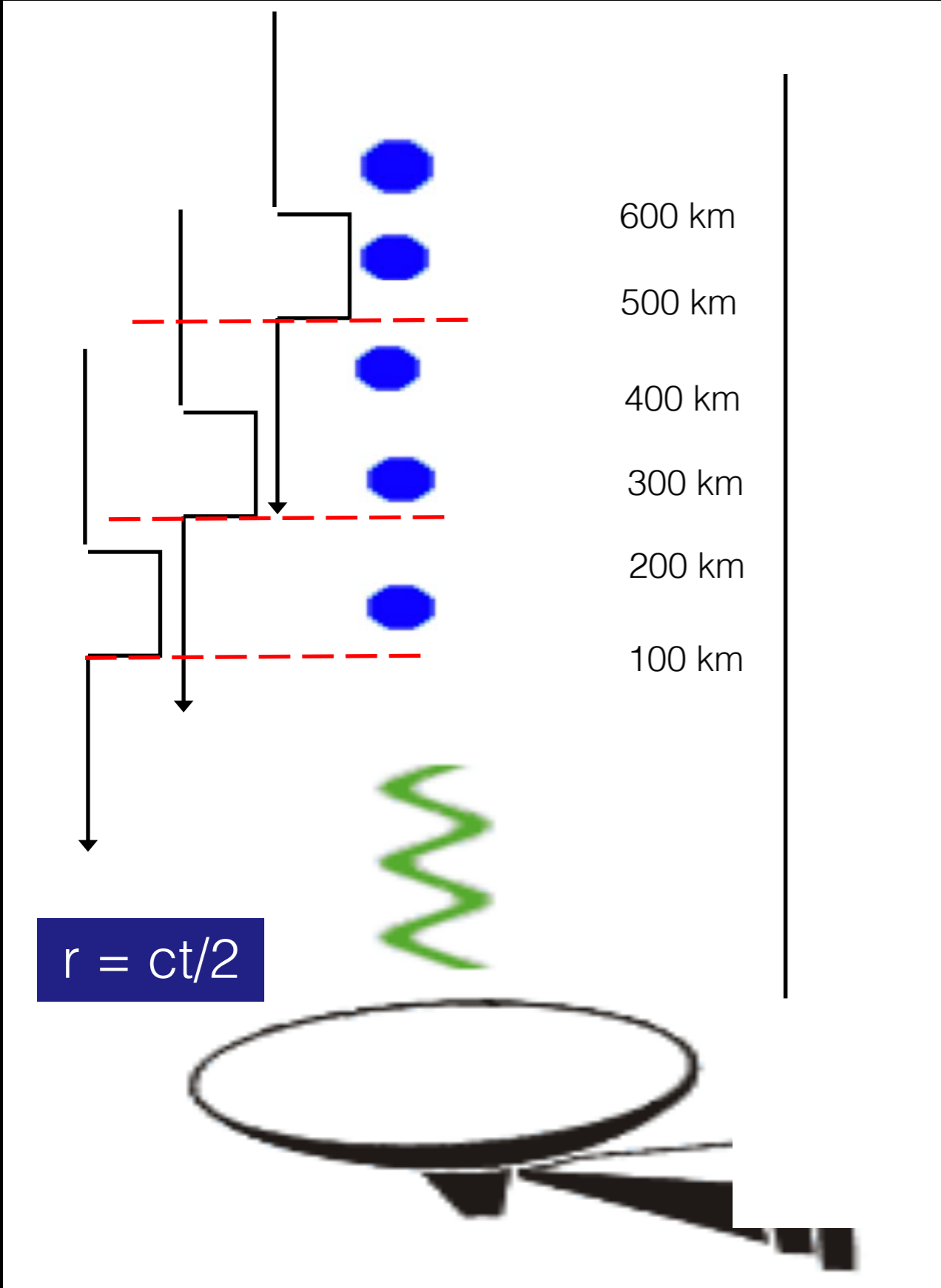


# Thermal fluctuating electrons



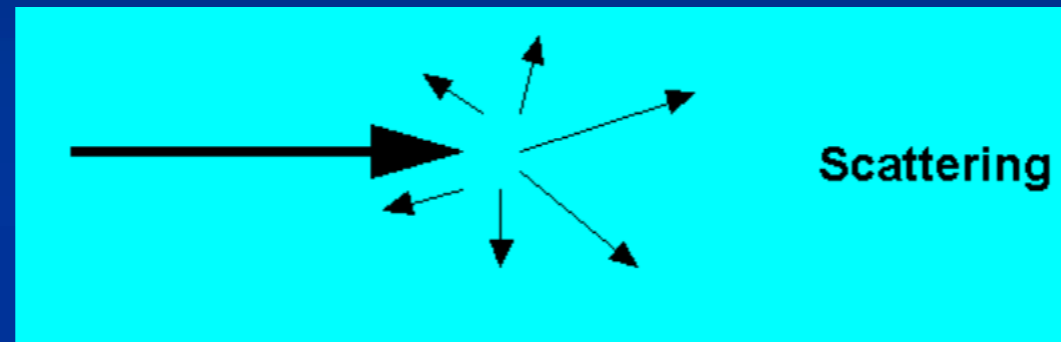
By Greg L at the English language Wikipedia, CC BY-SA 3.0,  
<https://commons.wikimedia.org/w/index.php?curid=1325234>



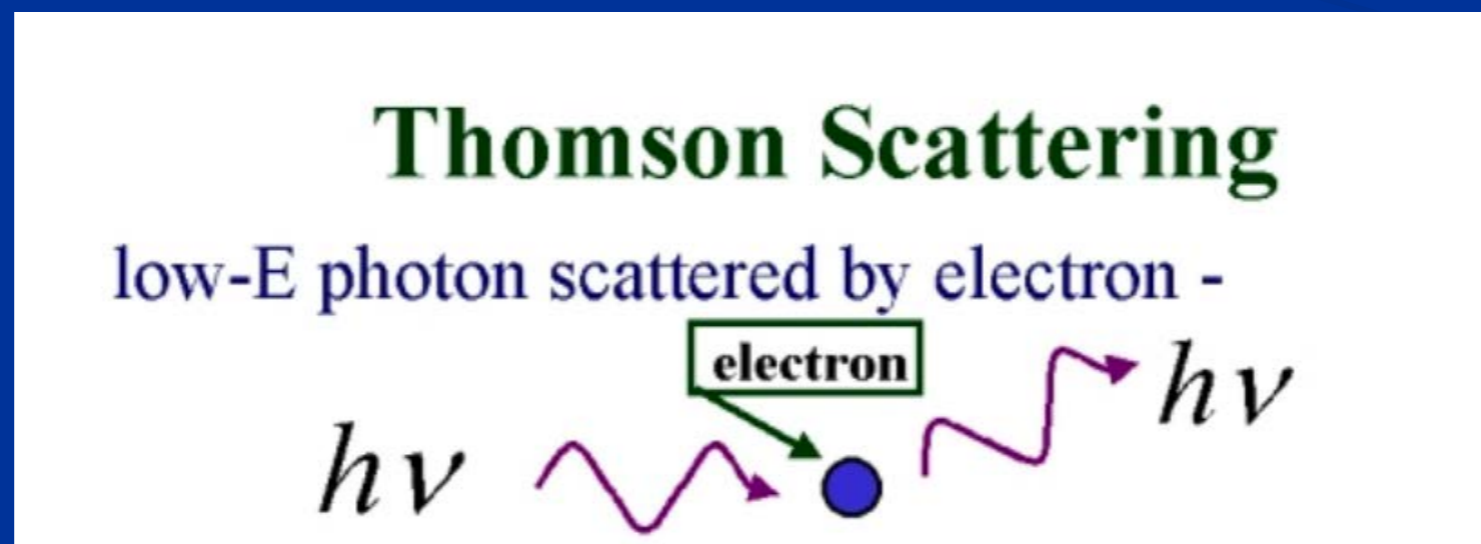


# Thomson scattering

- Thomson scattering is the elastic scattering of electromagnetic radiation by a free charged particle, as described by classical electromagnetism.



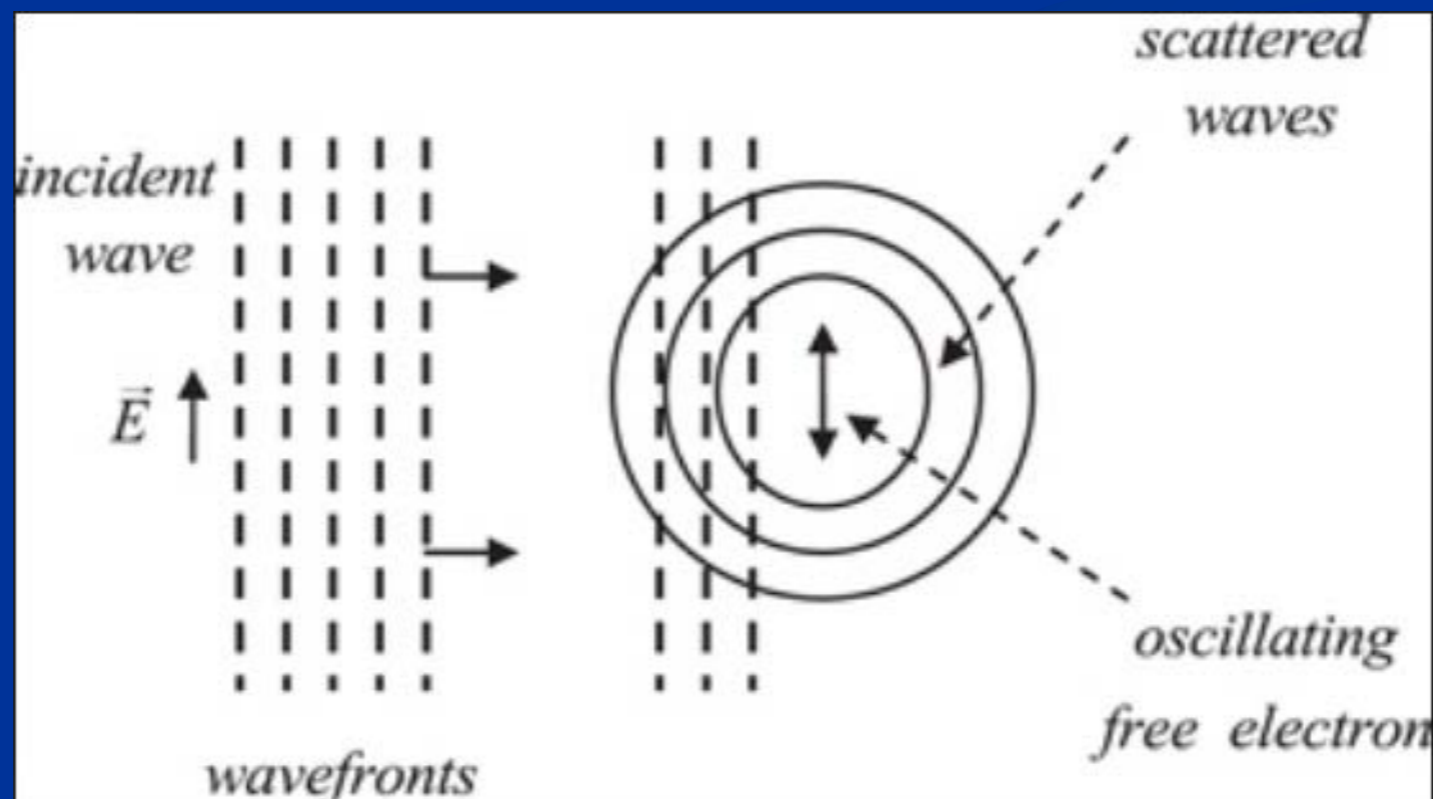
- In the low-energy limit, the electric field of the incident wave (radar wave) accelerates the charged particle, causing it, in turn, to emit radiation at the same frequency as the incident wave, and thus the wave is scattered.





# Thomson scattering (con.)

- As long as the motion of the particle is non-relativistic (i.e. its speed is much less than the speed of light), the main cause of the acceleration of the particle will be due to the electric field component of the incident wave, and the magnetic field can be neglected. The particle will move in the direction of the oscillating electric field, resulting in electromagnetic dipole radiation



# Thomson Scattering

$$E_x = E_0 e^{j(\omega t - kx)}$$

$$\vec{F} = q(\vec{E} + \vec{v} \times \vec{B})$$

$$v_x = -j \frac{q_e E_0}{m_e \omega} e^{j\omega t}$$

$$E_\phi = \frac{\mu_0 q_e^2}{4\pi m_e} \frac{\sin \phi}{r} e^{-jkr} E_0$$

$$\sigma_e = 4\pi \left( \frac{\mu_0 q_e^2}{4\pi m_e} \right)^2 \sin^2 \phi = 4\pi r_e^2 \sin^2 \phi$$
$$\approx 10^{-28} \sin^2 \phi \quad (\text{m}^2)$$

# Definition of Incoherent

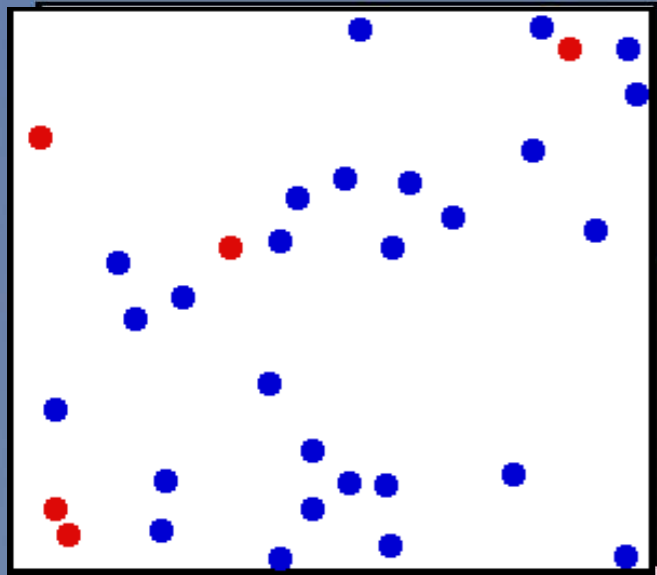
- Property of being coherent
- Antonym: incoherent
- Incoherent=Random
- Example: The drunk man made no sense. He was incoherent.
- In radar: Incoherent scatter is the process by which radio waves are randomly scattered by electrons in the ionosphere

# Definition of Incoherent

- Property of being coherent
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**Incoherent scatter is neither incoherent  
nor incomprehensible**

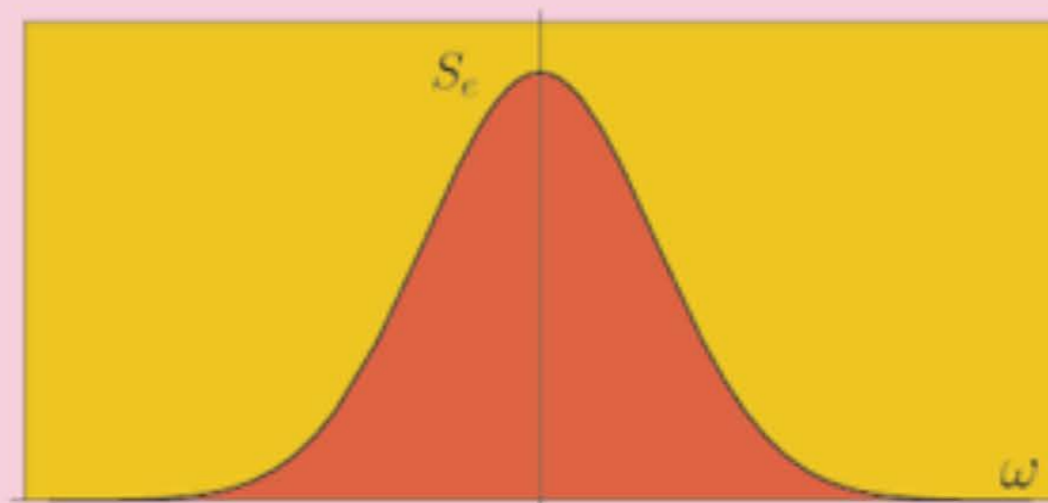
# For TRUE incoherent scatter...



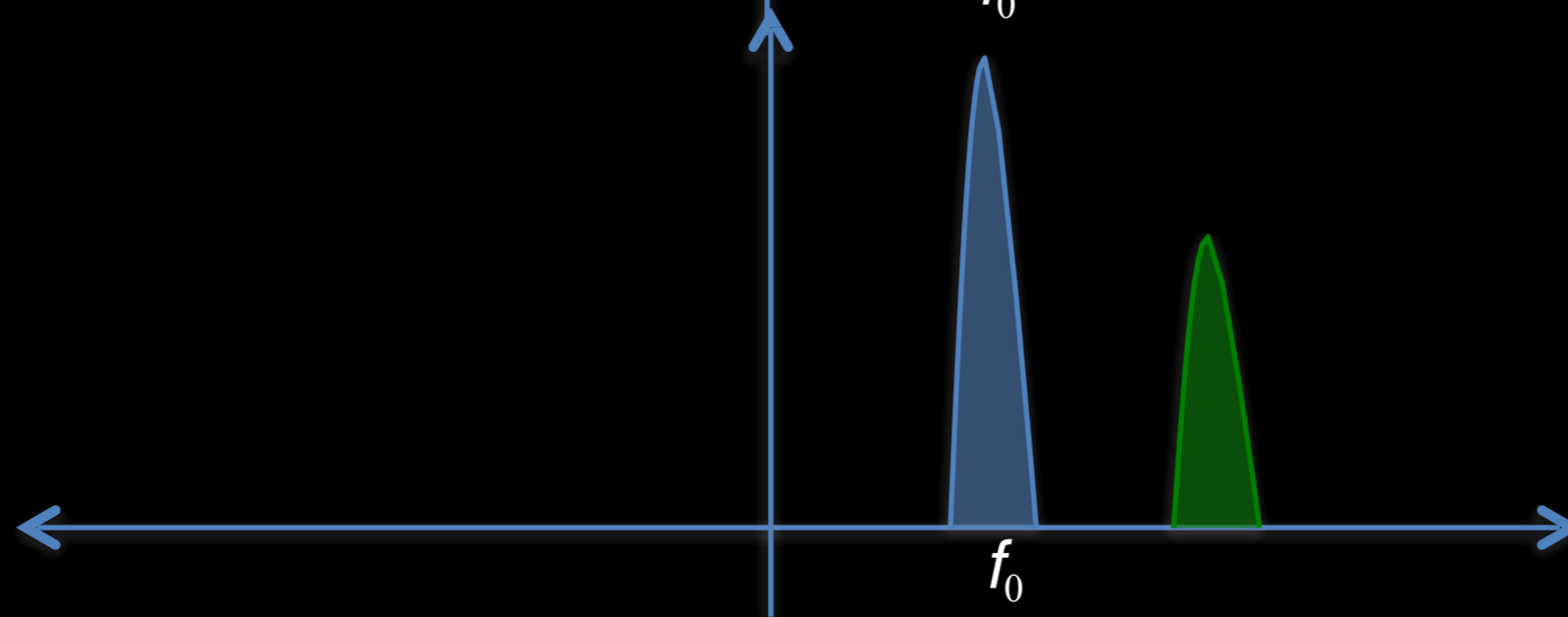
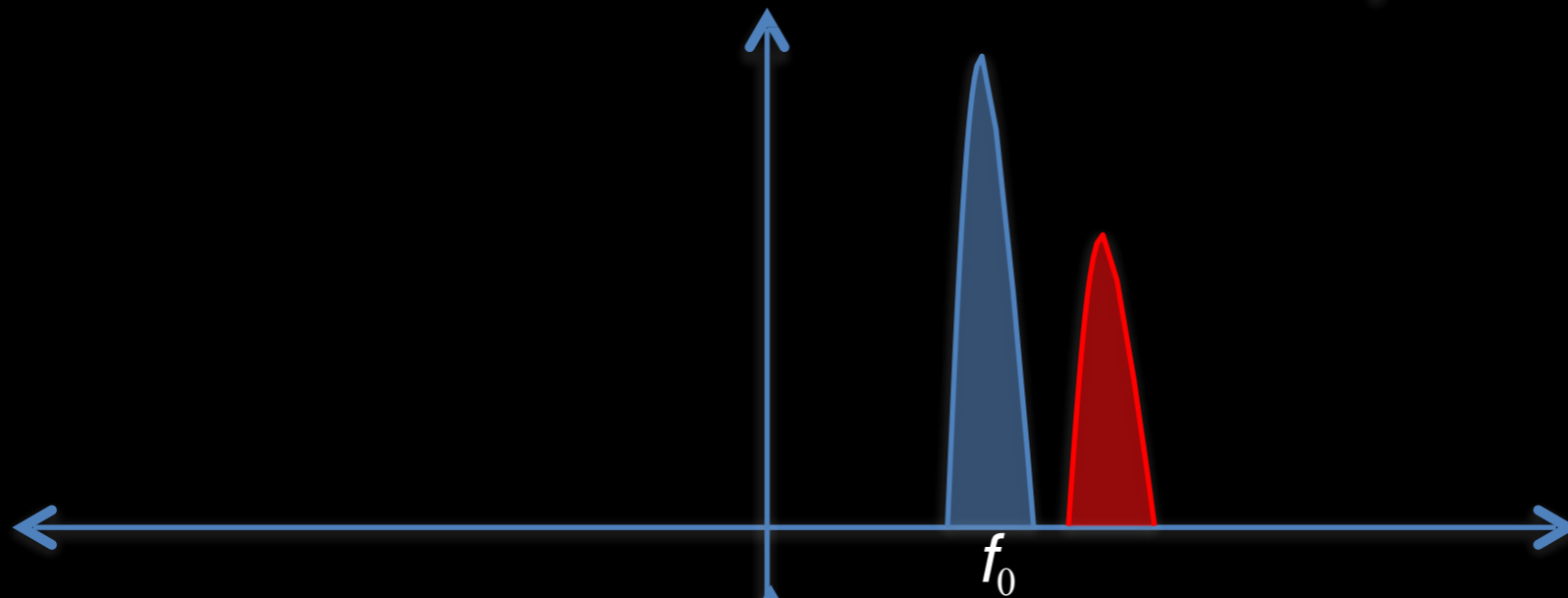
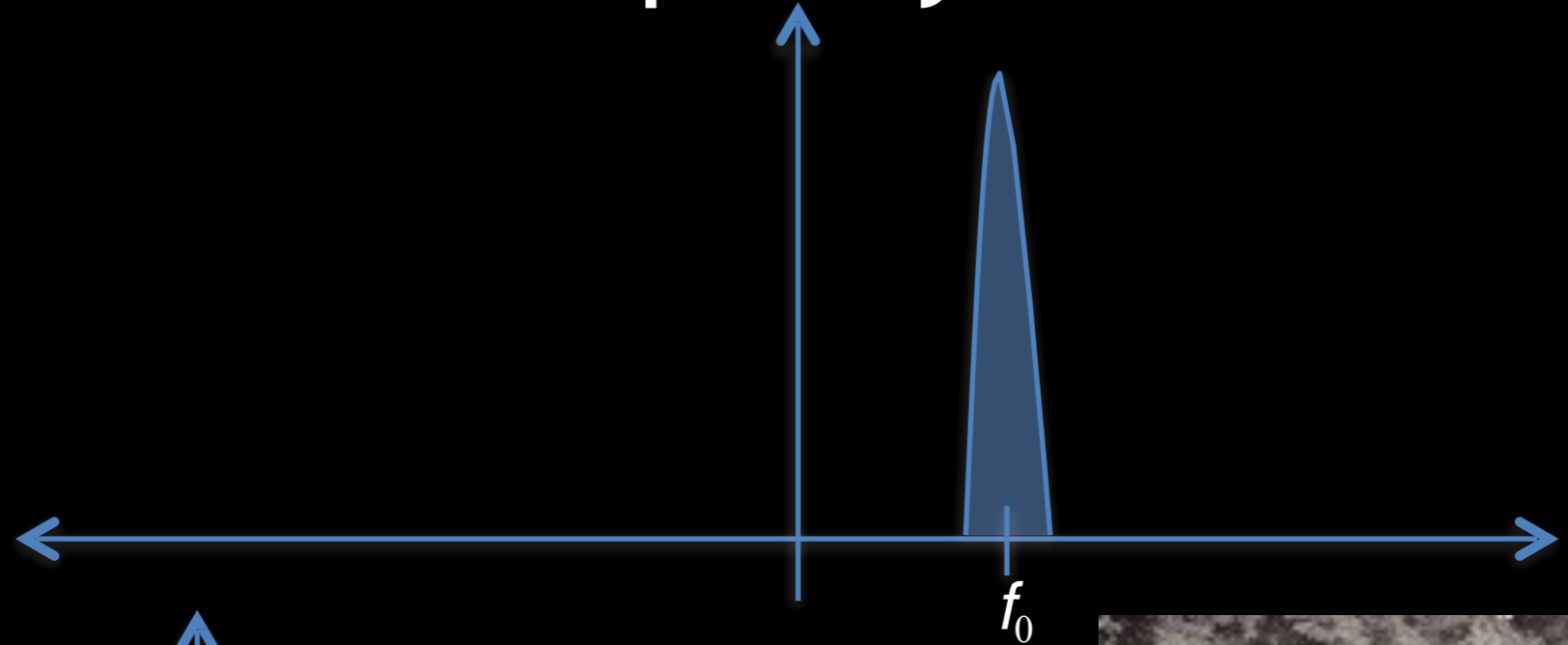
no collective interactions

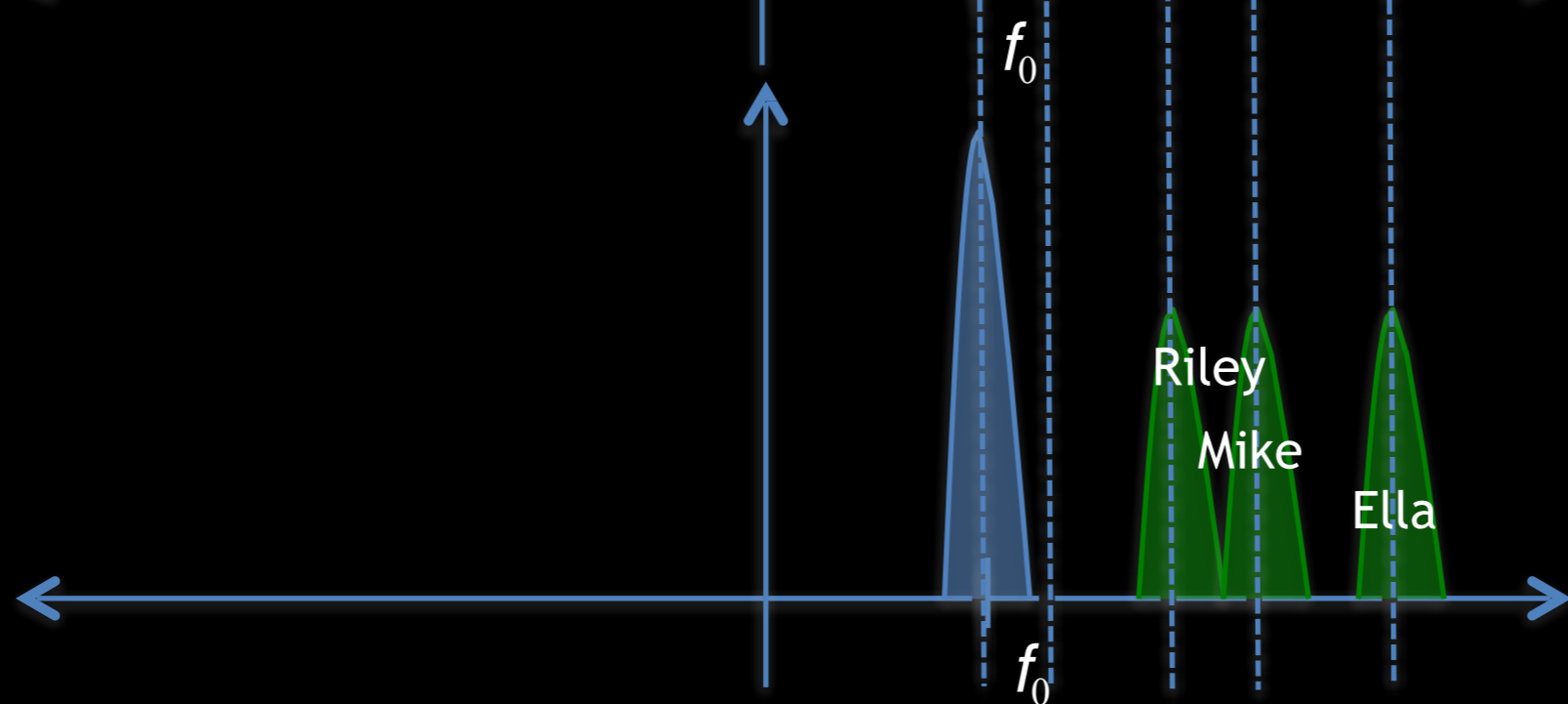
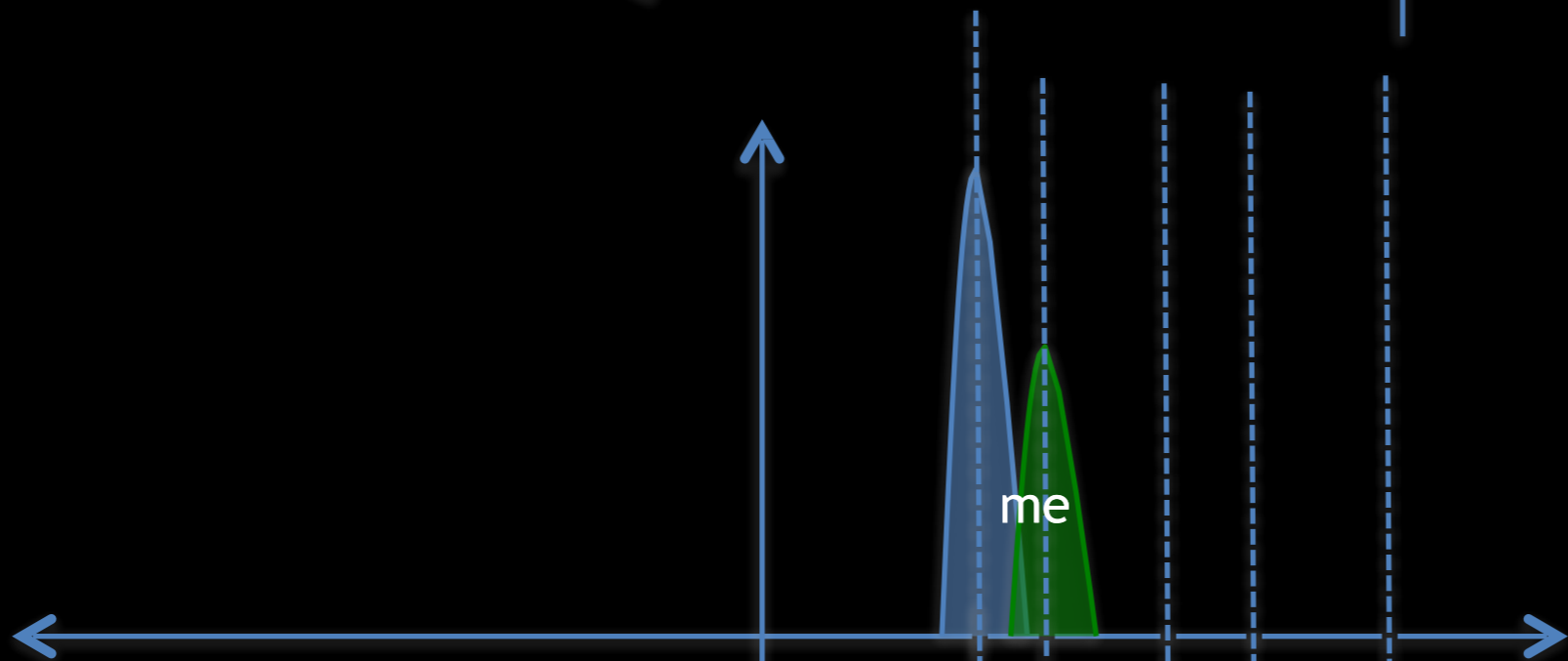
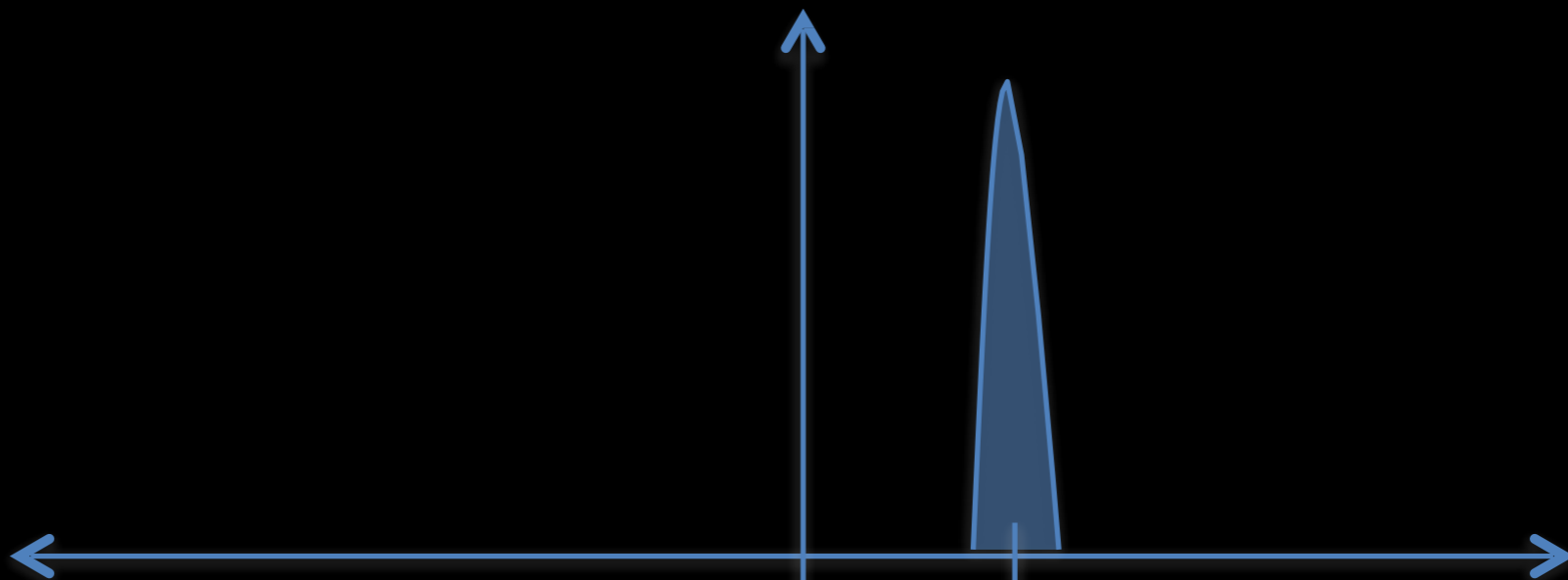
$$S_e(\mathbf{k}, \omega) = N_e \left| 1 - \frac{\chi_e(\mathbf{k}, \omega)}{\epsilon(\mathbf{k}, \omega)} \right|^2 \int d\mathbf{v} f_e(\mathbf{v}) \delta(\omega - \mathbf{k} \cdot \mathbf{v}) + N_i \left| \frac{\chi_e(\mathbf{k}, \omega)}{\epsilon(\mathbf{k}, \omega)} \right|^2 \int d\mathbf{v} f_i(\mathbf{v}) \delta(\omega - \mathbf{k} \cdot \mathbf{v})$$

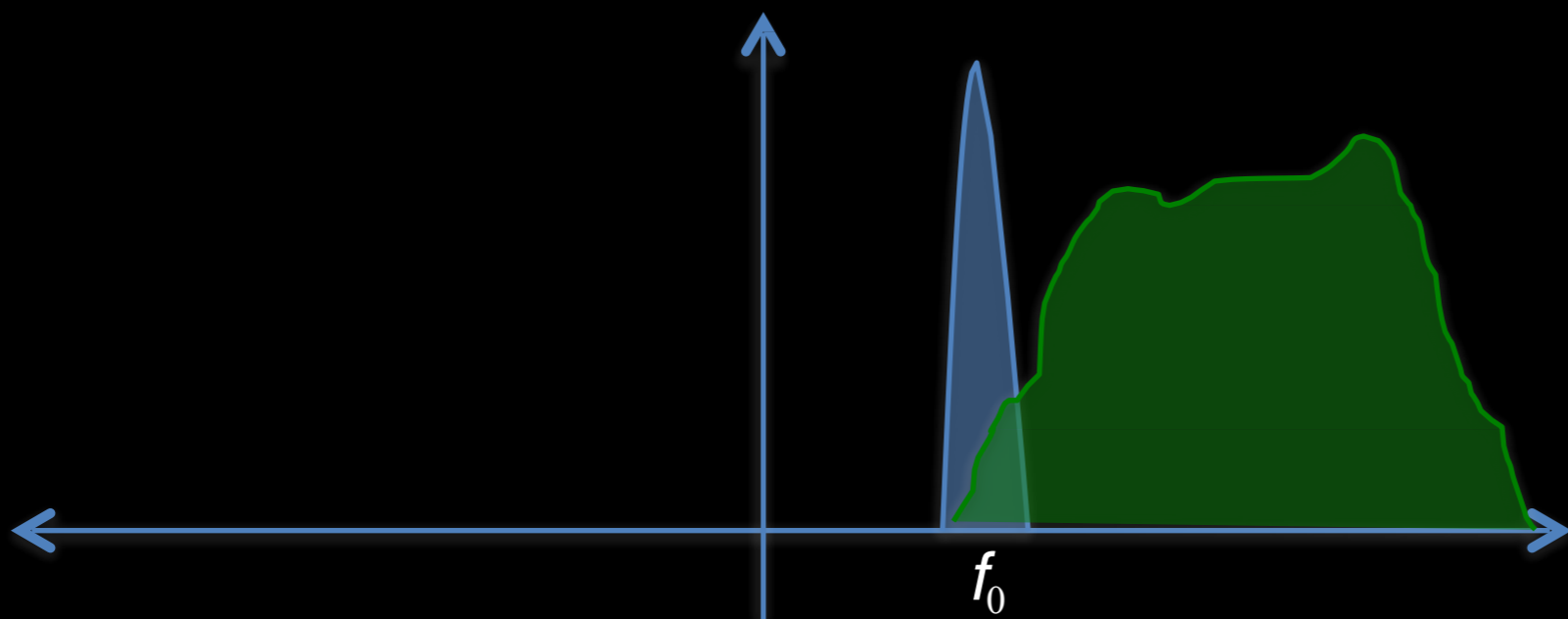
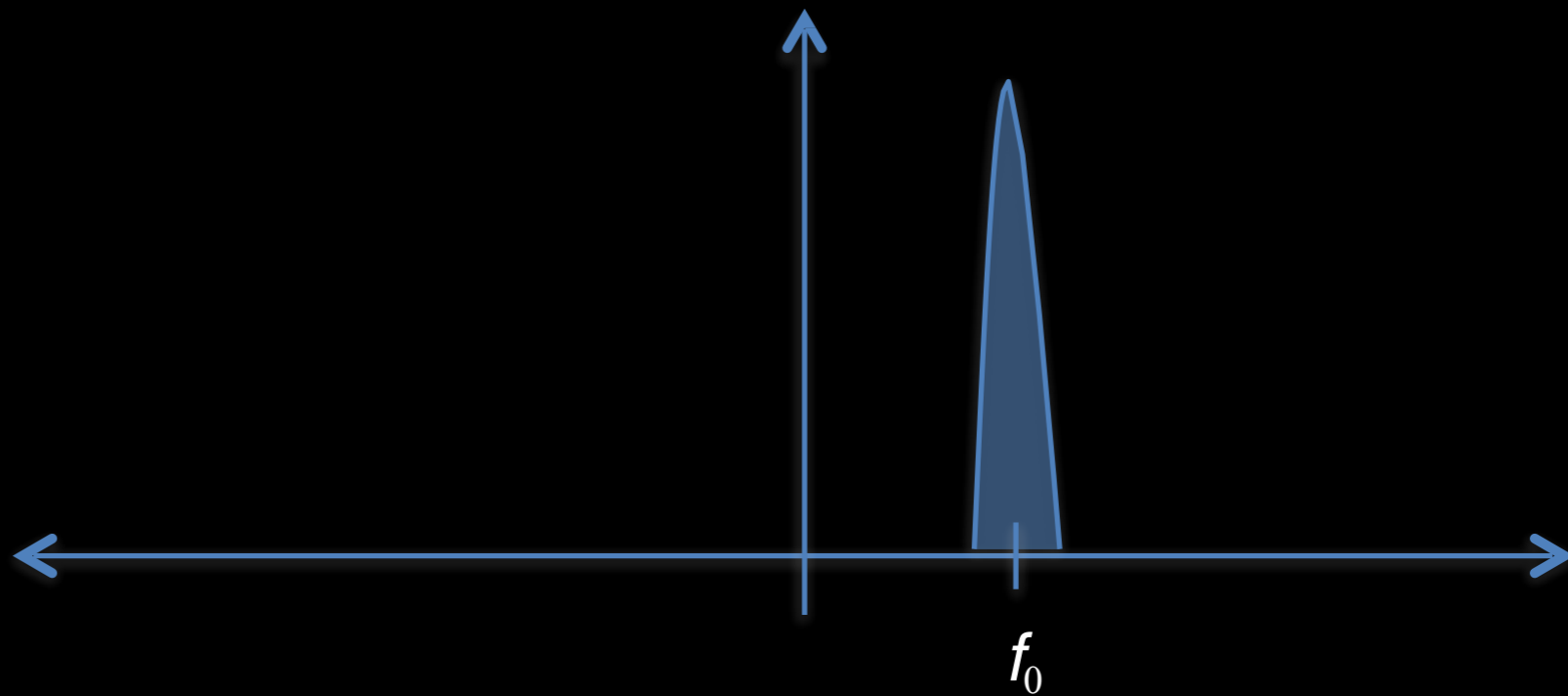
$$S_e(\mathbf{k}, \omega) = N_e \int d\mathbf{v} f_e(\mathbf{v}) \delta(\omega - \mathbf{k} \cdot \mathbf{v})$$



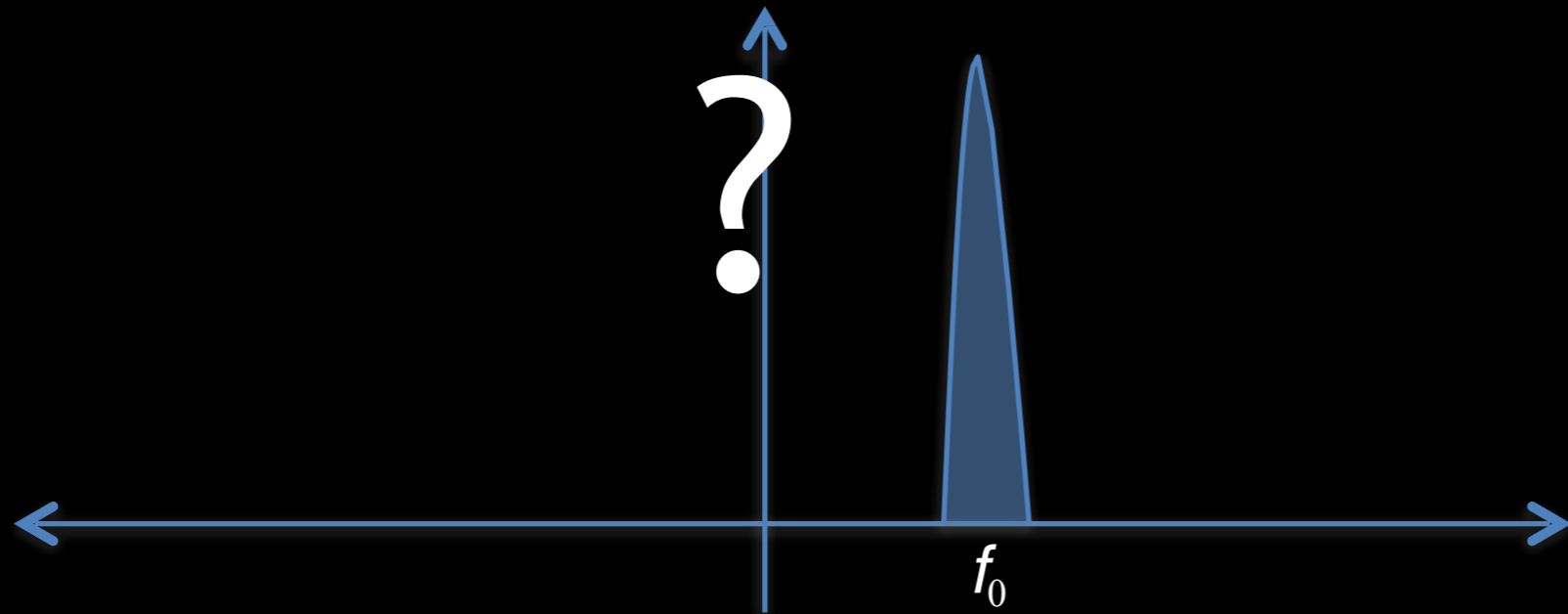
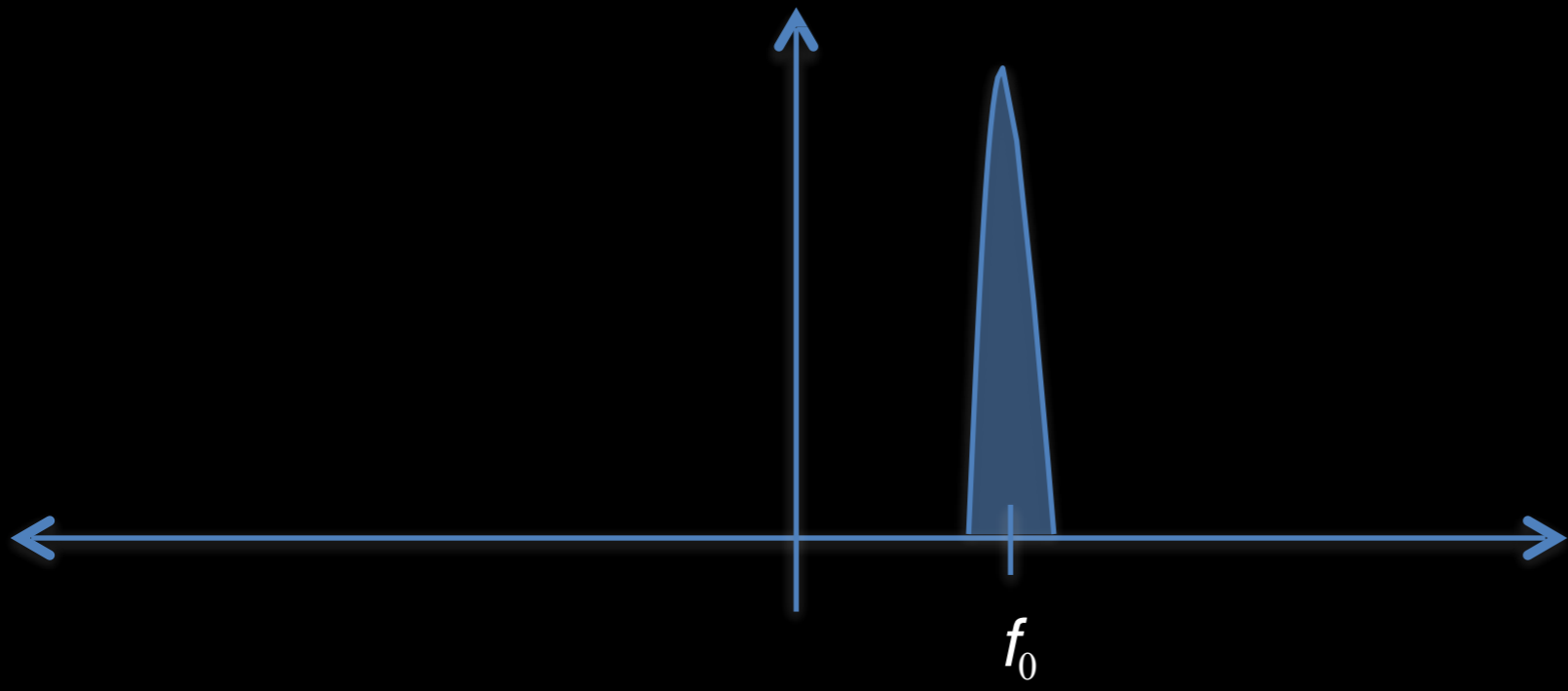
# Doppler Radar - frequency domain

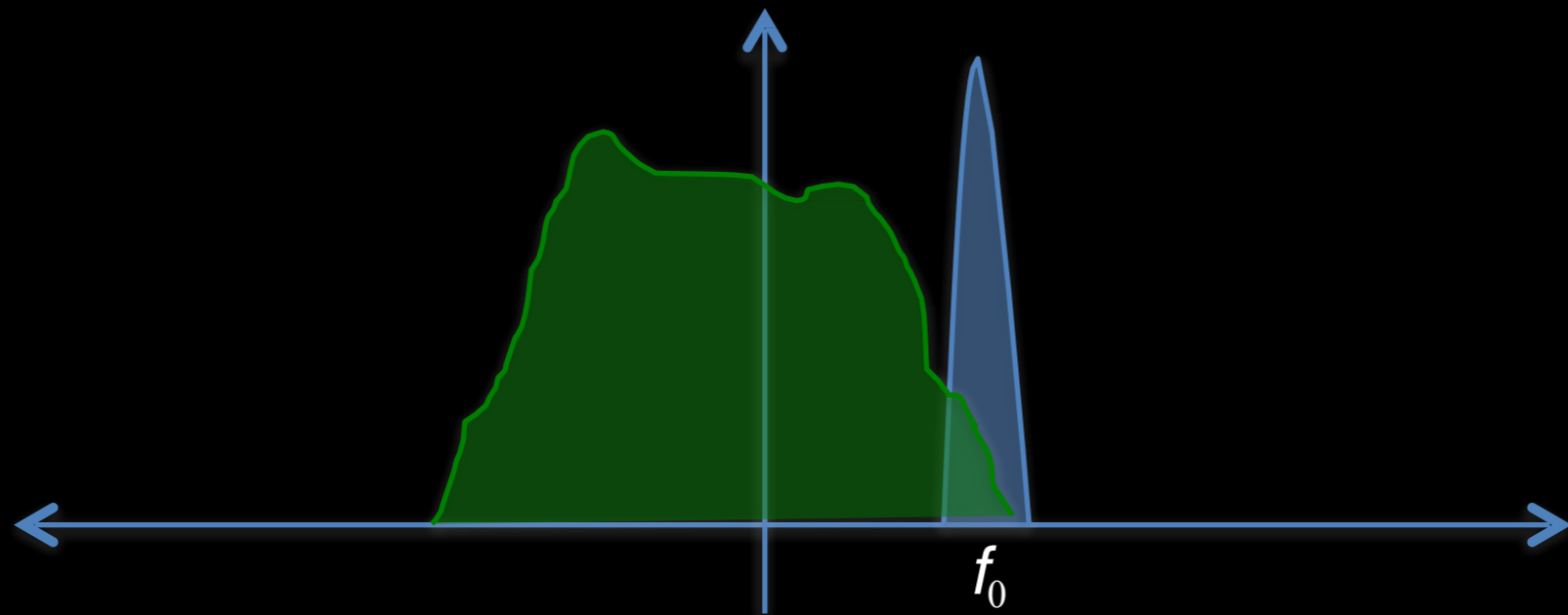
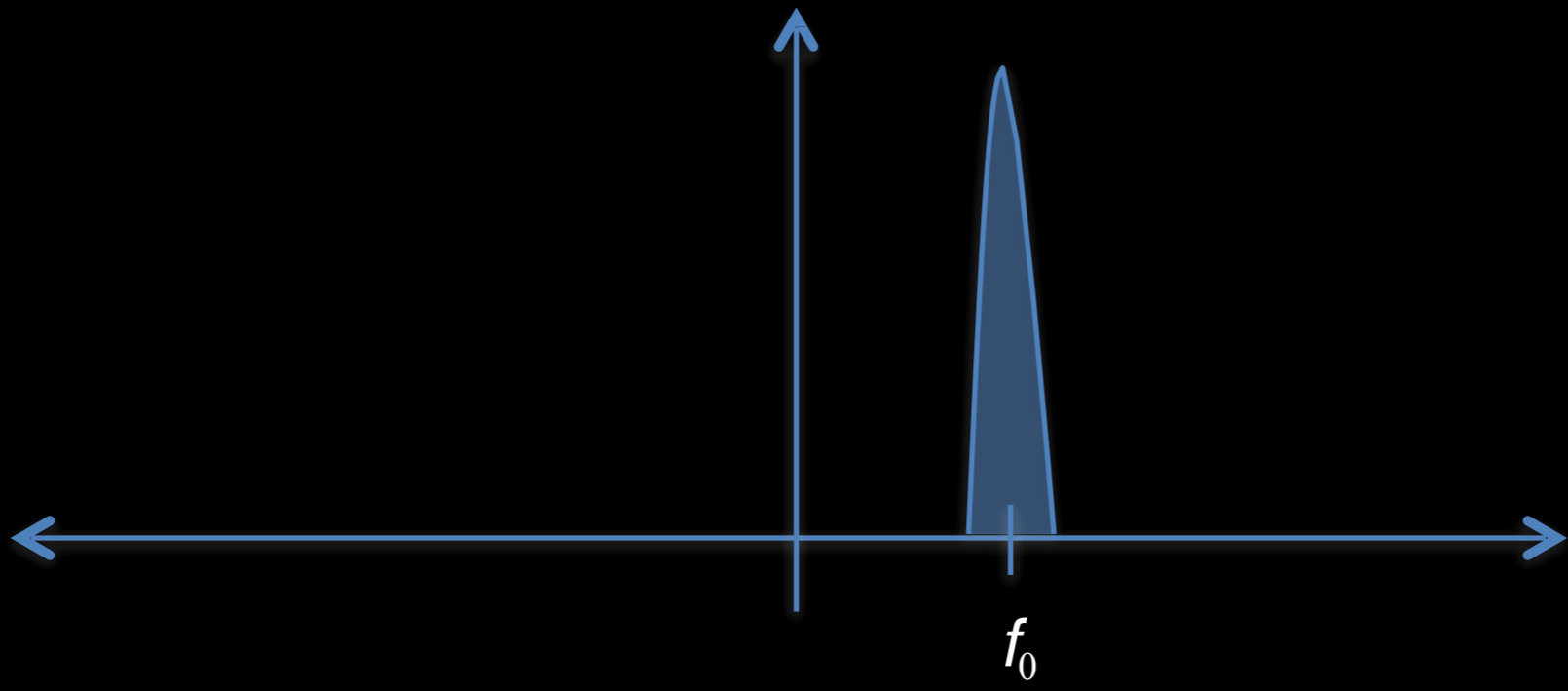




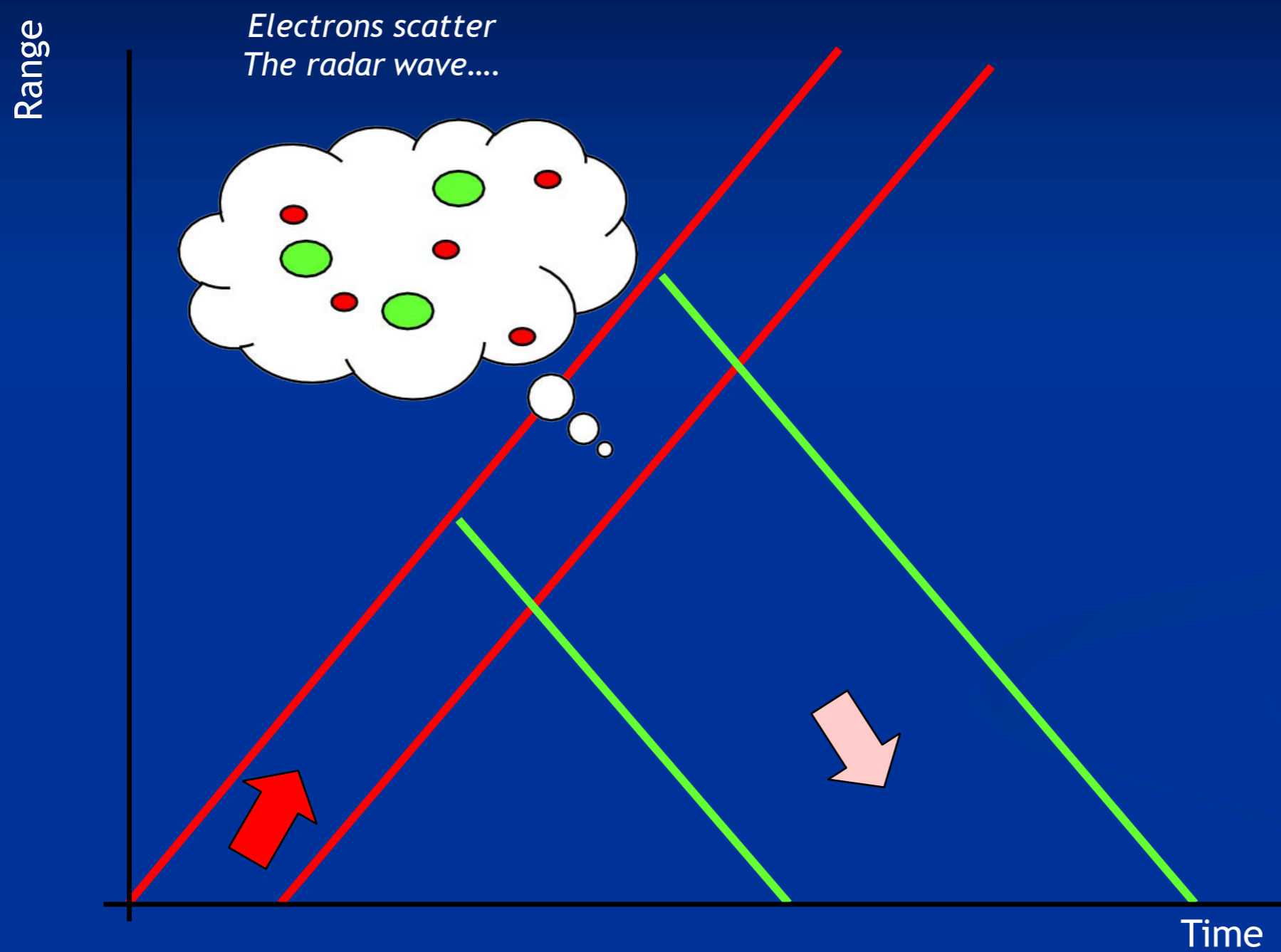








# How ISRs work



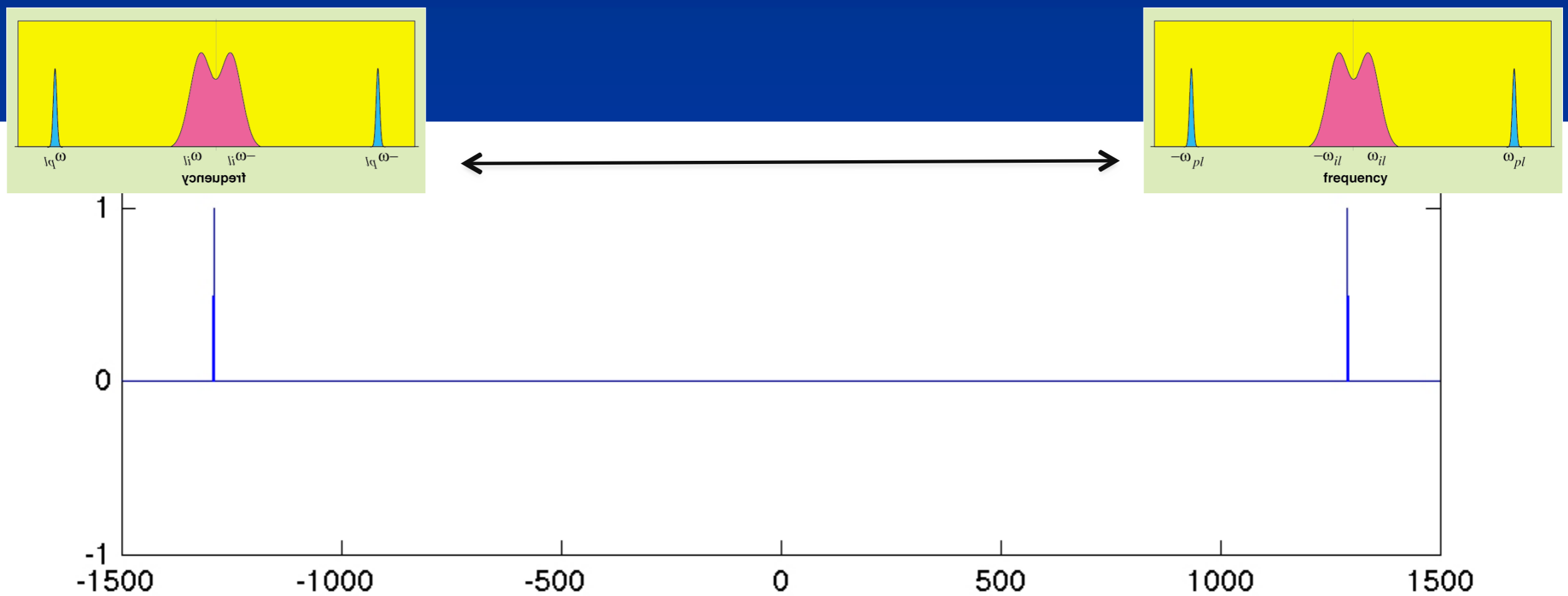
High power transmitter

Very sensitive receiver

Only  $\sim 0.000000000000000000000001\%$  of the transmitted power is returned!

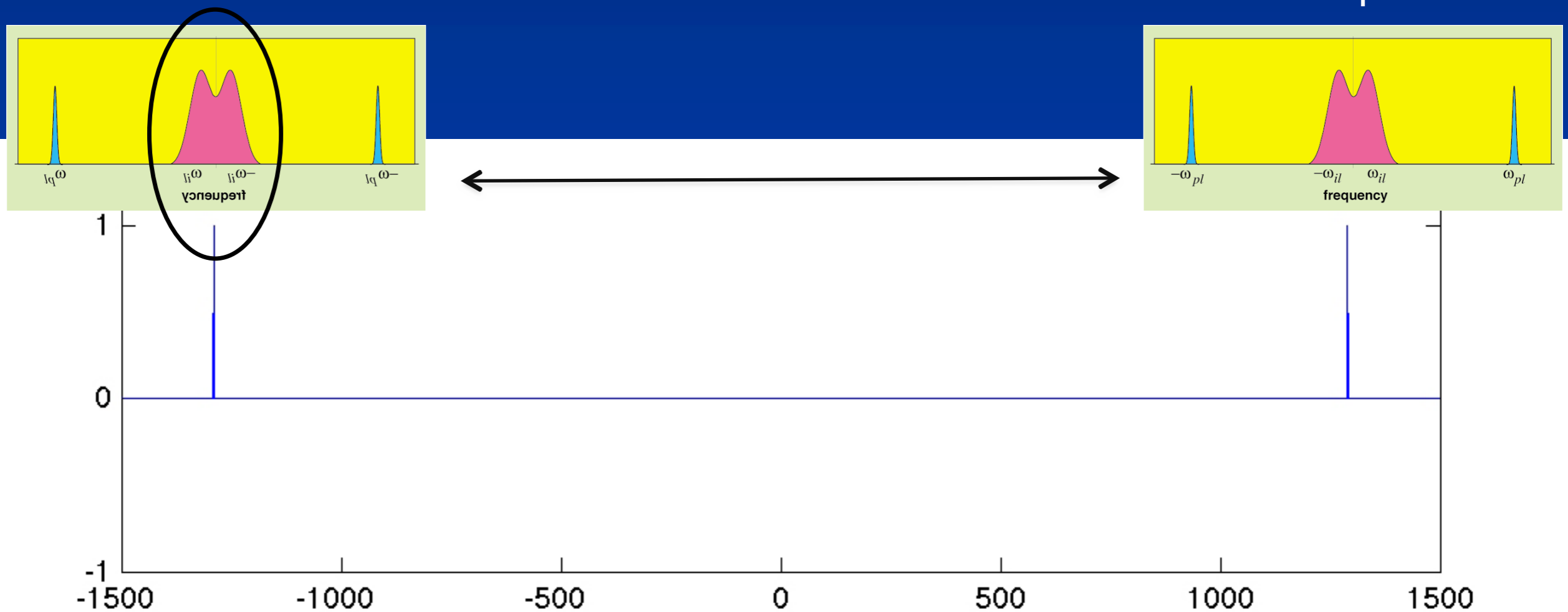
# IS Received Spectrum

The individual sides are **not** symmetric about their center frequencies!



# IS Received Spectrum

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# Outline

Distribution of ISRs across the world

Introduction to ISR principles and Incoherent Scatter Spectrum

→ Ion acoustic waves, Langmuir waves, Debye Spheres, Landau Damping .....

- **We only see scattering from the electrons**  
...but they also tell the story about the ion  
dynamics...

# Collective behavior

- There are a number of wave modes existing inherently in the ionospheric plasma



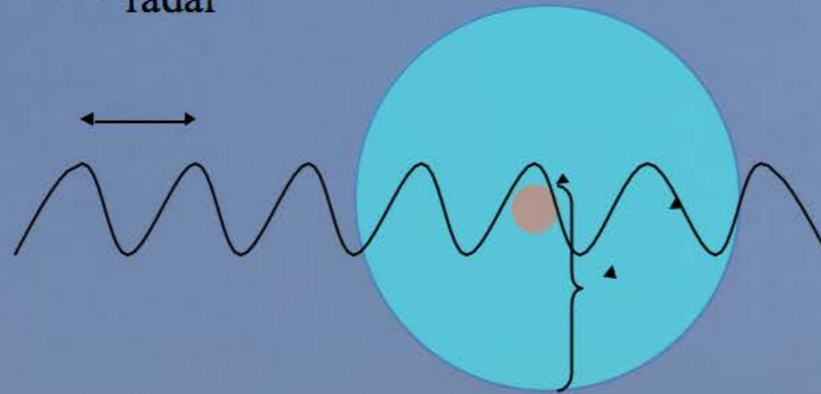
# Debye length dependence

Ion

Electron cloud

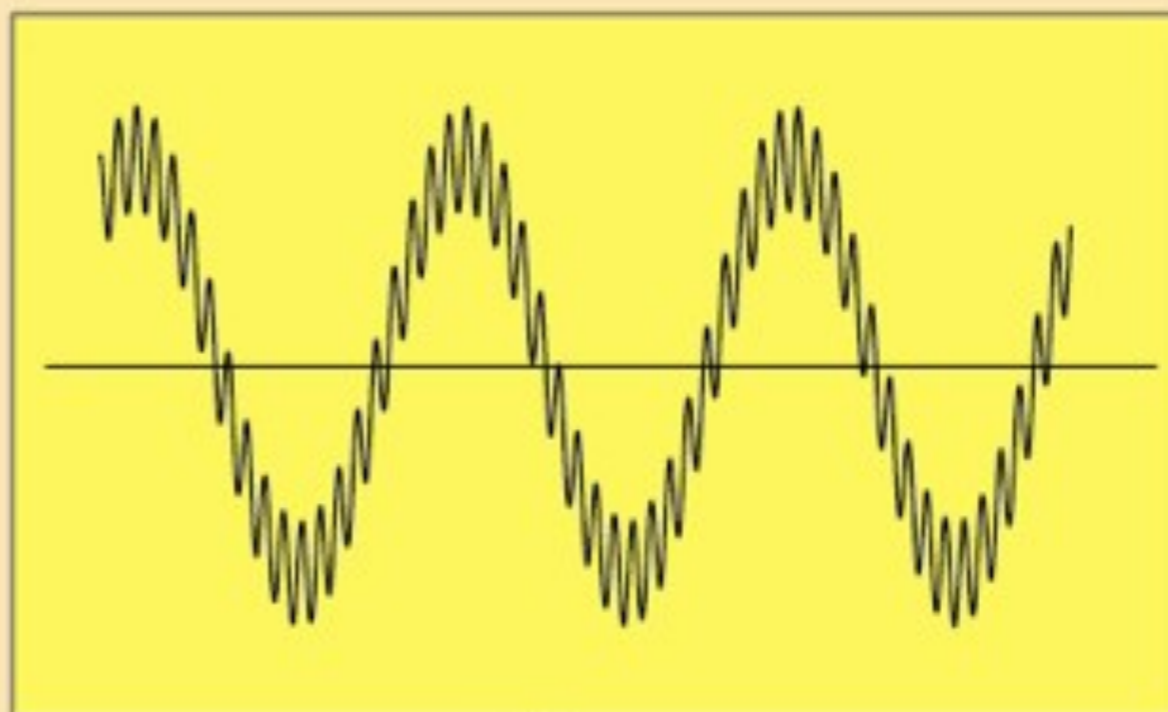
Debye length  $\lambda_D$

$$\lambda_{\text{radar}} \propto 1/k_{\text{radar}}$$

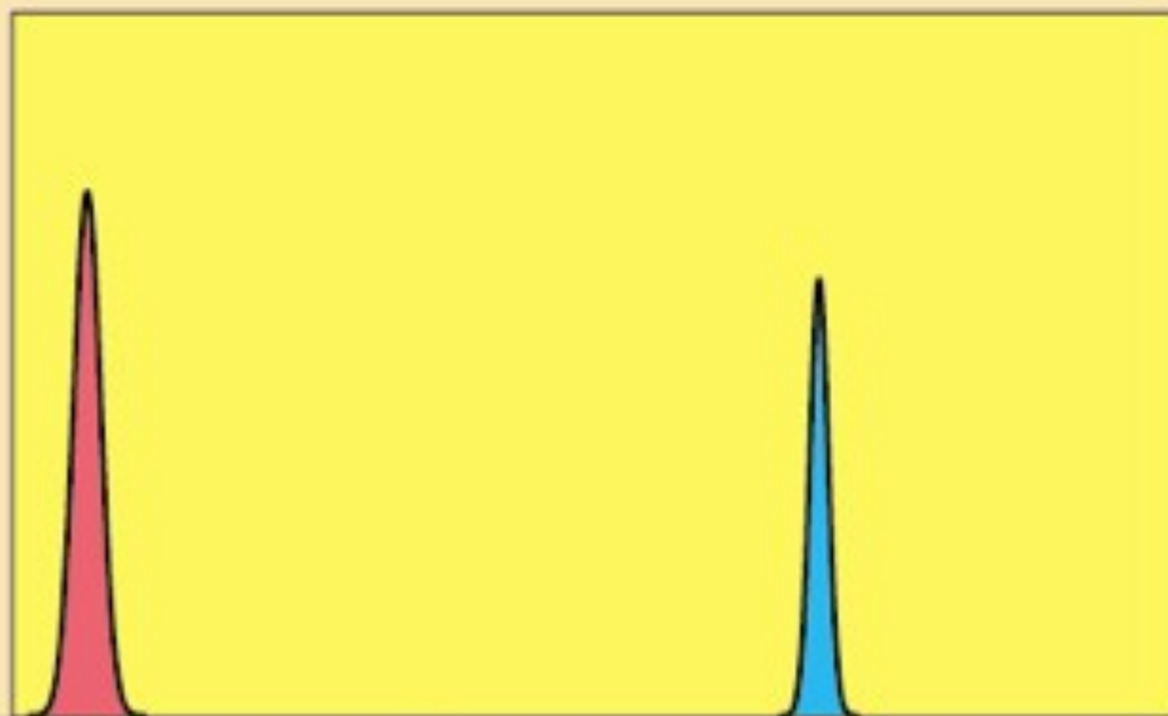


$$(\lambda_D / \lambda_{\text{radar}})^2 > 1$$

- $(k_{\text{radar}} \lambda_D)^2 > 1$
- No collective interactions

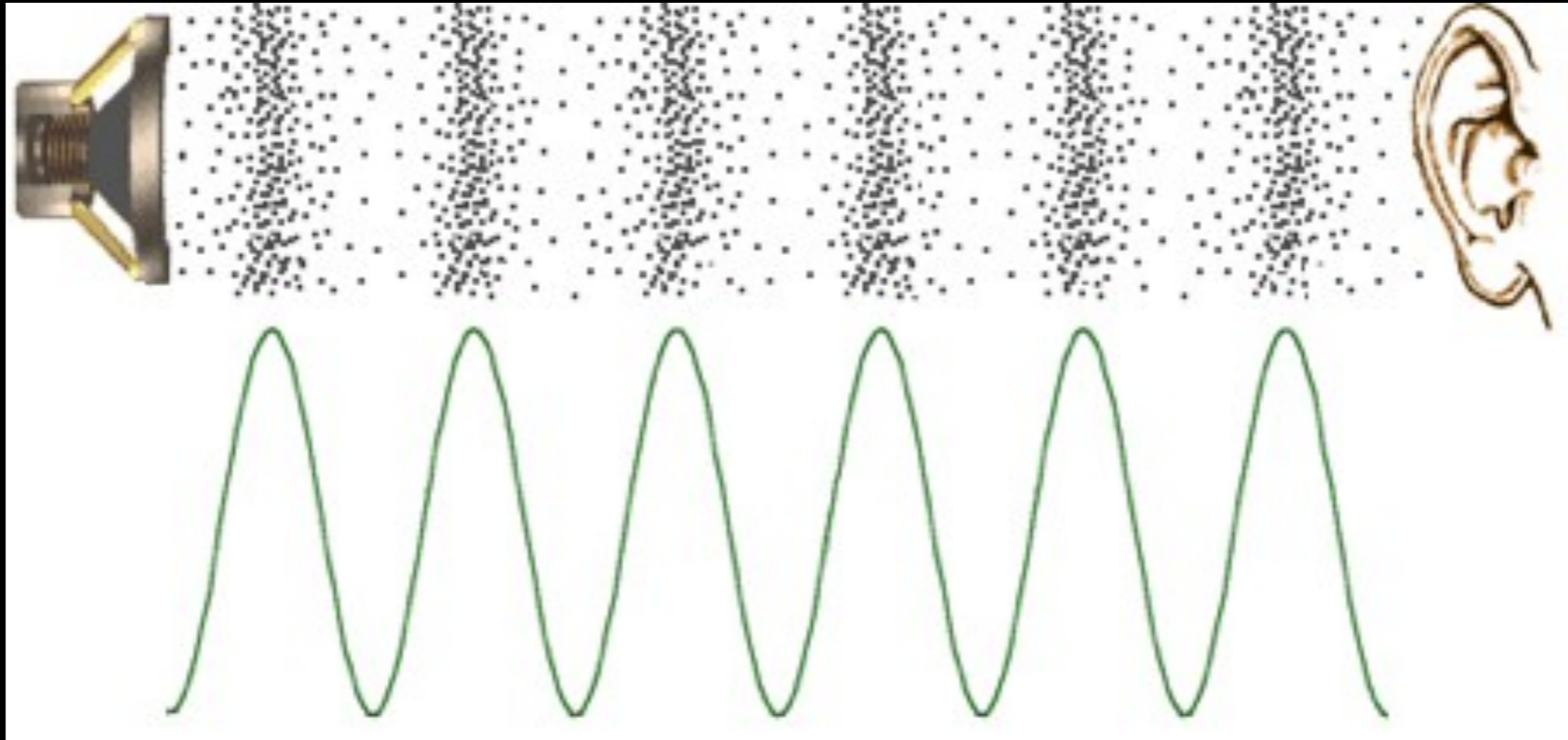


time

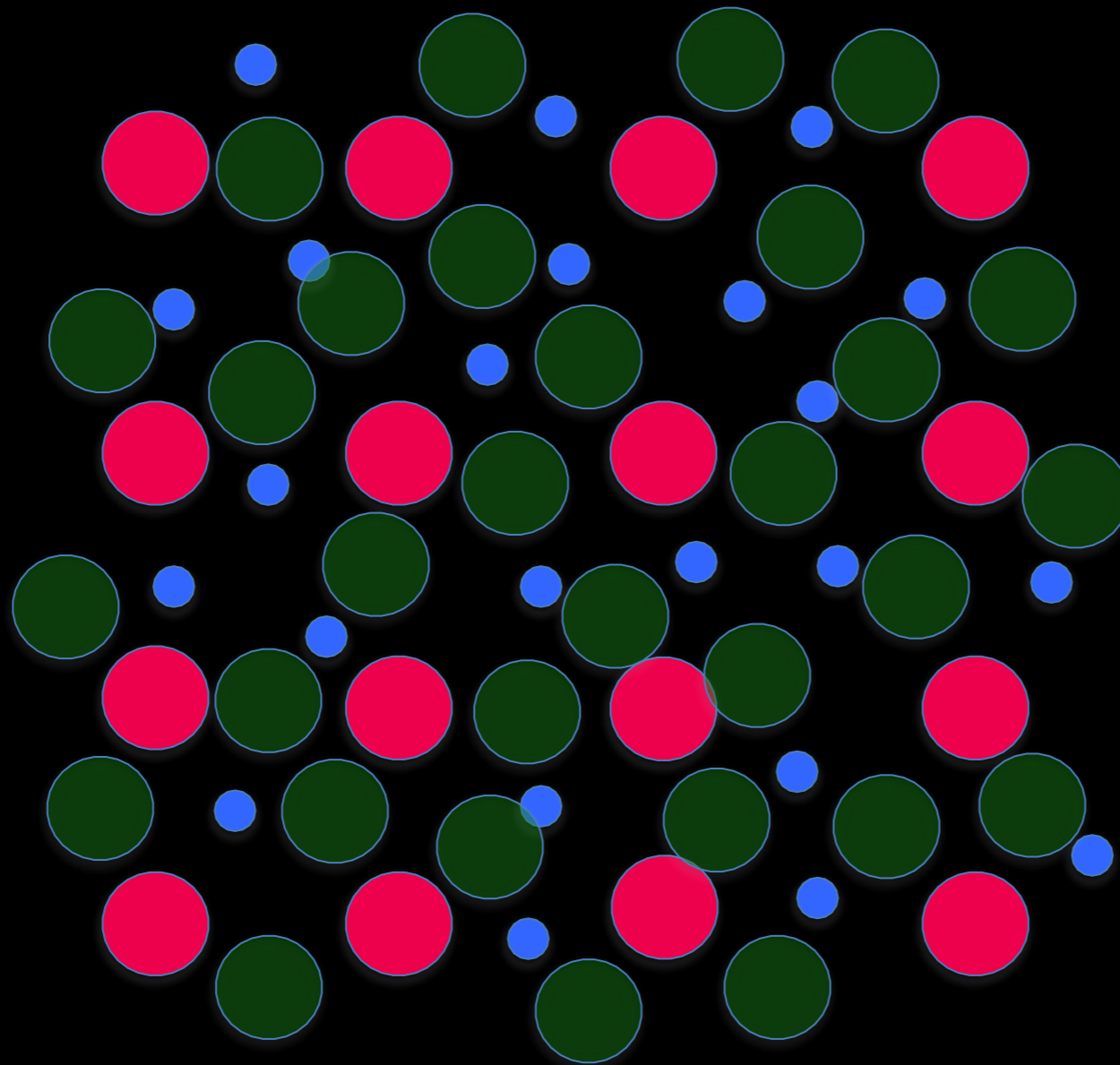


frequency

# Ion Acoustic Waves



“Pressure” waves in the ion density



Neutrals



Positive Ions



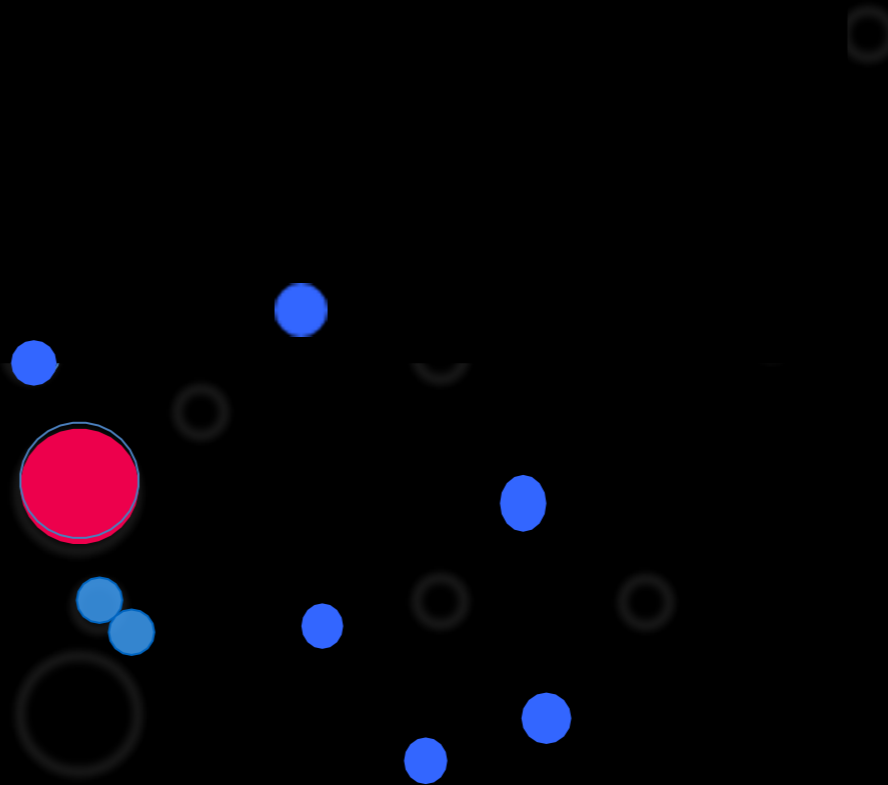
Electrons



Positive Ions



Electrons

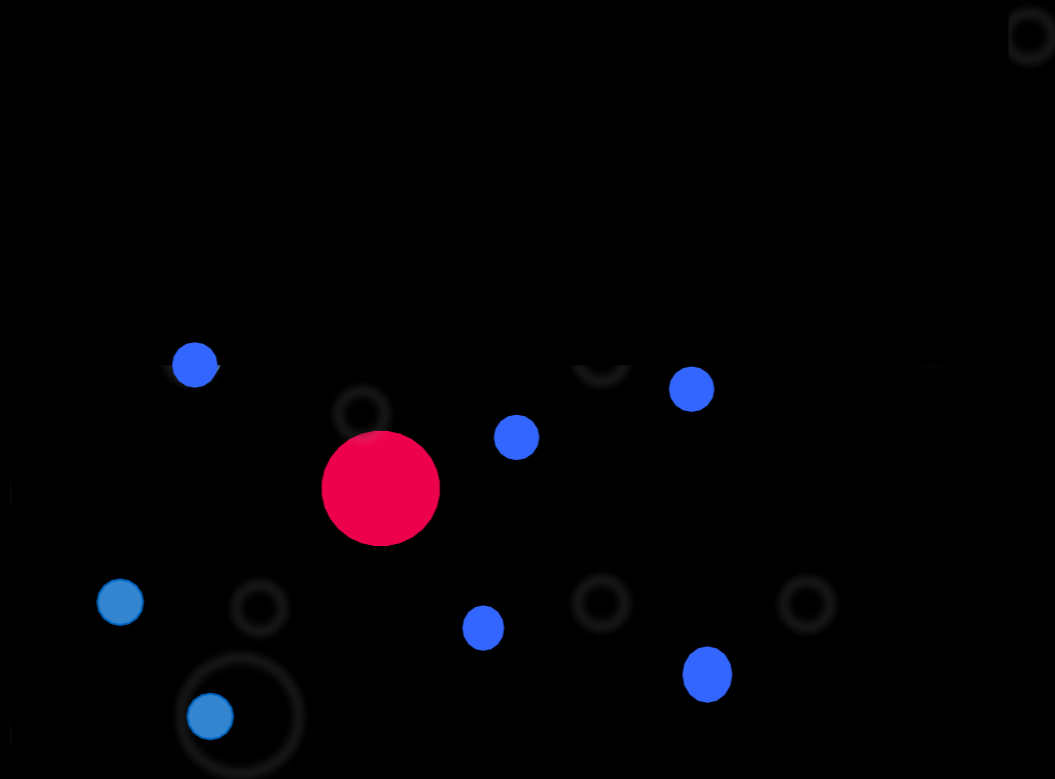




Positive Ions



Electrons

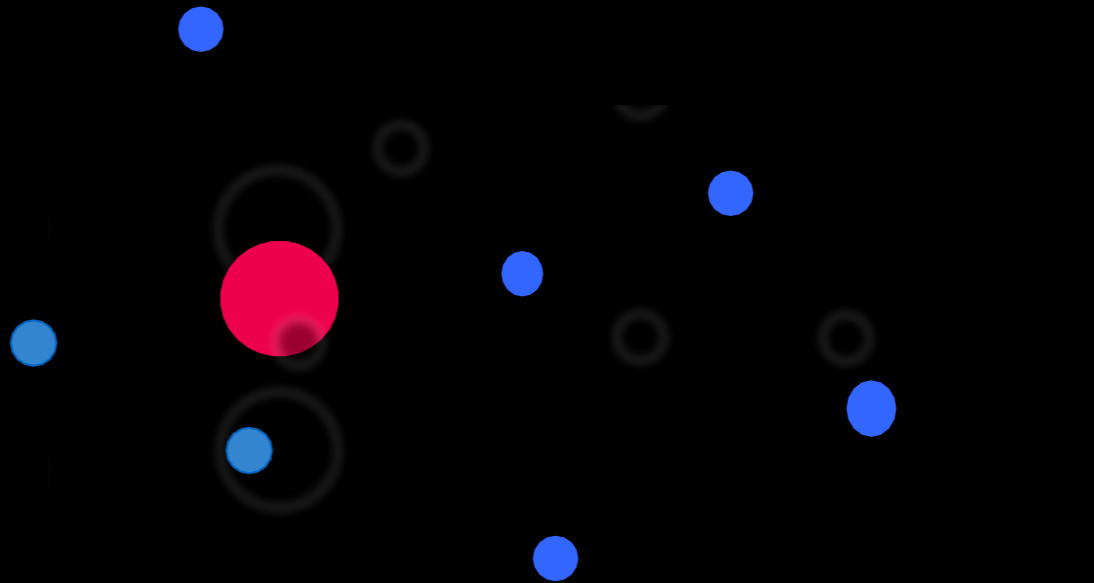




Positive Ions



Electrons

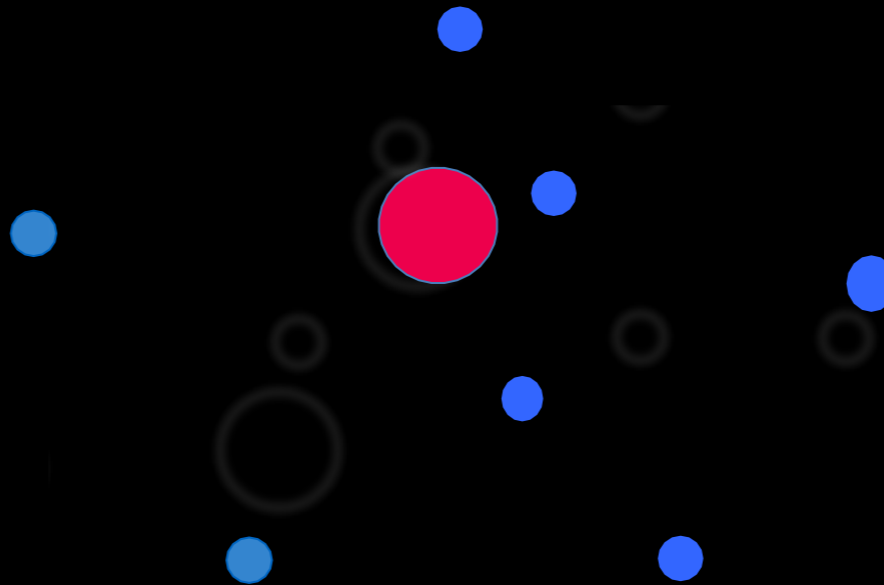




Positive Ions



Electrons



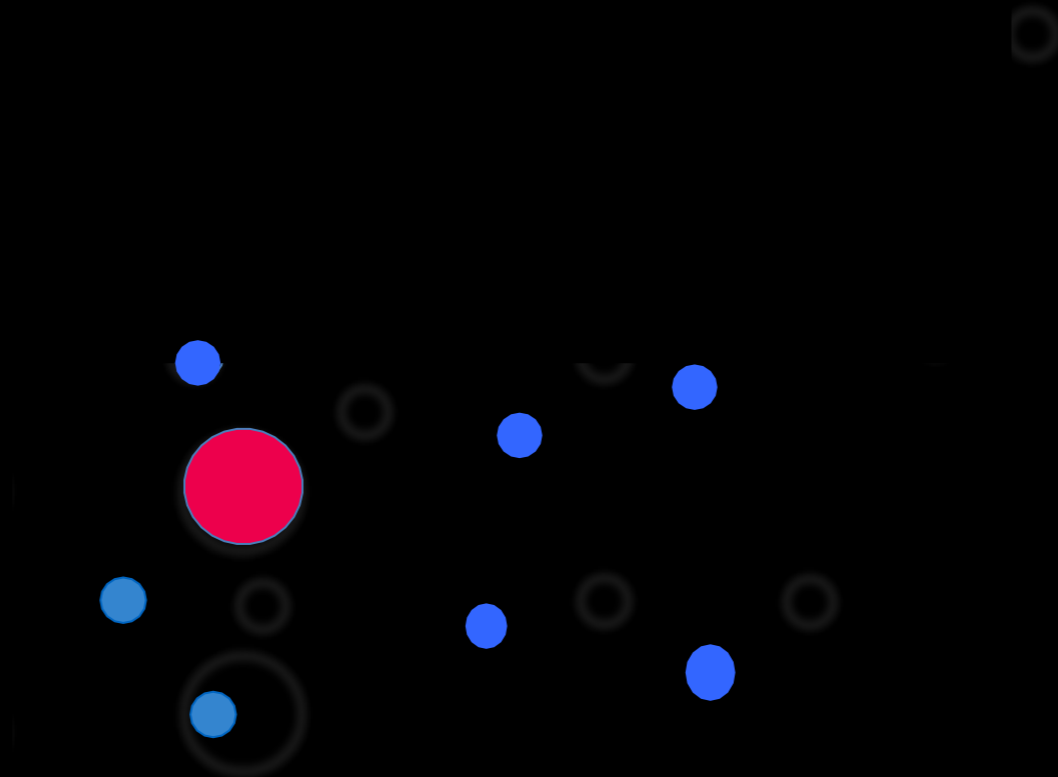




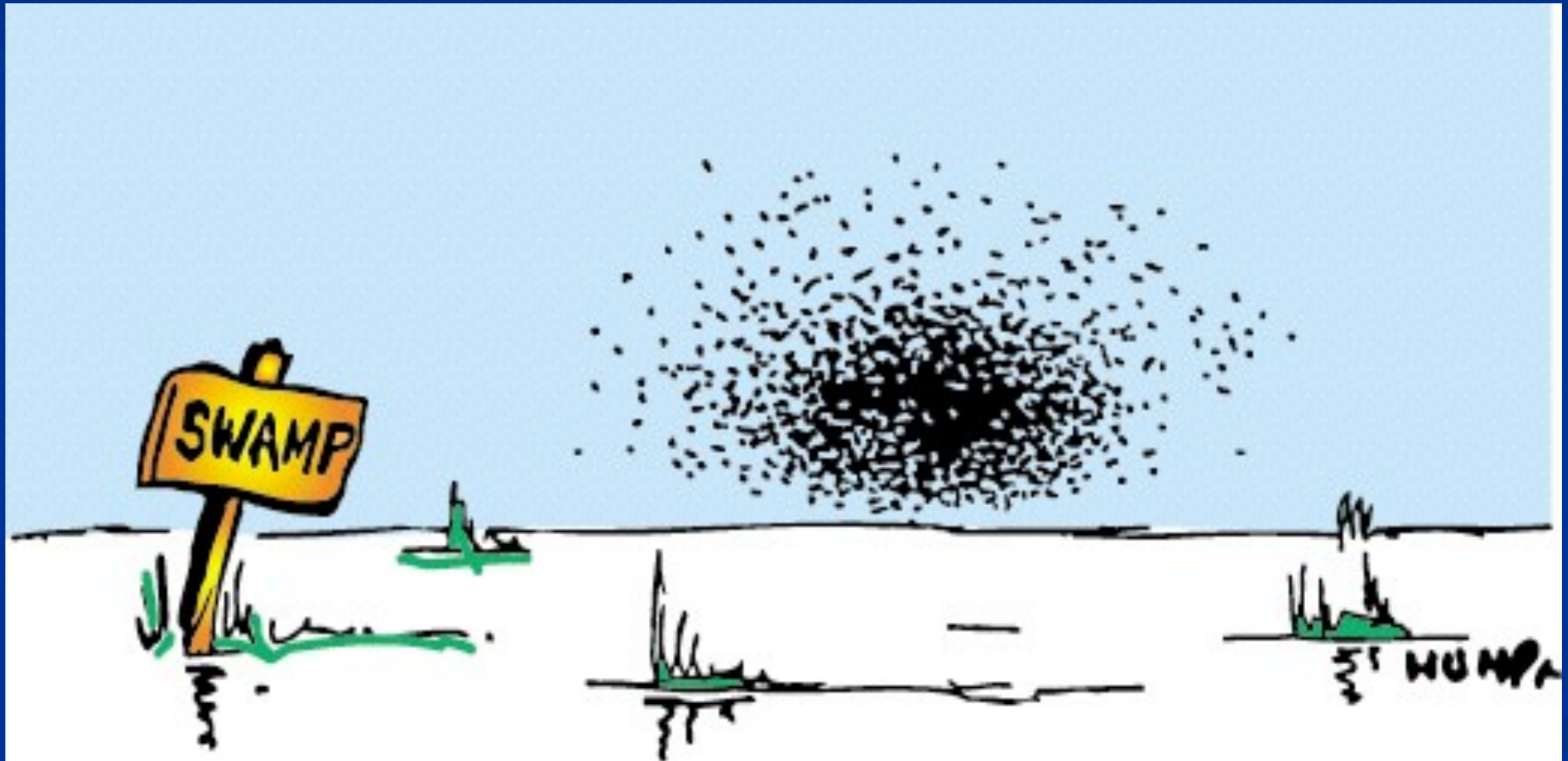
Positive Ions



Electrons



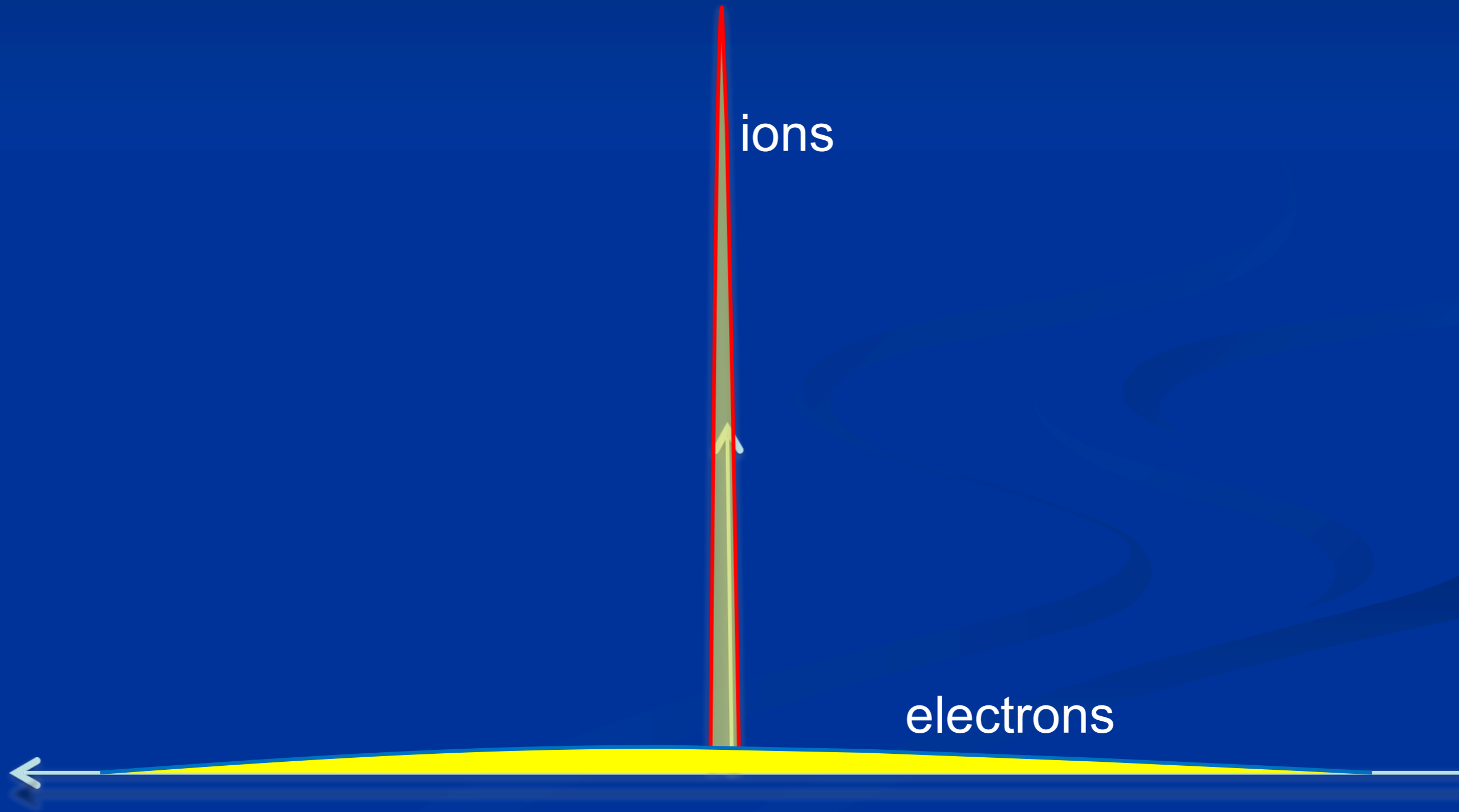
# Incoherent scattering: the short story



# Incoherent scattering: the short story



The ionospheric ions acts as  
sloooooow pacers for the electron gas



# Plasma Frequency ~

$$\omega_p^2 = \frac{n_0 e^2}{m \epsilon_0}$$



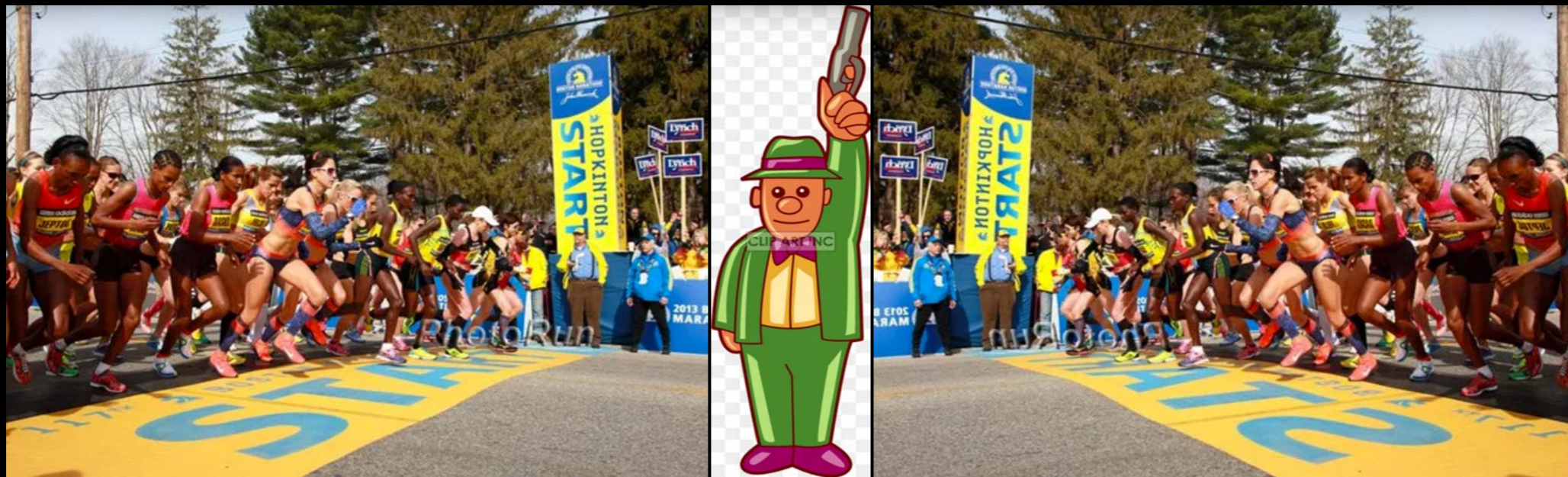
# Plasma Frequency ~

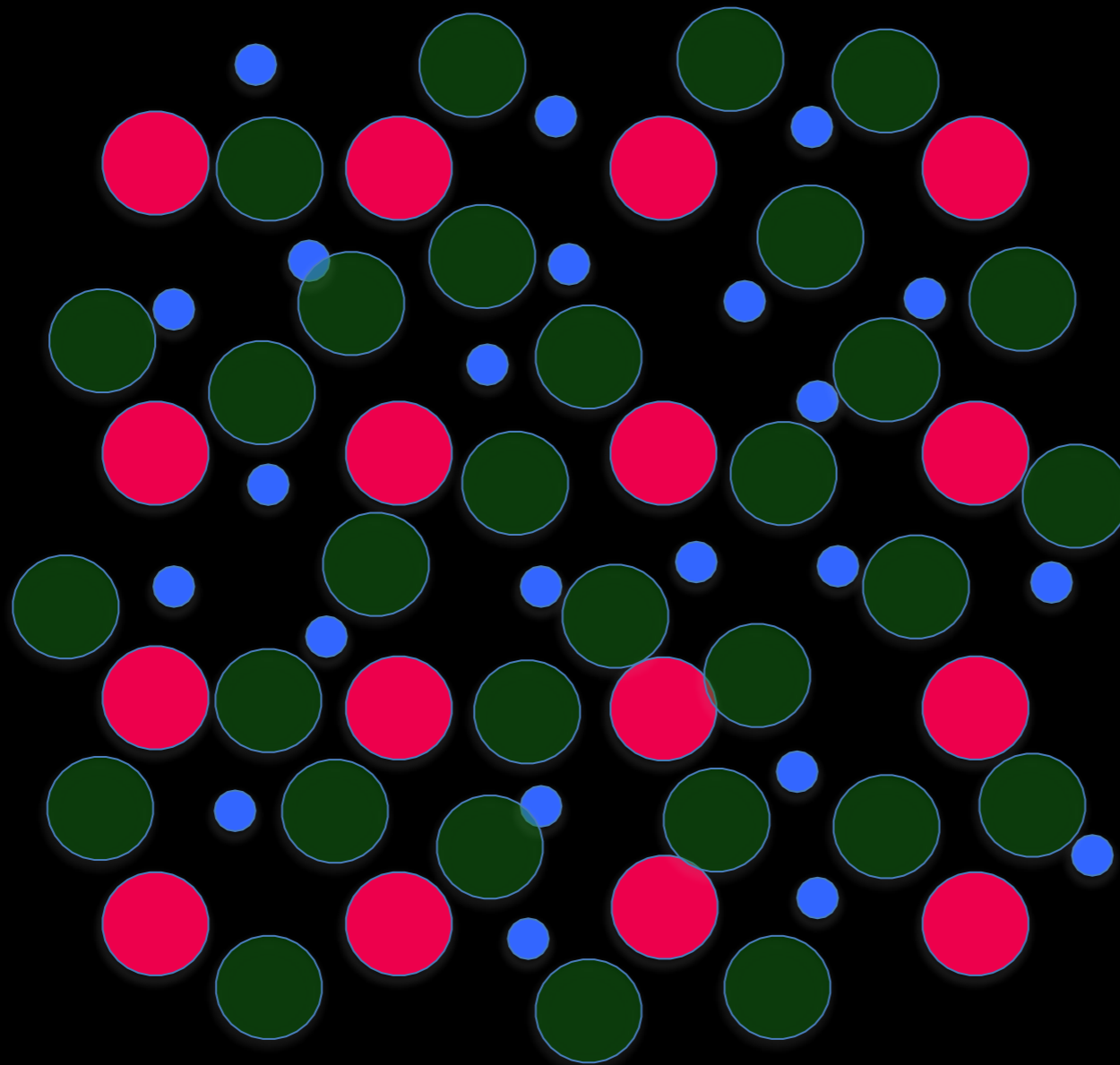
$$\omega_p^2 = \frac{n_0 e^2}{m \epsilon_0}$$



# Plasma Frequency ~

$$\omega_p^2 = \frac{n_0 e^2}{m \epsilon_0}$$





Neutrals

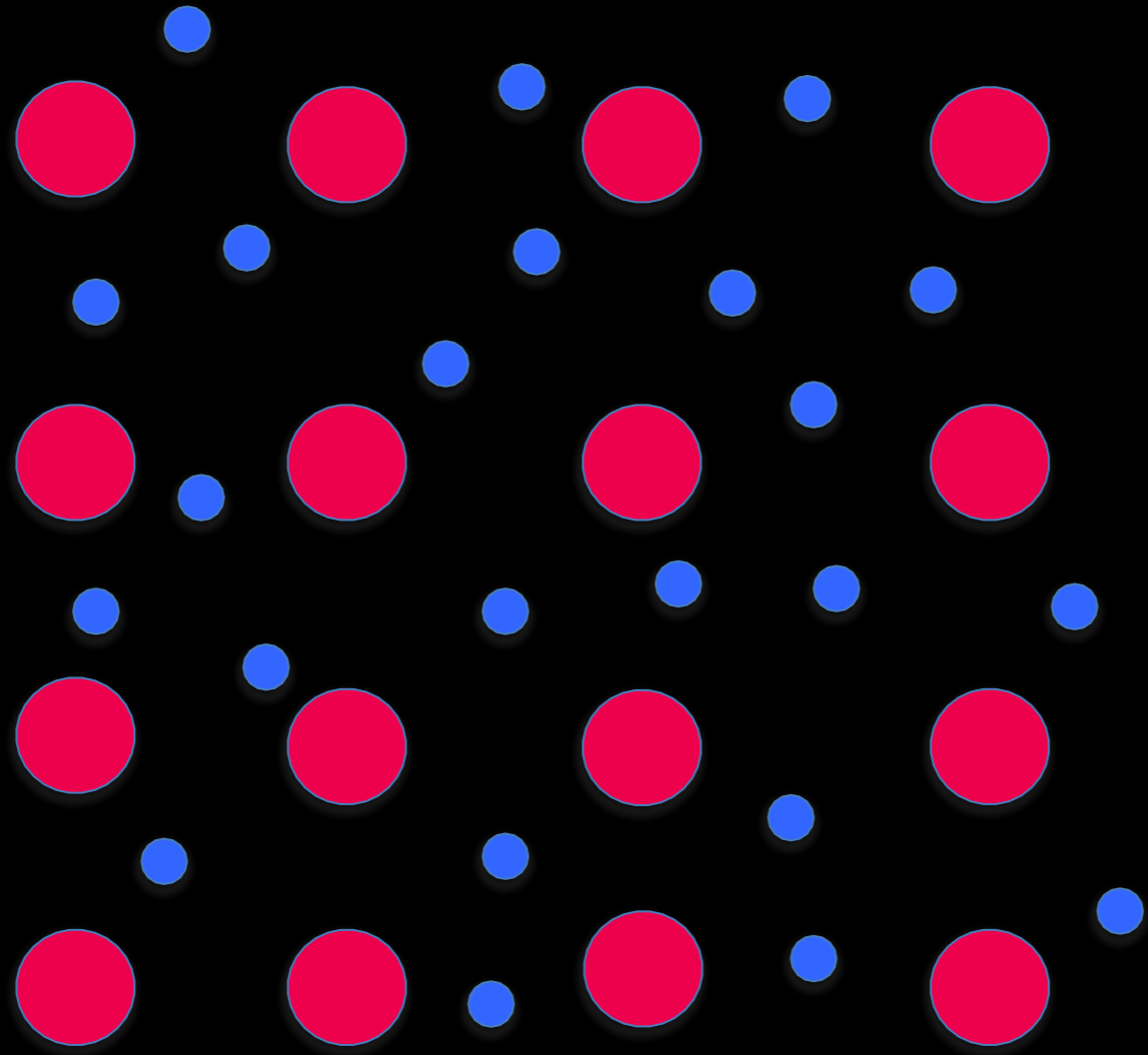


Positive Ions



Electrons

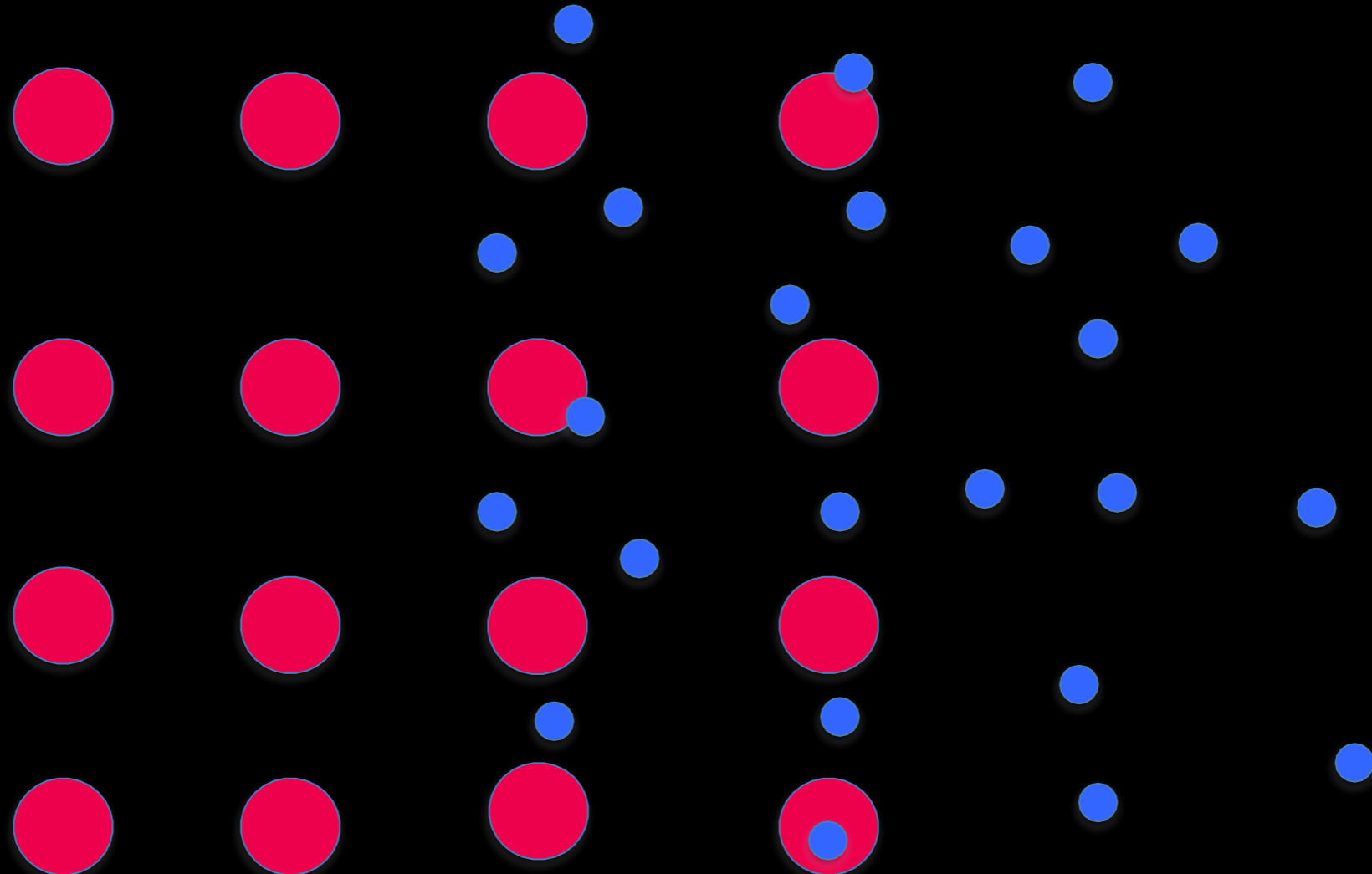




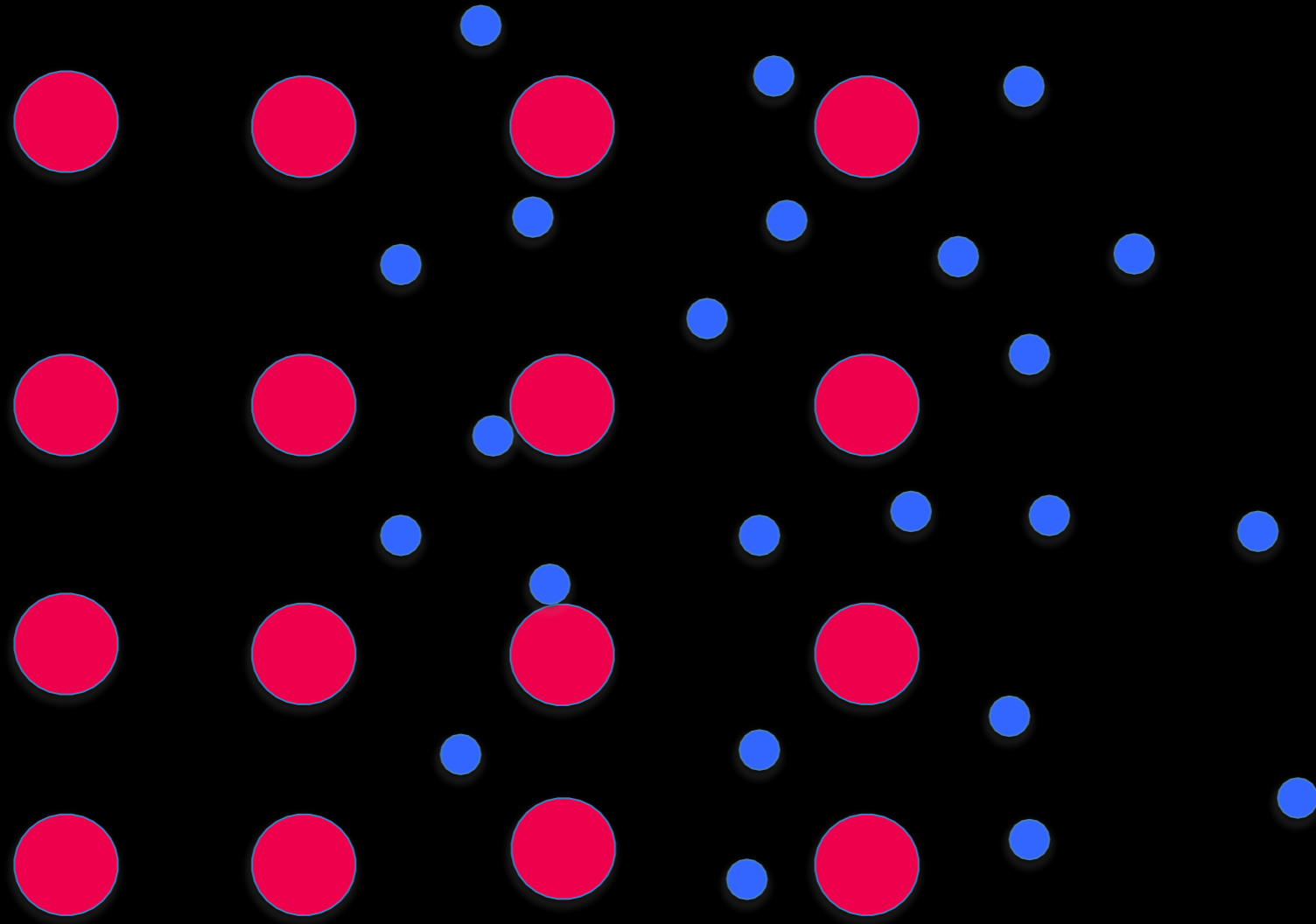
Positive Ions



Electrons



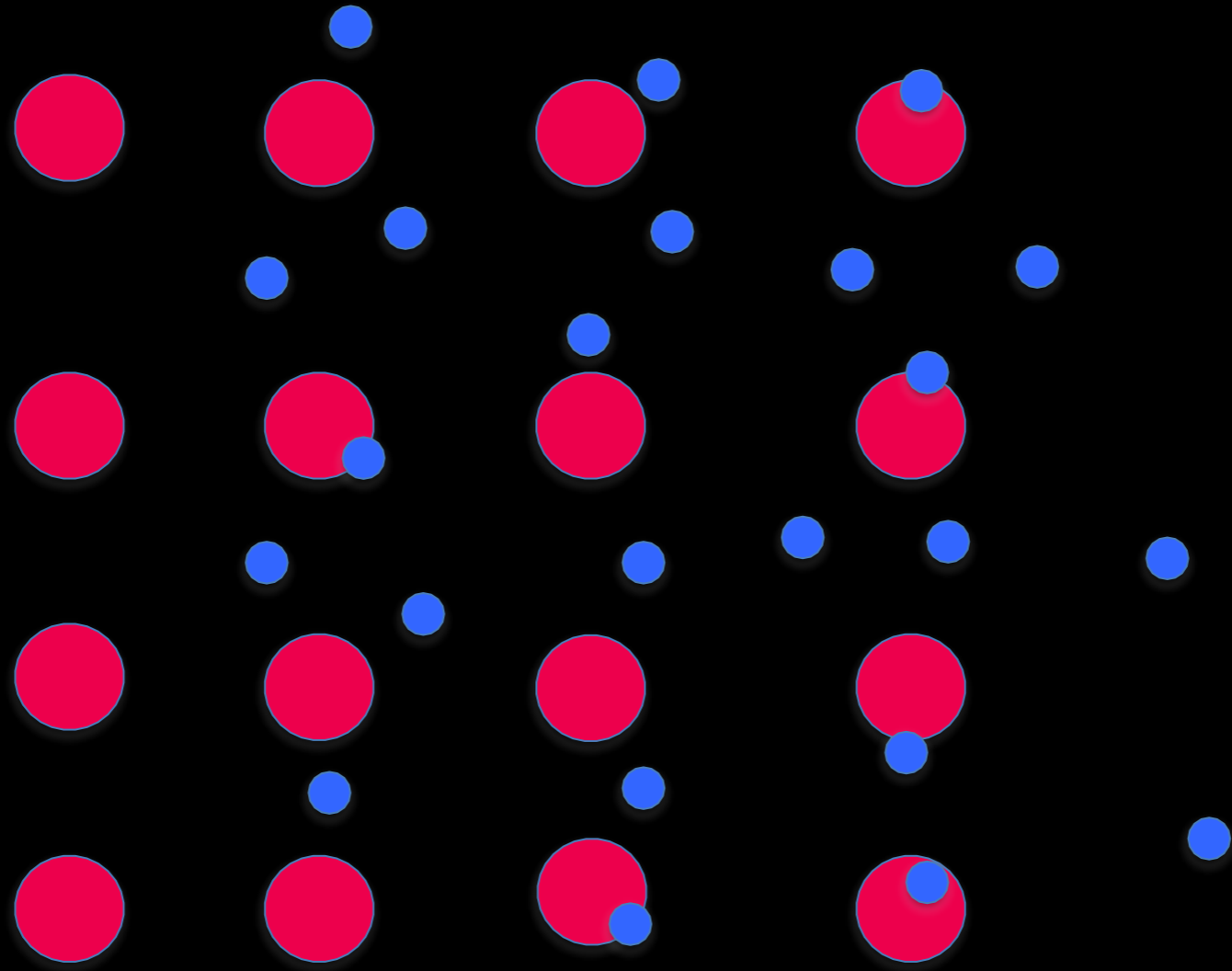
-  Positive Ions
-  Electrons



Positive Ions



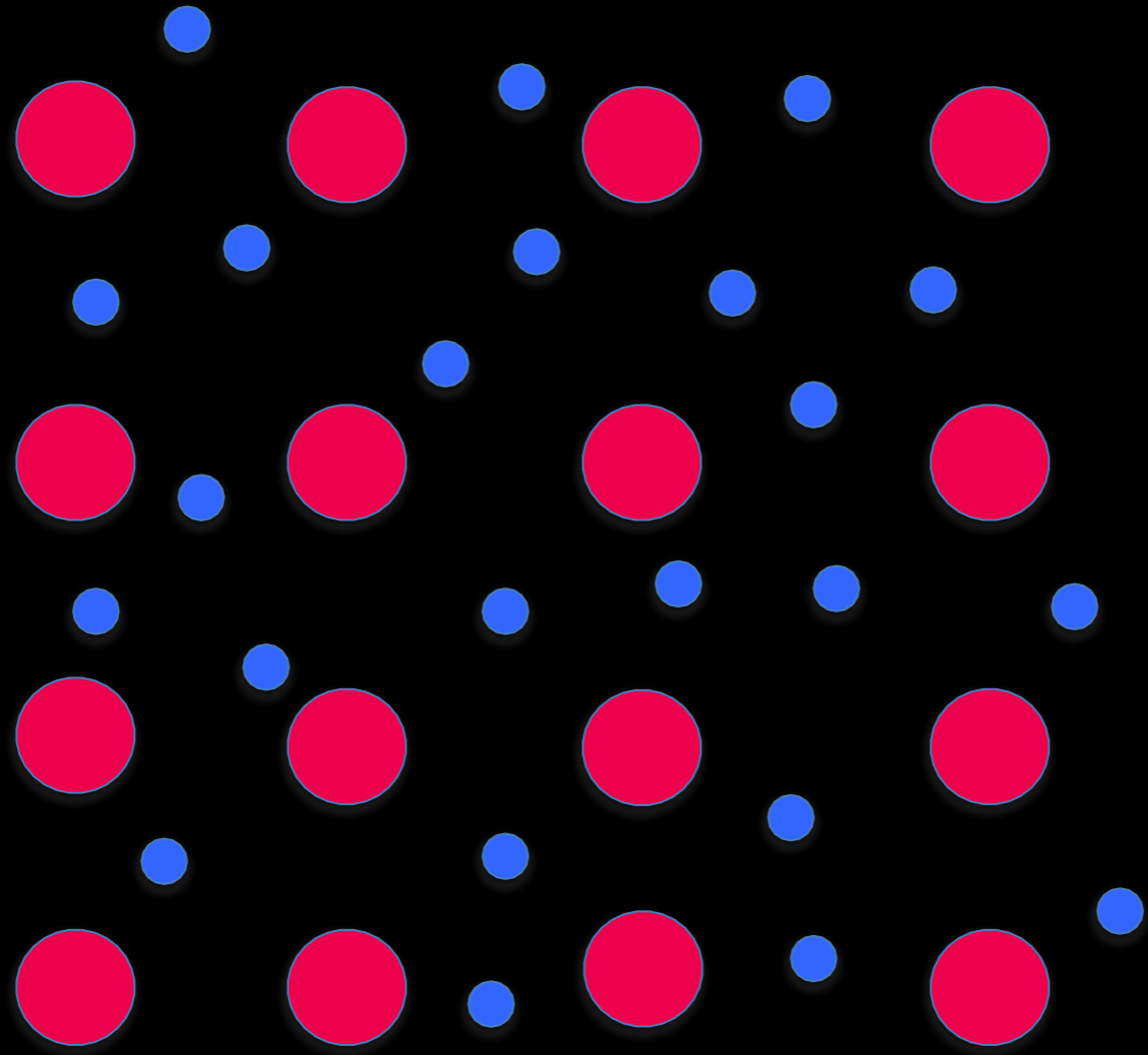
Electrons



Positive Ions



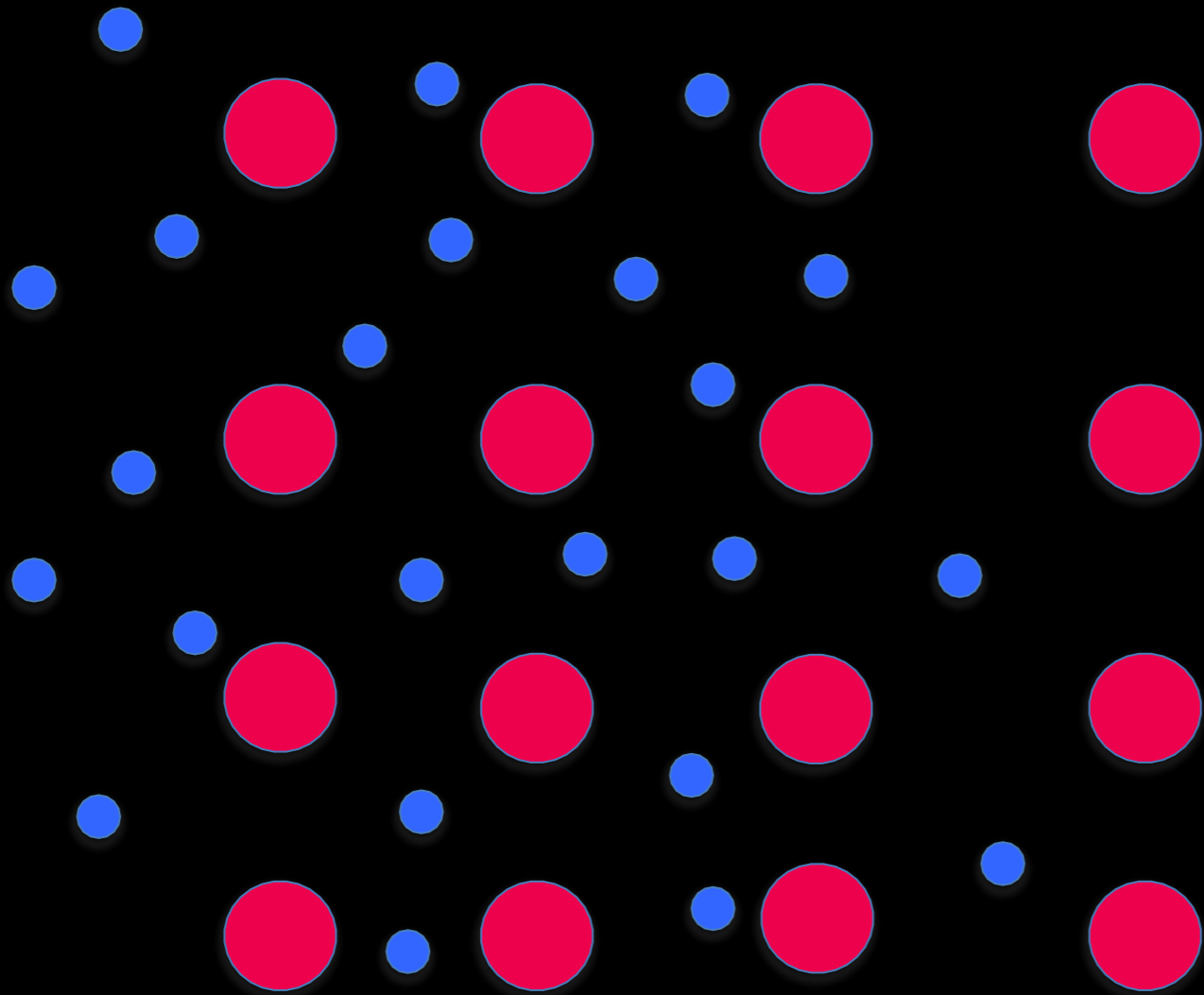
Electrons



Positive Ions



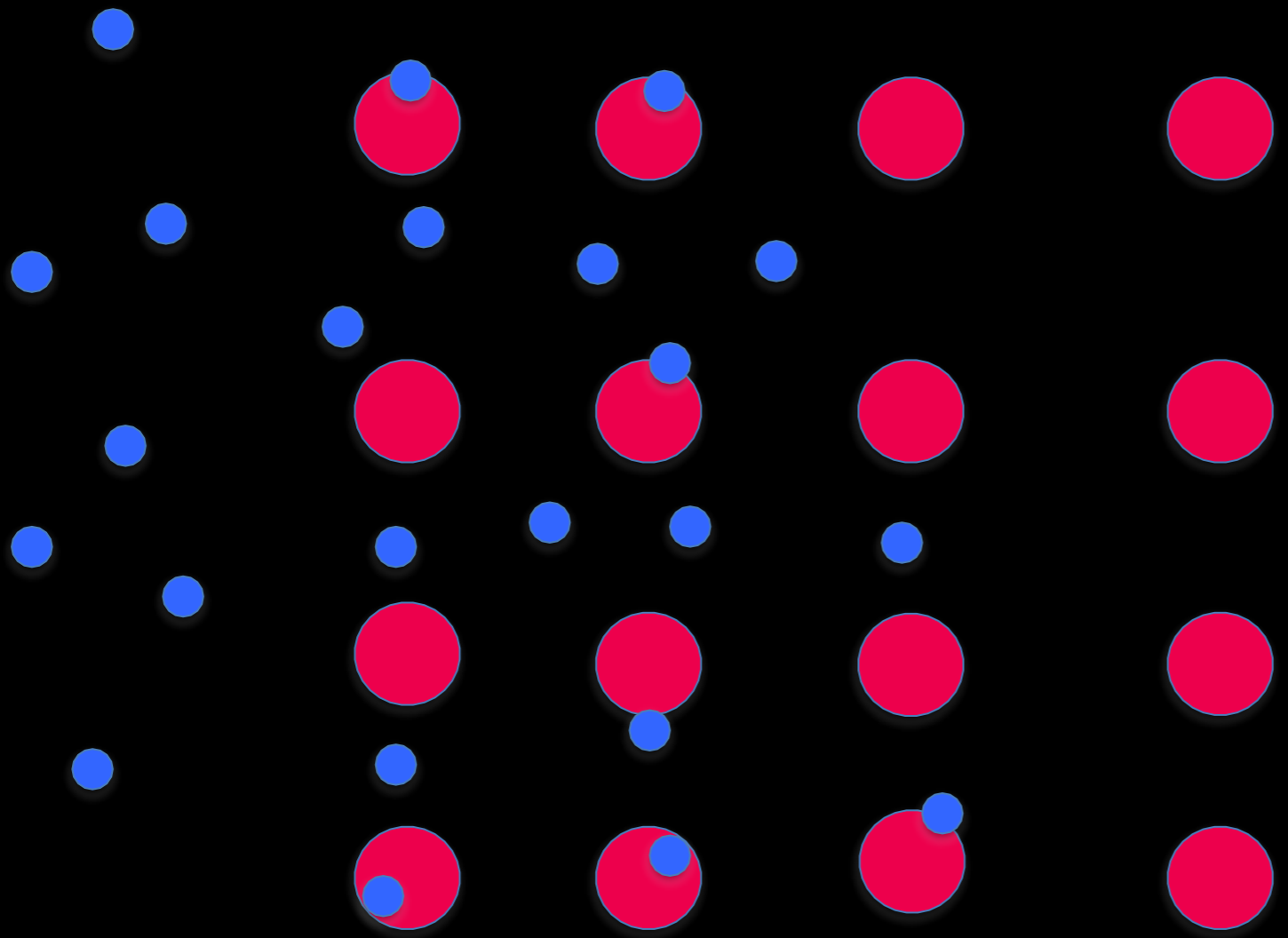
Electrons



Positive Ions



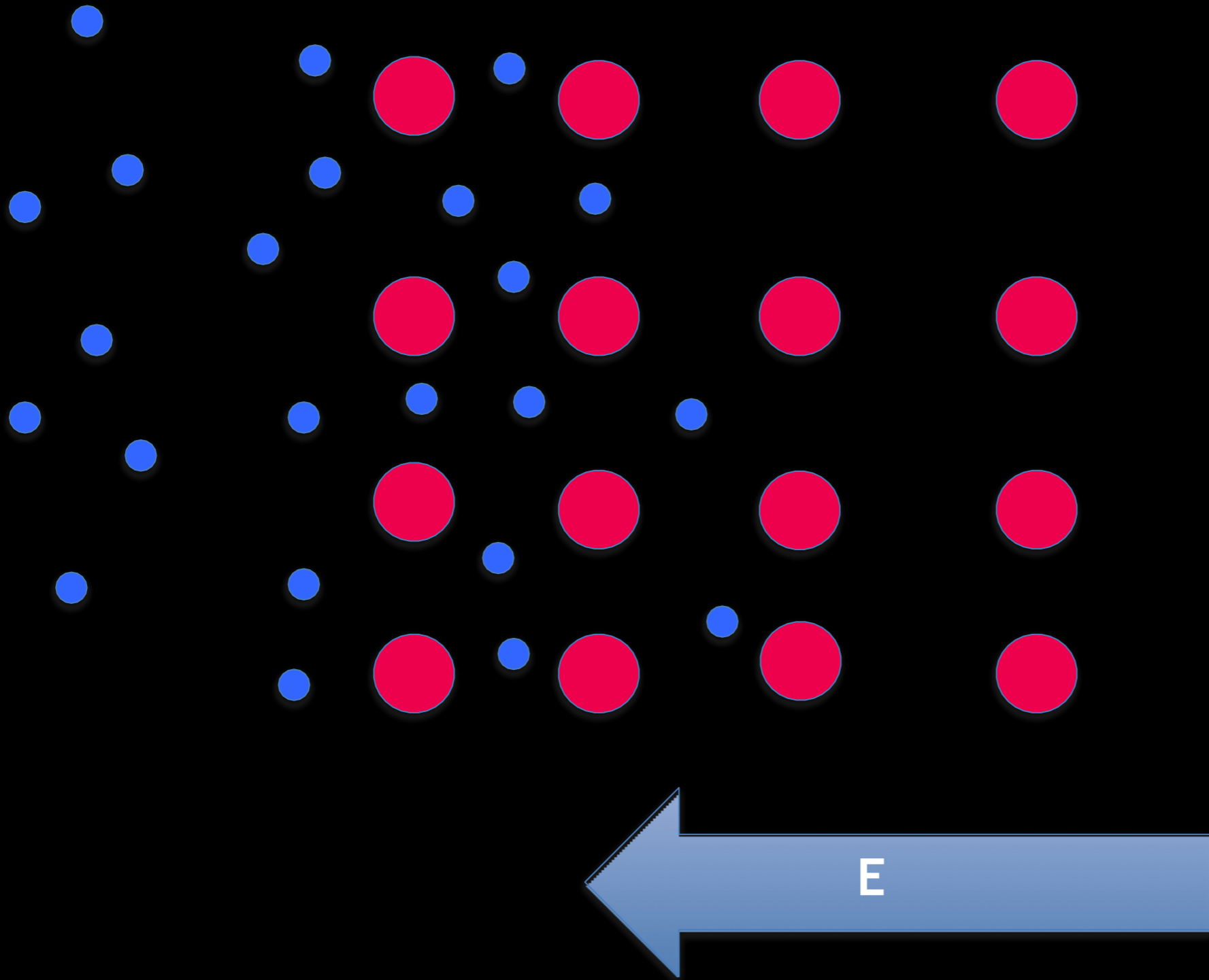
Electrons



Positive Ions



Electrons

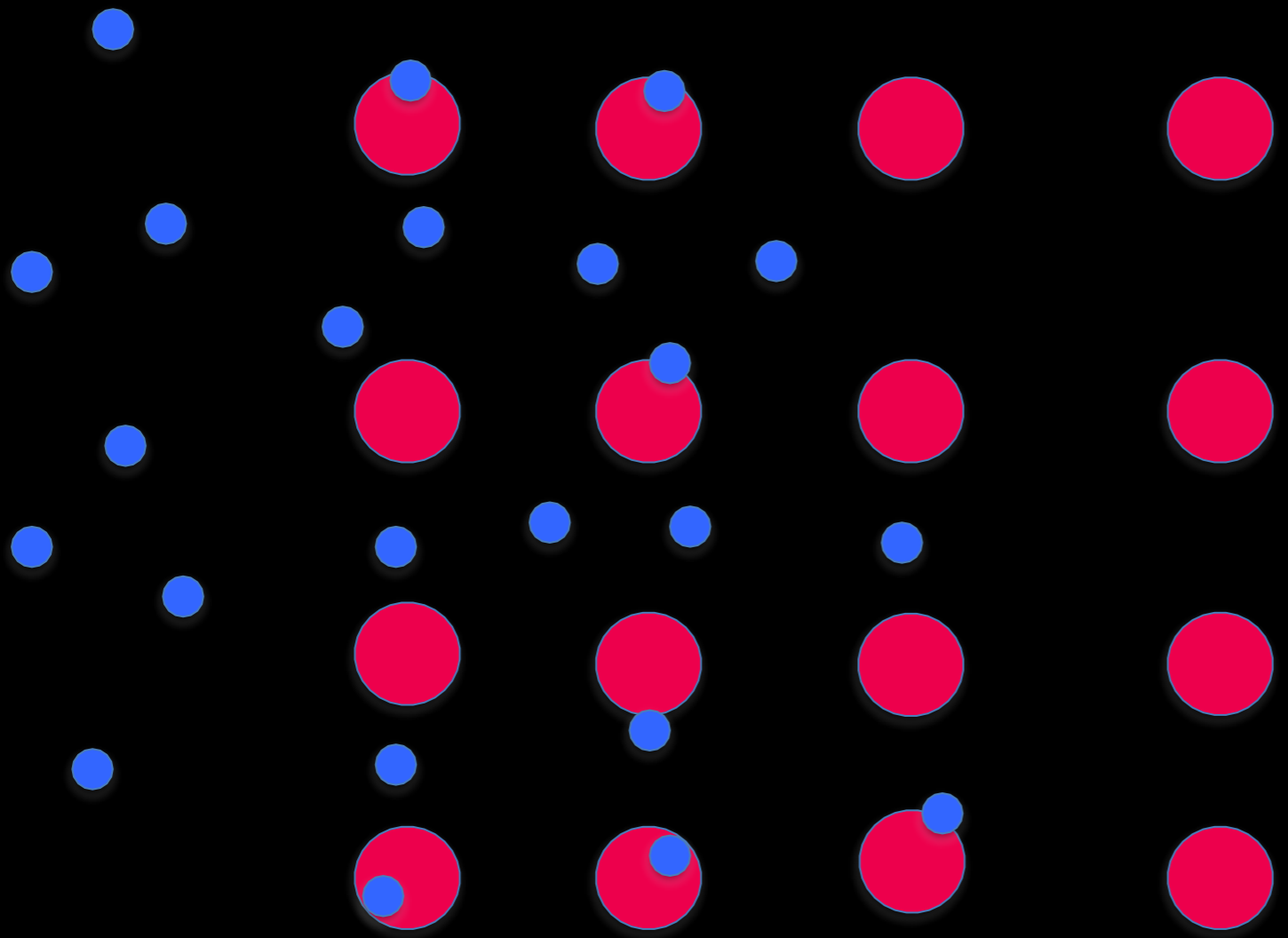


Positive Ions



Electrons

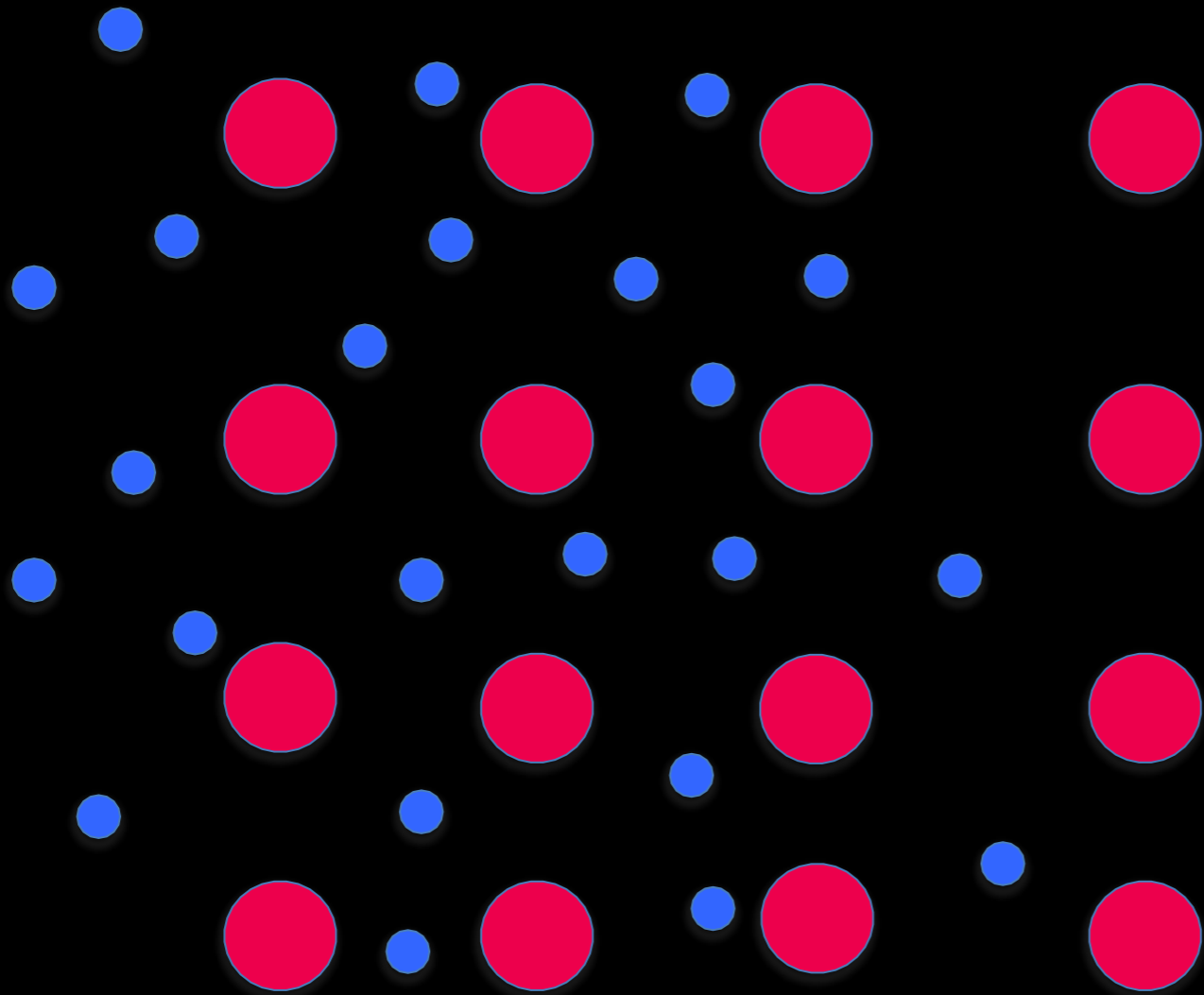




Positive Ions



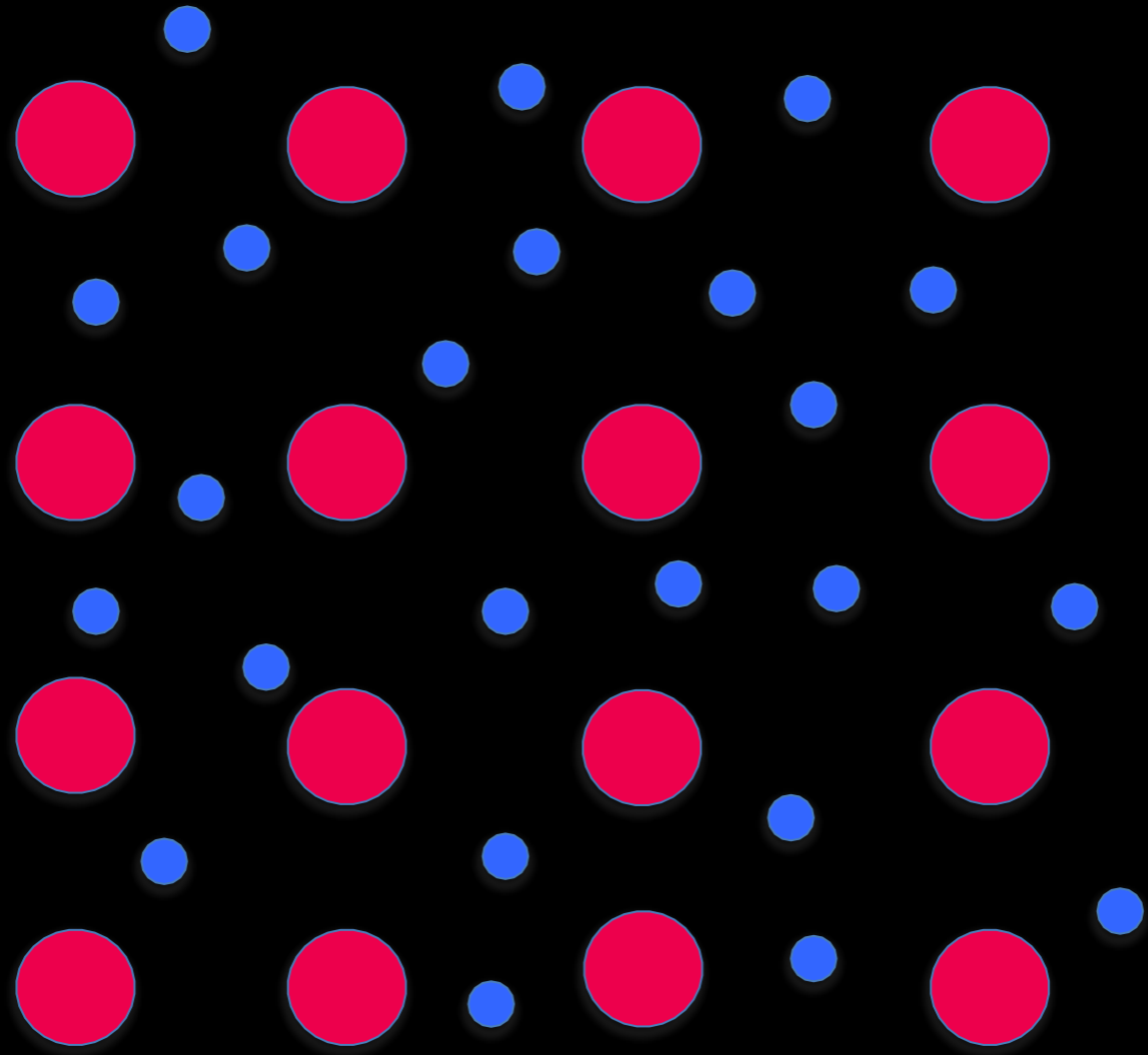
Electrons



Positive Ions



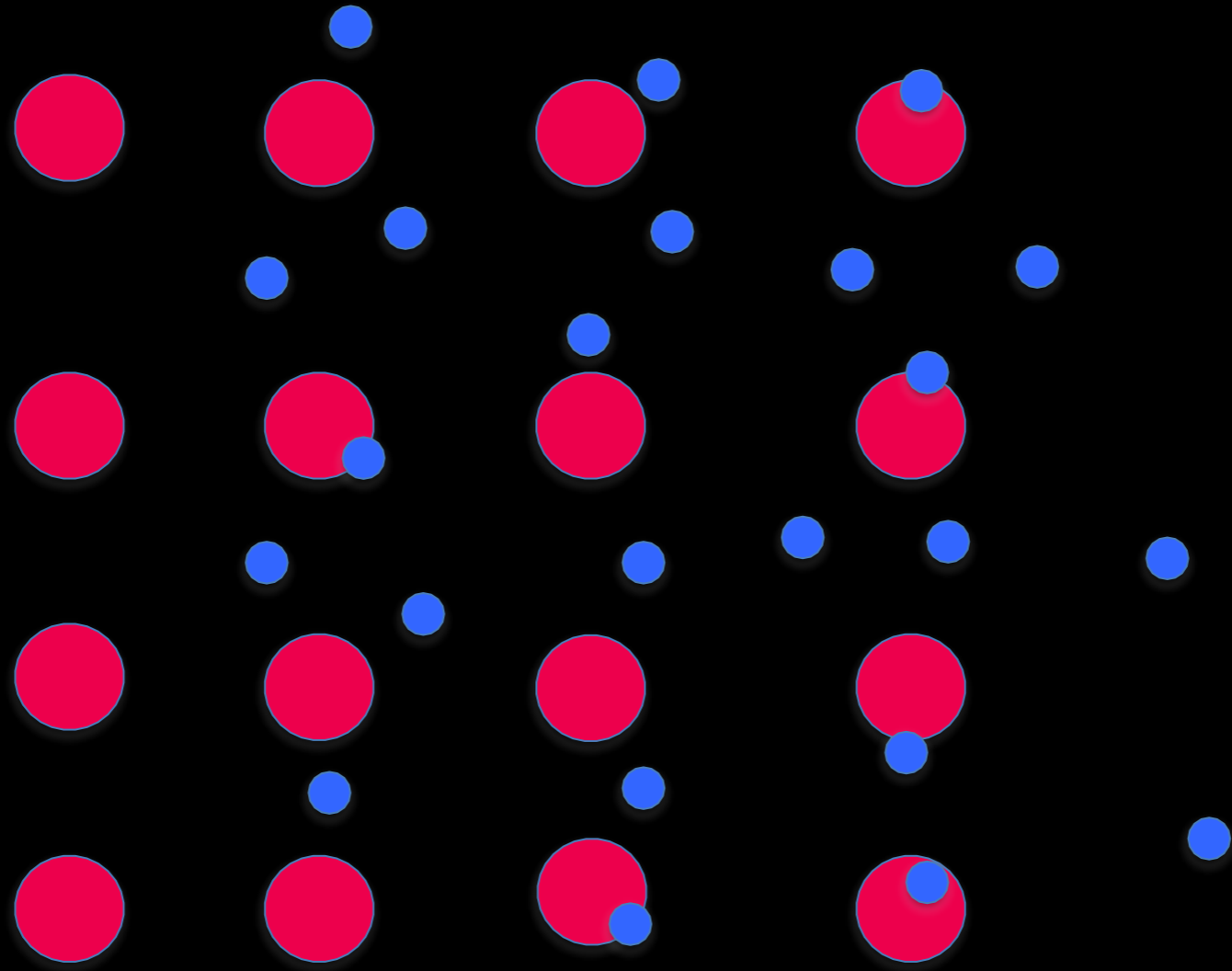
Electrons



Positive Ions



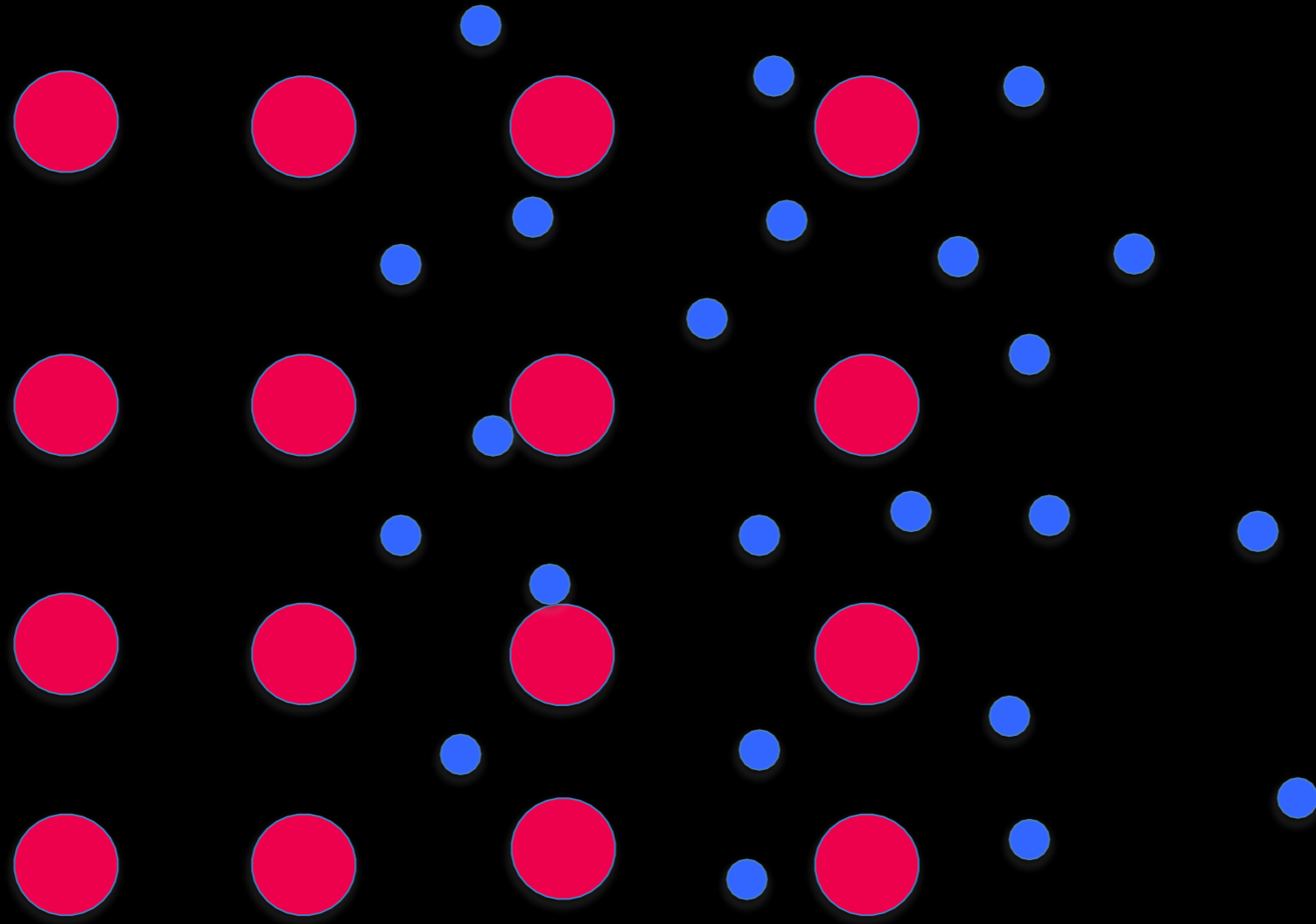
Electrons



Positive Ions



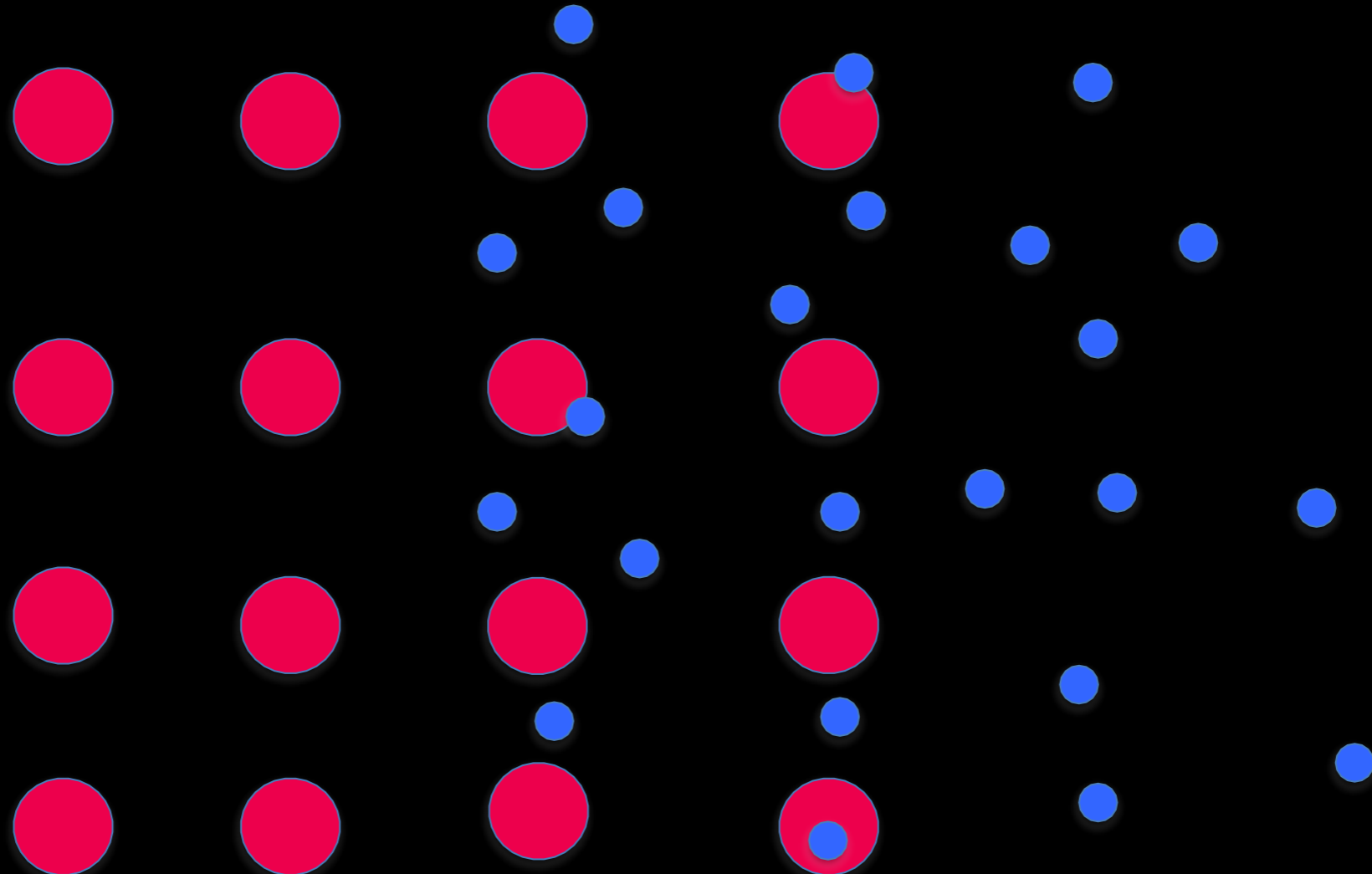
Electrons



Positive Ions



Electrons



Positive Ions



Electrons

The electron gas oscillates at a natural plasma oscillation frequency



# Langmuir (or plasma) waves

High-frequency electrostatic waves

- Dispersion relation:

$$\omega_r = (\omega_{pe}^2 + 3k^2 v_{the}^2)^{1/2} = \omega_p (1 + 3k^2 \lambda_{De}^2)^{1/2}$$

$$\omega_i = -c \frac{\omega_{pe}}{(k\lambda_{De})^3} \exp(-\frac{1}{2} k^2 \lambda_{De}^2), \quad c = \sqrt{\frac{\pi}{8}} e^{-3/2}$$

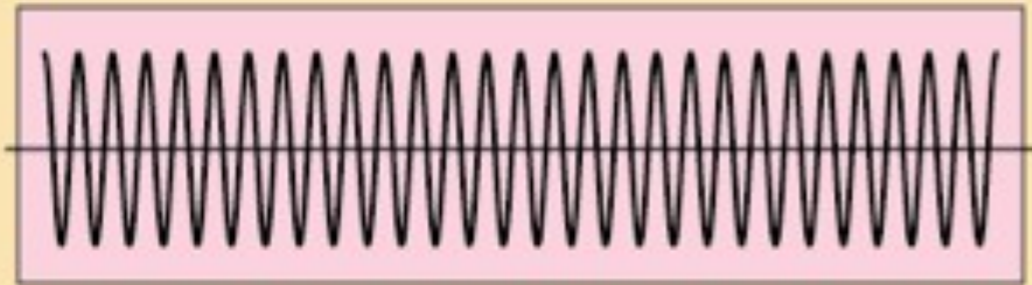


# Ion acoustic waves

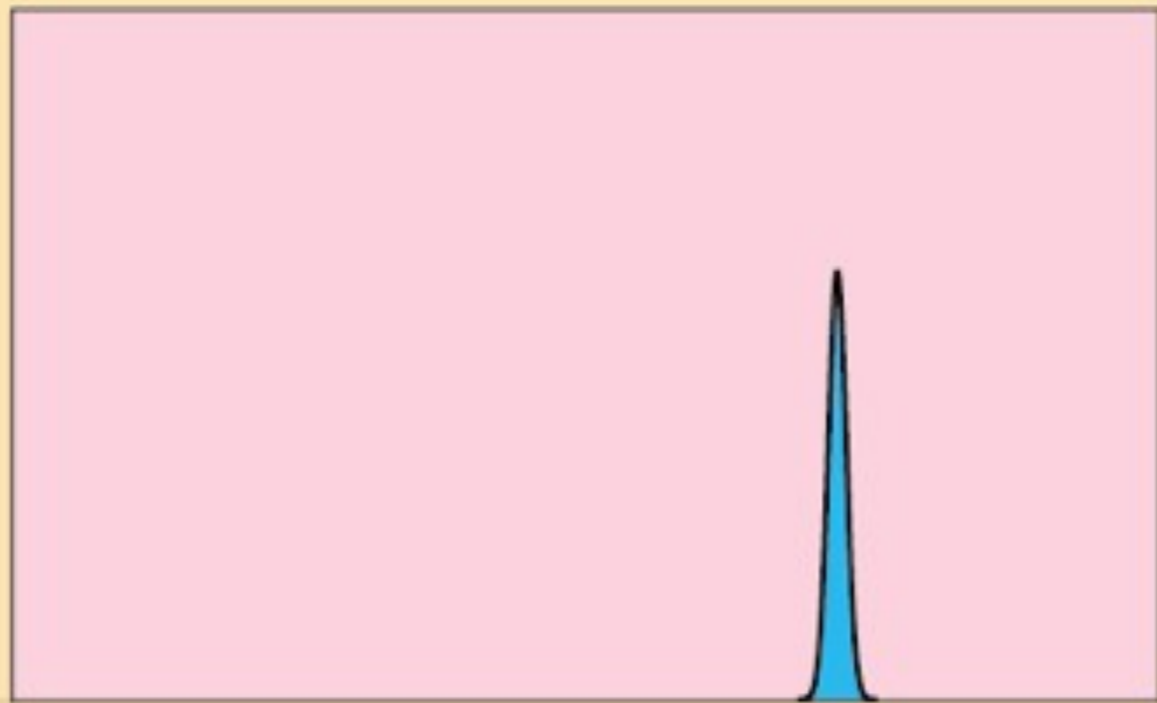
\* Ion acoustic waves:

$$(63) \quad \omega_r = \frac{u c_s}{1 + u^2 \lambda_{De}^2}, \quad c_s = \left( \frac{k_B T_e + 3k_B T_i}{m_i} \right)^{1/2}$$

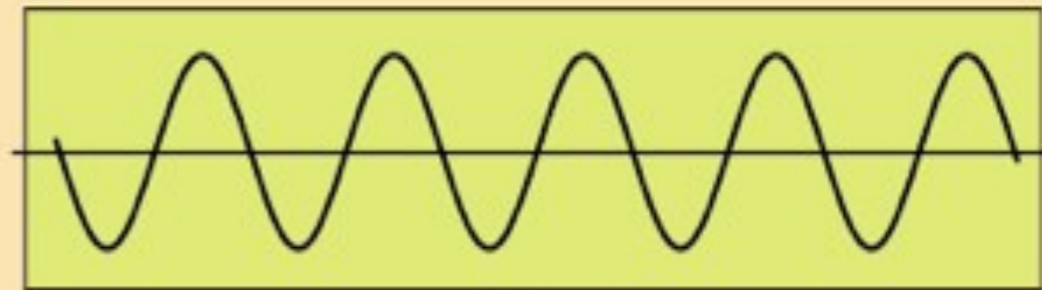
$$(64) \quad \omega_i = -\sqrt{\frac{\pi}{8}} \frac{\omega_r}{(1 + u^2 \lambda_{De}^2)^{3/2}} \left[ \left( \frac{T_e}{T_i} \right)^{3/2} \exp \left( -\frac{T_e/T_i}{2(1 + u^2 \lambda_{De}^2)} + \sqrt{\frac{m_e}{m_i}} \right) \right]$$



time



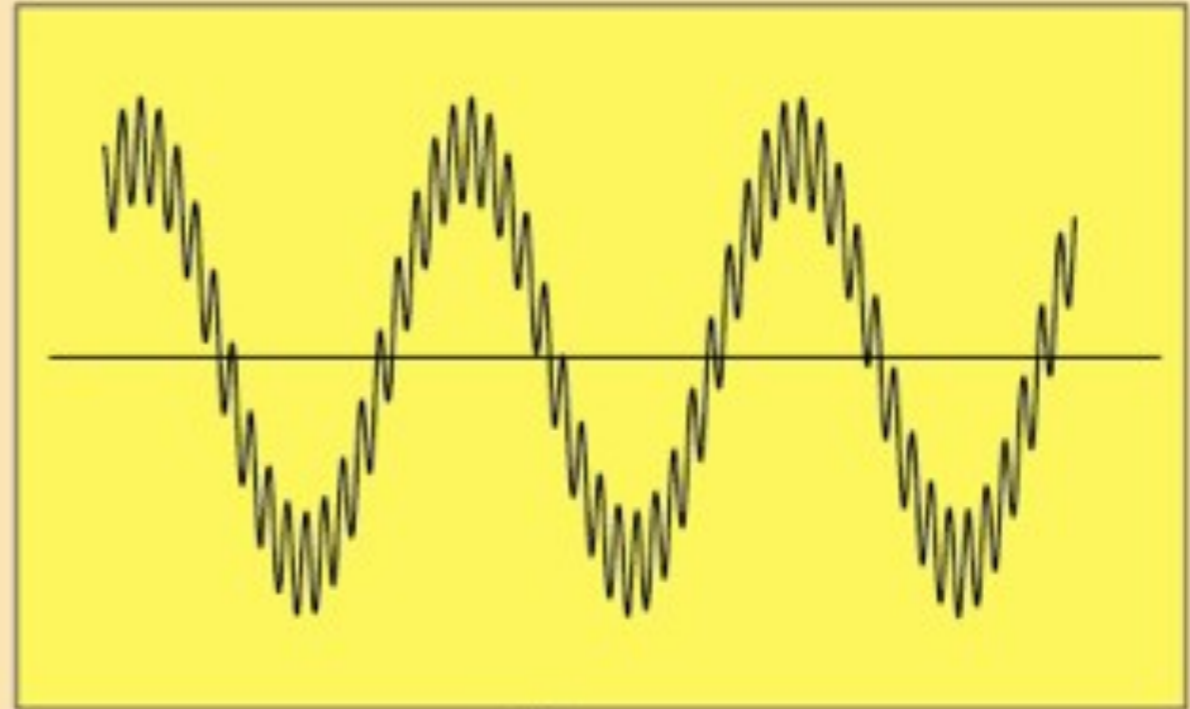
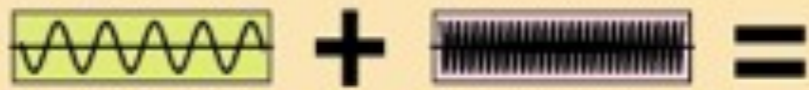
frequency



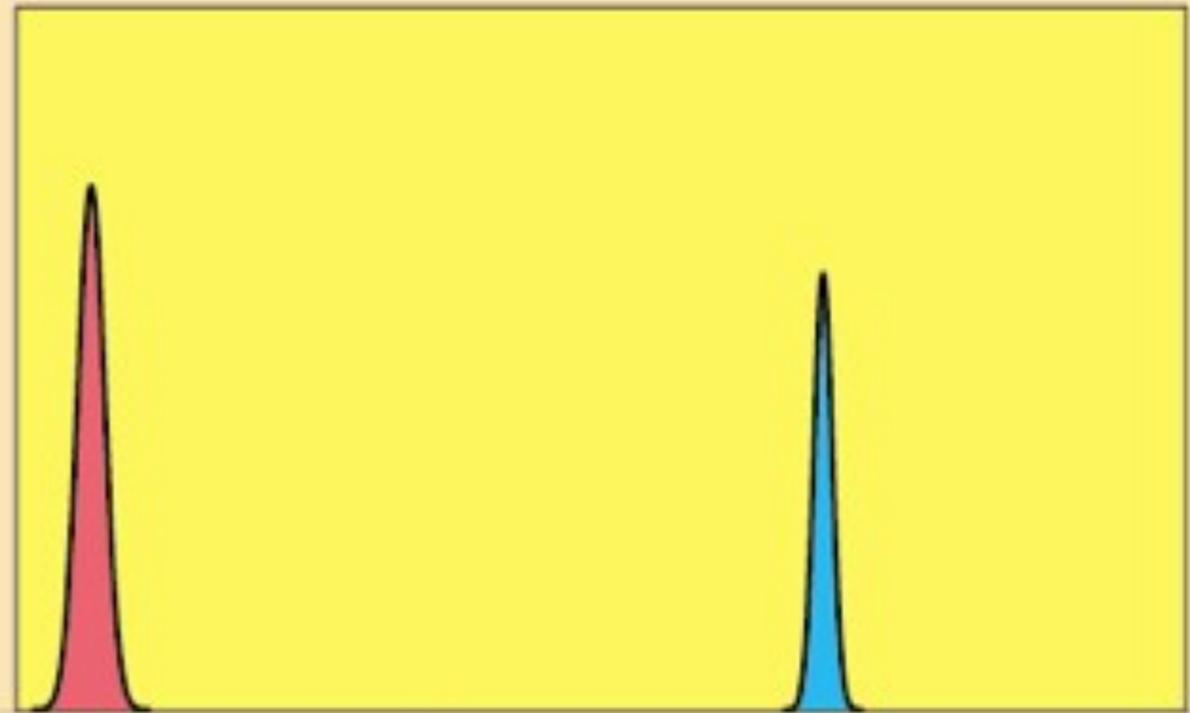
time



frequency

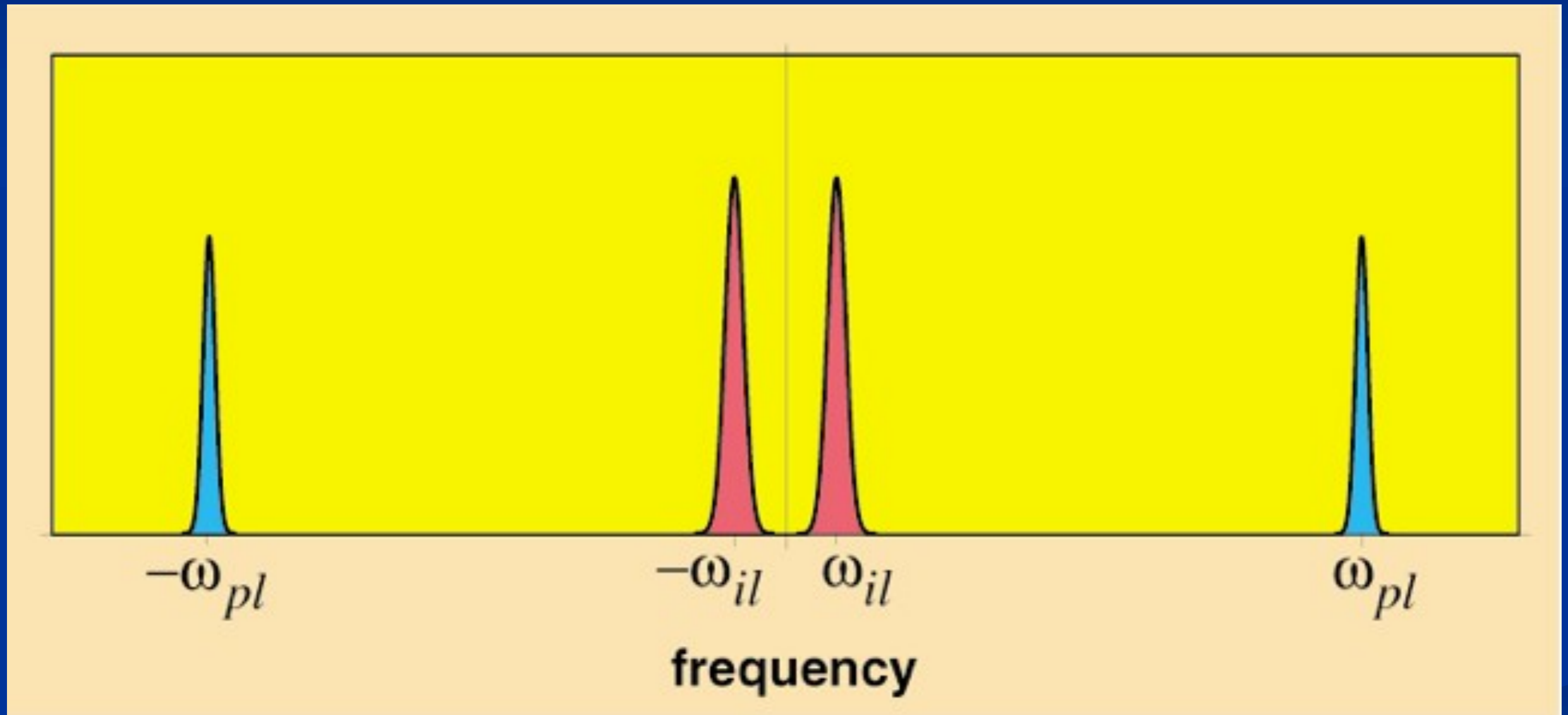


time

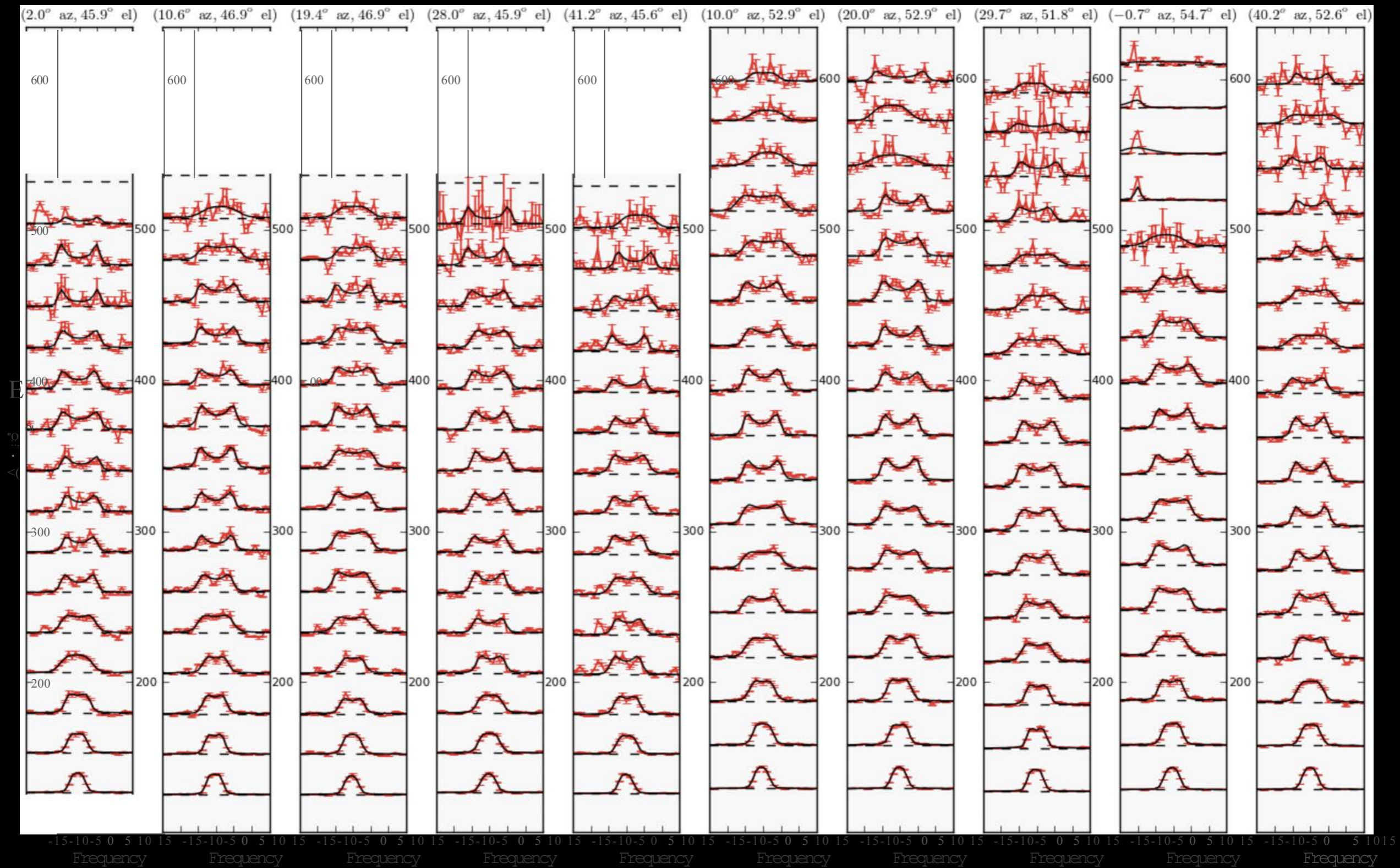


frequency

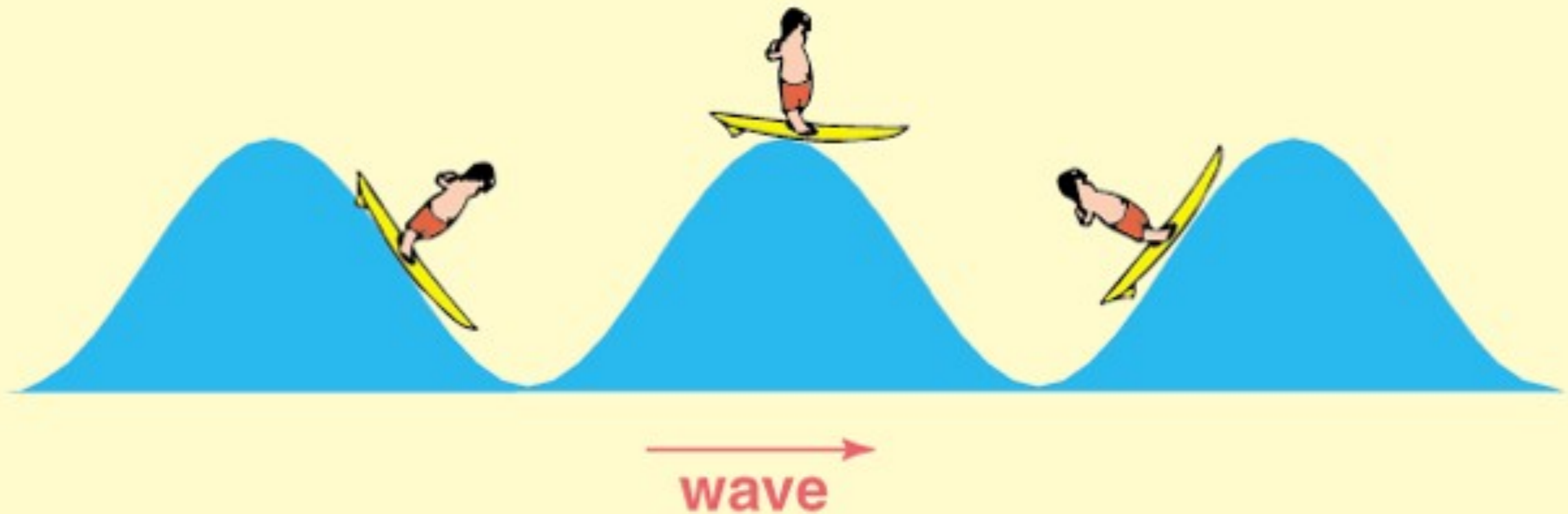
# Plasma Wave Approach (cont'd)



8-1-2012 9.520-9.538 UT



# Landau wave-particle interactions



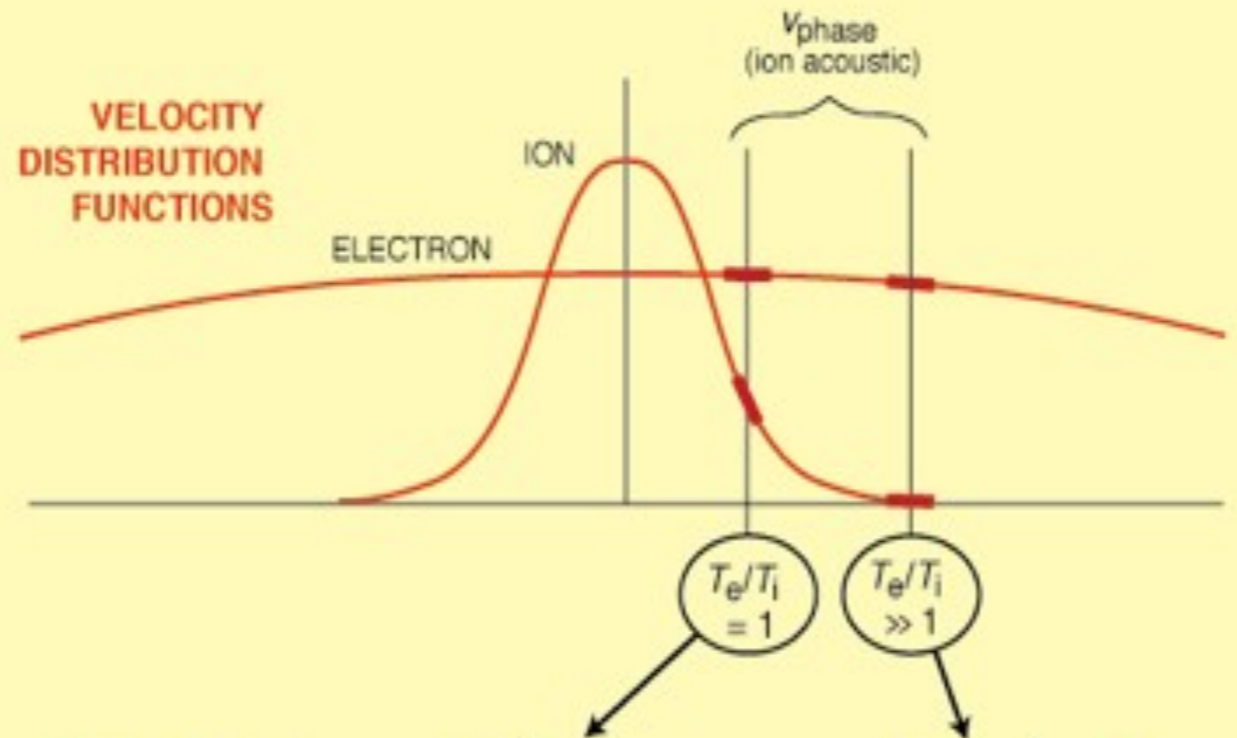
particle  
gains  
energy

wave  
gains  
energy

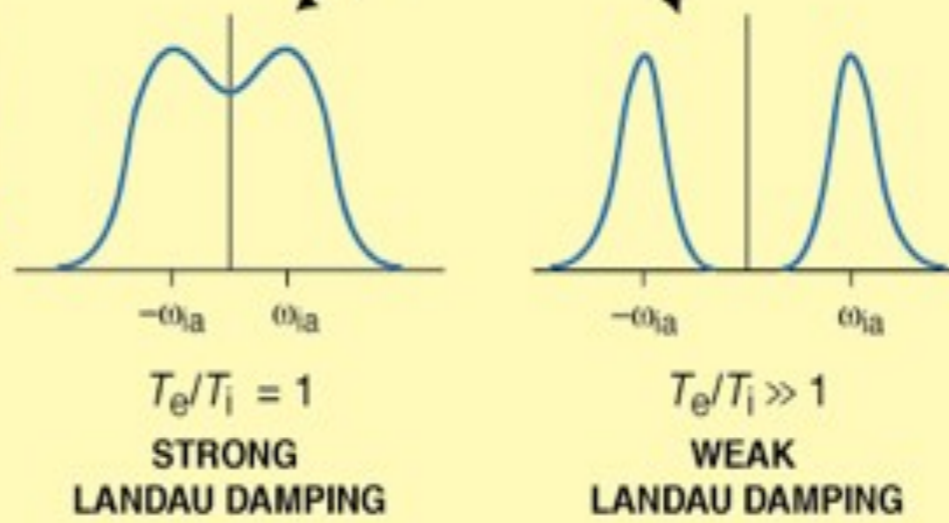
# THE EFFECT OF LANDAU DAMPING ON THE INCOHERENT SCATTER ION LINE SPECTRUM

**ION-ACOUSTIC  
DISPERSION  
EQUATION**

$$\omega_{ia} = k v_{\text{phase}} = k \left( \frac{T_e + 3T_i}{m_i} \right)^{1/2}$$

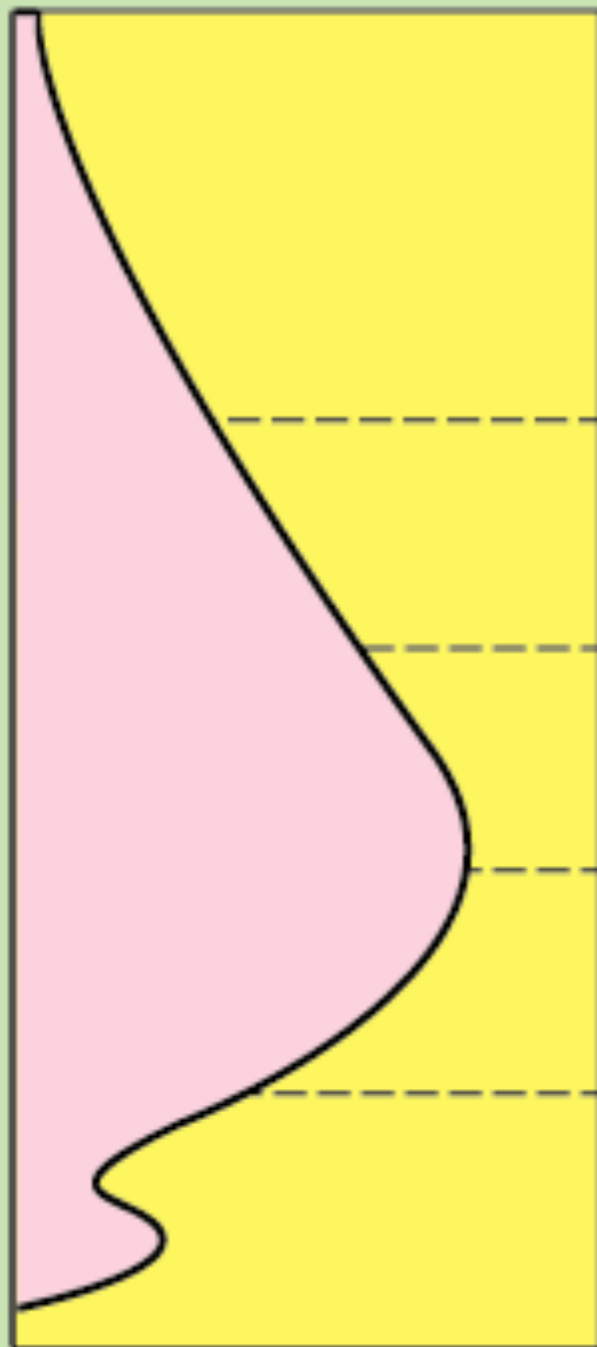


**INCOHERENT  
SCATTER  
ION LINE  
SPECTRA**

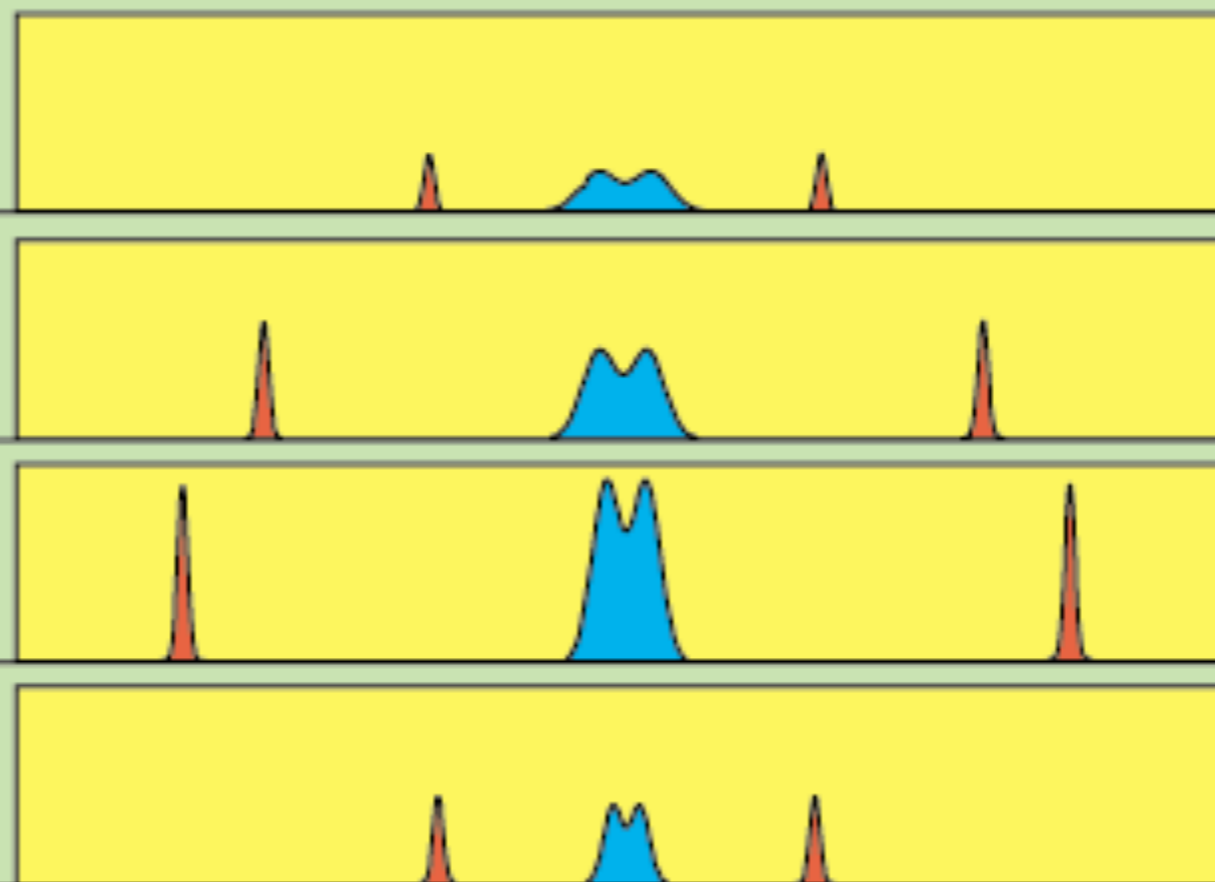


# electron density profile

height



# Incoherent scatter spectra



spectral amplitude

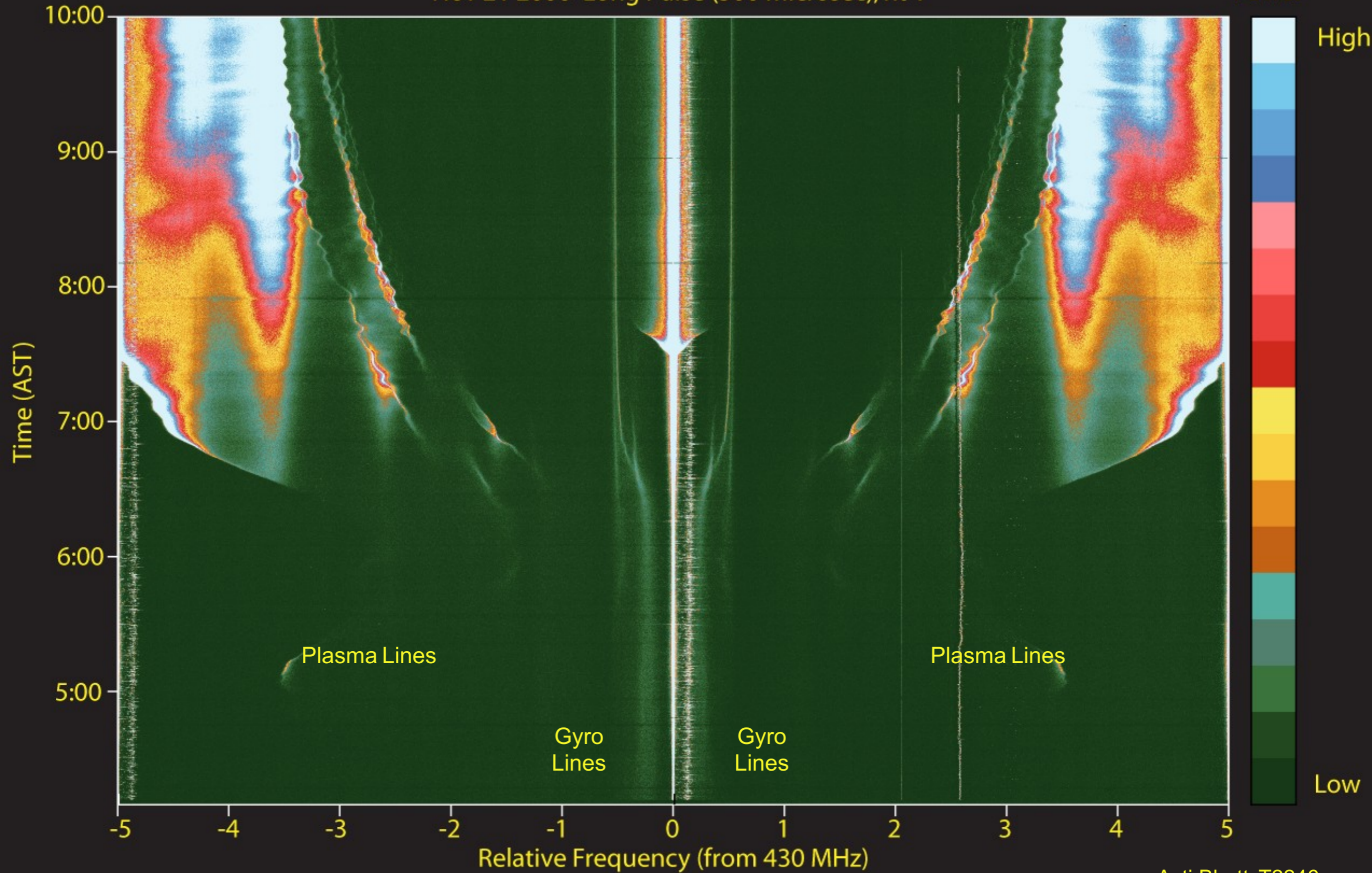


← frequency →



Nov 21 2006 Long Pulse (500 microsec), ht 1

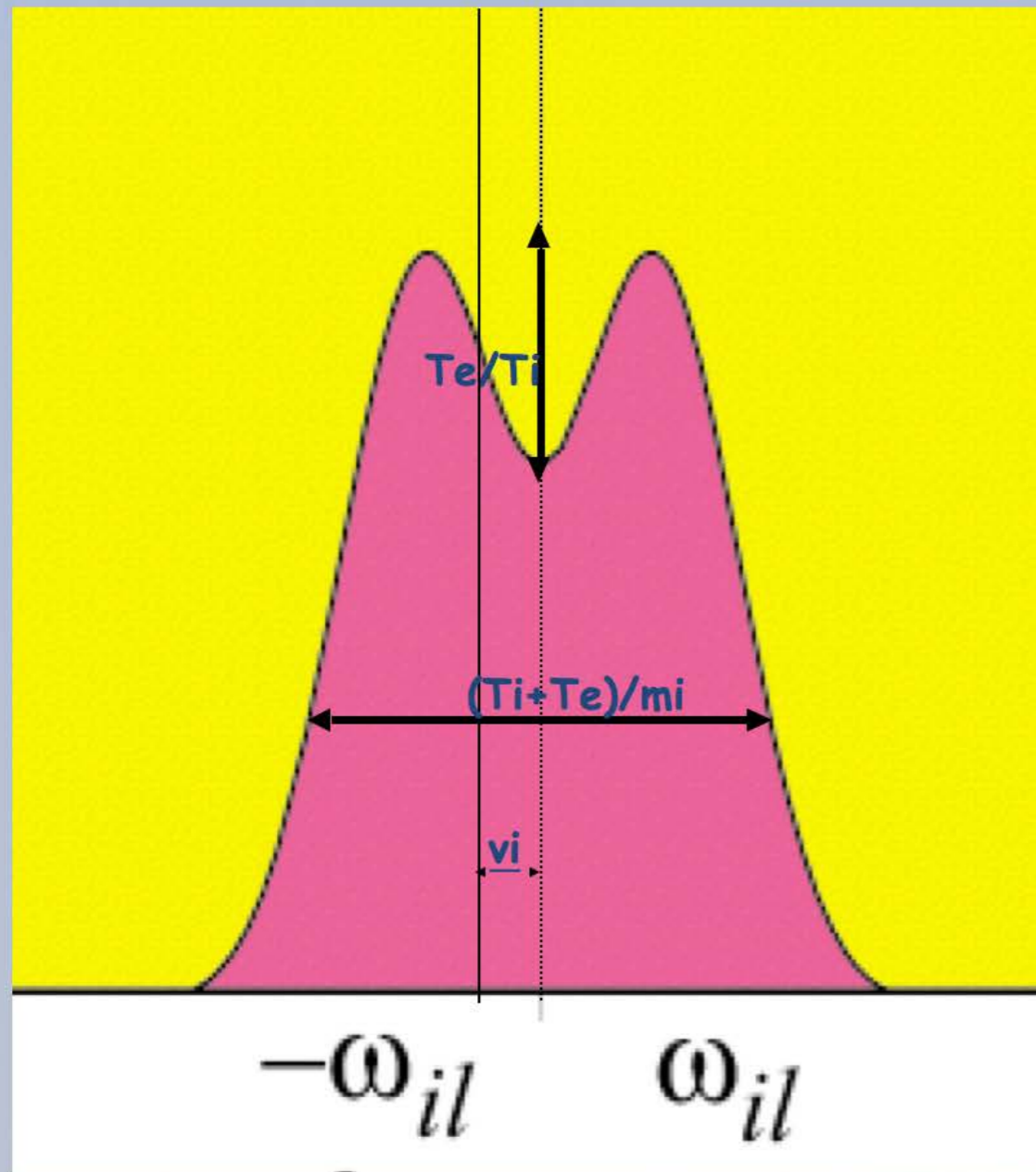
Power



Asti Bhatt, T2246

Arecibo Sensitivity: The 305 m dish, 2.5 MW of power, and  $T_{\text{sys}}$  of about 80 K (condition dependent) provide high time resolution on even weak features such as the gyro line. The data above are centered on the E region. The strong plasma line after sunrise is “leakage” from the low F region. The complicated behavior of the gyro line is probably due to multiple layers, but is not completely understood.!

...or to sum up...



- Ion (and electron) temperature ( $T_i$  and  $T_e$ ) to ion mass ( $m_i$ ) ratio from the width of the spectra
- Electron to ion temperature ratio ( $T_e/T_i$ ) from “peak\_to\_valley” ratio
- Electron (= ion) density from total area (corrected for temperatures)
- Ion velocity ( $v_i$ ) from the Doppler shift

Plasma Line  $S_{PL}(\mathbf{k}, \omega)$

Ion Line  $S_{IL}(\mathbf{k}, \omega)$

$$S_e(\mathbf{k}, \omega) = N_e \left| 1 - \frac{\chi_e(\mathbf{k}, \omega)}{\epsilon(\mathbf{k}, \omega)} \right|^2 \int d\mathbf{v} f_e(\mathbf{v}) \delta(\omega - \mathbf{k} \cdot \mathbf{v}) + N_i \left| \frac{\chi_e(\mathbf{k}, \omega)}{\epsilon(\mathbf{k}, \omega)} \right|^2 \int d\mathbf{v} f_i(\mathbf{v}) \delta(\omega - \mathbf{k} \cdot \mathbf{v})$$

$$\epsilon(\mathbf{k}, \omega) = 0$$

$$\omega_{pl}(k) \approx \omega_{pe} (1 + 3\lambda_D^2 k^2)$$

$$\omega_{ia}(k) \approx k \sqrt{\frac{T_e + 3T_i}{m_i}}$$

