# PMSE Experiment Using PFISR Group Four(ier) ~ ISR School 2019

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# Background

PMSE are strong radar echoes from the cold summer mesosphere at high latitudes (La Hoz et al., 2006)

It was first reported as: Mesospheric echoes

Seasonal variation

# Fig. 1. Height of occurrence of mesospheric echoes on a daily basis from February 1979 to December 1980.

## First report of PMSE (Ecklun and Balsley, 1981)



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# Motivation

Learn to correctly use the Poker Flat Incoherent Scatter Radar (PFISR) observations to further investigate the Ionospheric D-region plasma signatures related to polar mesospheric summer echoes (PMSE).

Identify PMSE in the data and their main characteristics.

Where and when to observe PMSE?

• Typically 80-90 km

- Spatially patchy
- Polar regions

• Occurs May through August

Existing AMISR experiment, MSWinds26

• 130 us, 13 baud Barker code, 10 us baud, 5 us sampling

• High range resolution in D and E regions

• Alternating codes and uncoded long pulses for higher altitudes

• 4 beams (typically)



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Experiment summary:

- AMISR MSWinds26 modified to use 17 beams
- 1200 to 1400 LT observation window

Operating in this mode:

- Allows for high resolution in D and E regions
- Provides large spatial coverage
- Some E and F region measurements for context

# Barker code













## Good SNR!!

PMSE Experiment Using PFISR ISR 2019

# Long Pulse



## Range ambiguity function





Removing ambiguity function from ACF



Echo signal in a IPP



ACF and Power Spectrum at altitude



# Power spectrum for all altitudes

Echo power is strong Range resolution is poor



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Removing ambiguity function from ACF







## Power spectrum for all altitudes

Echo power is very weak Range resolution is the best ,which equals to baudlength!

ACF and Power Spectrum at altitude

# SNR comparison



The Signal-to-Noise Ratio of Long Pulse is Maximum, so we can detect higher altitude with long pulse.

But We need to see strong echoes with good resolution at low altitude, so we need to choose Barker code

## Conditions

- Annual Cycle
  - May August.
  - Peak in July.
- Diurnal Cycle
  - Peak at 12UT and 2UT.
- Cold temperatures

## (Williams et al, 1995) (Lübken and von Zahn, 1991)



## MESOPAUSE TEMPERATURE

1. JULY

MESOPAUSE ALTITUDE

CIRA 1986

1.OCT.

250

200

150

100 L

[K]

emperature

1.APR

## Conditions

- Water vapour
  - Less ppmv requires lower temperature as in early May.
  - 82 ~90 km altitude range shrinks as season progresses.



- (1) low temperatures and water vapor availability allow the ice particles to grow
- (2) charge aerosol number density is comparable to the local plasma density
- (3) aerosol's low mobility "implies that small-scale structures in their distribution can persist for a long time without being smeared out by diffusion.

$$\log_{10} P_{sat} = 12.537 - \frac{2663.5}{T}$$

## Marti and Manersberger (1993):

## How PMSE looks to PFISR

- Altitude range 80 ~90 km
- Frequency Spectra
  - Narrow doppler.
  - Strong response.



Nicolls et al, 2009



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Beam 6 PDS at ~84 km (range)

# Conclusion

We used PFISR to study Polar Mesospheric Summer Echoes

To observe the PMSEs, we used 17 beams and chose a time slot near 12:00 LT

PMSEs are found in high latitude regions during the summer in the mesosphere

To observe a PMSE, there must be water vapor, dust particles, and ice aerosols

We observed a PMSE around 20:00 UT at an altitude of about 80 km

The particularly strong season for PMSEs is coming to an end