

# ISR Data Analysis and Fitting 1

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# Outline

- ① Introduction
- ② Electron Density From Power
- ③ Inverse Problems
- ④ Fitting ISR Data
- ⑤ Derived Parameters

# Quick Review

Rough summary of topics introduced in previous lectures:

- **Hardware:**
  - Send megawatt pulses, receive femtowatt signals
- **Stochastic Processes:**
  - Voltage samples of received signals are correlated Gaussian random variables
- **Autocorrelation Function (ACF):**
  - Plasma parameters are encoded in the second moment
- **Ambiguity:**
  - Measurement technique influences the measurement
- **ISR Theory:**
  - Relationship between ACF (equivalently, the power spectrum) and ionospheric plasma parameters:  $N_e$ ,  $T_e$ ,  $T_i$ ,  $V_{los}$

# Quick Review

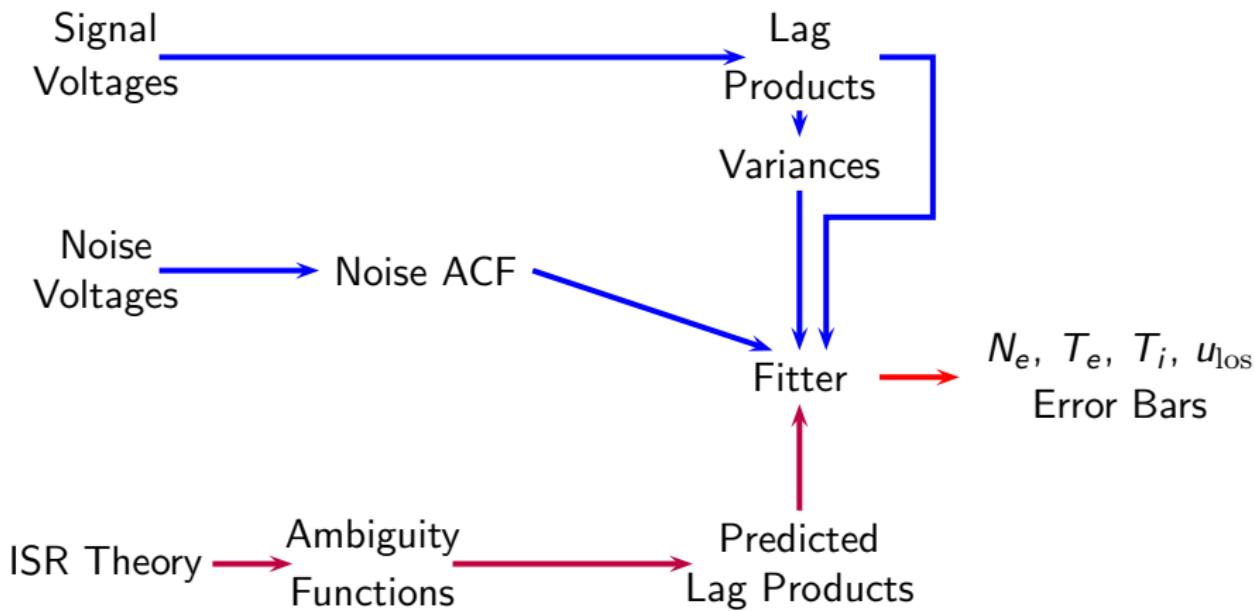
Rough summary of topics introduced in previous lectures:

- **Hardware:**  $\Leftarrow$  a lot of engineering
  - Send megawatt pulses, receive femtowatt signals
- **Stochastic Processes:**  $\Leftarrow$  statistics
  - Voltage samples of received signals are correlated Gaussian random variables
- **Autocorrelation Function (ACF):**  $\Leftarrow$  more statistics
  - Plasma parameters are encoded in the second moment
- **Ambiguity:**  $\Leftarrow$  signal processing
  - Measurement technique influences the measurement
- **ISR Theory:**  $\Leftarrow$  plasma physics
  - Theoretical relationship between ACF (equivalently, the power spectrum) and ionospheric plasma parameters:  $N_e$ ,  $T_e$ ,  $T_i$ ,  $V_{los}$

# Quick Review

But, how do we put it all together to get  $N_e$ ,  $T_e$ ,  $T_i$ ,  $V_{los}$ ? Why should we trust the values we get?

# The ISR Signal Processing Chain

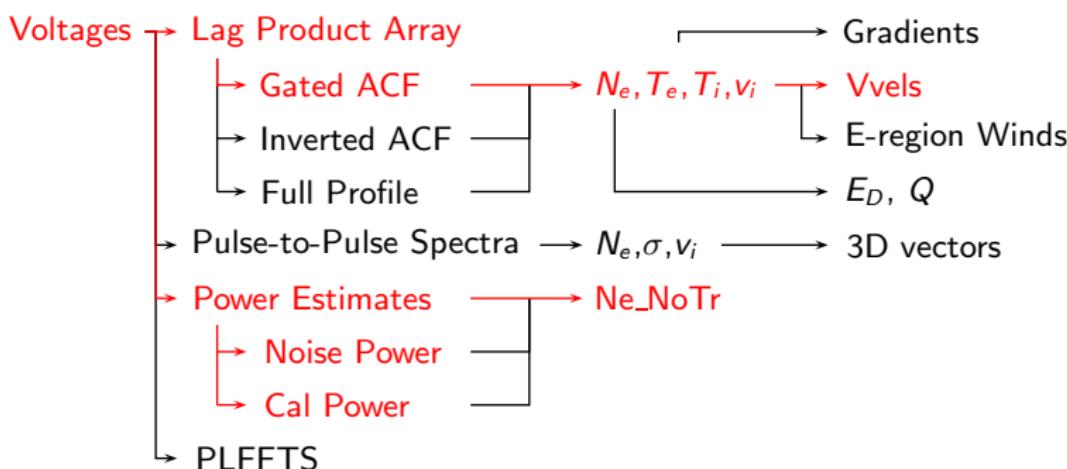


(provided by Roger Varney)

# ISR Data Levels

## Summary of ISR data products:

Level 0 → Level 1 → Level 2 → Level 3



## Electron Density From Power

# Electron Density Determination

- ISR Equation for Power received (Watts)

$$P_{\text{Rx}} = \frac{P_{\text{Tx}} \tau_p}{R^2} K_{\text{sys}} \frac{N_e}{(1 + k^2 \lambda_{De}^2) (1 + k^2 \lambda_{De}^2 + T_e/T_i)}$$

$$\lambda_{De} = \sqrt{\frac{\epsilon_0 k_B T_e}{e^2 N_e}} \quad k = \frac{4\pi}{\lambda_{\text{Tx}}} \quad R = \text{Range} \quad \tau_p = \text{Pulse Length (s)}$$

- $P_{\text{Rx}}$  in Watts: determined by comparing relative power received to direct signal injection (cal pulses)
- $K_{\text{sys}}$ : the “System Constant” involves antenna gain, effective area, etc. For PFISR  $K_{\text{sys}} \sim 10^{-19} \text{ m}^5 \text{s}^{-1}$ .
- Can determine  $K_{\text{sys}}$  by comparing estimated  $N_e$  to absolute  $N_e$  measurements, e.g.:
  - Ionosonde  $f_{0F2}$
  - ISR plasma line frequency
  - Faraday rotation (e.g. Jicamarca)

# Electron Density Estimation

Can estimate electron density using estimates of received power!

# Estimators

**Estimator:** An estimator is a statistic. A statistic is a function used to estimate a parameter from a sample.

- Expectation value (mean):

$$E[\hat{X}] = X$$

- Mean-squared Error:

$$MSE = E[(\hat{X} - X)^2] + (E[\hat{X}] - X)^2$$

- Mean-squared Error: Variance + Bias

## Notation:

- Measurement:  $\tilde{V}$
- Estimate:  $\hat{V}$

# Power Estimation

Given K samples of  $\tilde{v}_i$ ,

$$\hat{P} = \frac{1}{K} \sum_{i=0}^{K-1} \tilde{v}_i \tilde{v}_i^*, \quad \text{Var} \left\{ \hat{P} \right\} = \frac{P^2}{K}$$

but in general there is noise in  $\tilde{v}_i$ , so we need to estimate the noise:

$$\hat{N} = \frac{1}{K_{noise}} \sum_{i=0}^{K_{noise}-1} \tilde{v}_{noise,i} \tilde{v}_{noise,i}^*$$

and remove it:

$$\hat{S} = \frac{1}{K} \sum_{i=0}^{K-1} \tilde{v}_i \tilde{v}_i^* - \hat{N}$$

$$\text{Var} \left\{ \hat{S} \right\} = \frac{(S + N)^2}{K} + \frac{N^2}{K_{noise}}$$

# Power Estimation

Generally, we can design our experiment so that  $K_{noise} \gg K$  then:

$$\frac{\delta \hat{S}}{S} = \frac{1}{\sqrt{K}} \left( 1 + \frac{1}{S/N} \right)$$

For example,  $\frac{\delta \hat{S}}{S} = 0.5$  with a  $S/N = 0.1$  requires  $K = 484$ .

This assumes the data and noise samples are taken far apart and are uncorrelated.

# Received Power Estimation

Power is in units of  $(\text{ADC counts})^2$ , but we need units of watts for ISR radar equation! So we inject a known noise power from a calibration noise source. Then, received power is estimated with:

$$\hat{P}_{\text{Rx}} = k B_{\text{Rx}} T_{\text{cal}} \frac{\hat{S}}{\hat{C}}$$

$$\hat{C} = \frac{1}{K_{\text{cal}}} \sum_{i=0}^{K_{\text{cal}}-1} \tilde{v}_{\text{cal},i} \tilde{v}_{\text{cal},i}^* - \hat{N}$$

with a variance of:

$$\text{Var} \left\{ \hat{P}_{\text{Rx}} \right\} = (k B_{\text{Rx}} T_{\text{cal}})^2 \left( \frac{\text{Var} \left\{ \hat{S} \right\}}{C^2} + \frac{S^2 \text{Var} \left\{ \hat{C} \right\}}{C^4} \right)$$

where  $B_{\text{Rx}}$  is the receiver bandwidth and  $T_{\text{cal}}$  is the noise temperature of the calibration noise source.

# Electron Density Estimation

Recall:

$$P_{\text{Rx}} = \frac{P_{\text{Tx}} \tau_p}{R^2} K_{\text{sys}} \frac{N_e}{(1 + k^2 \lambda_{De}^2) (1 + k^2 \lambda_{De}^2 + T_e/T_i)}$$

- Define:

$$\zeta = \frac{2}{(1 + k^2 \lambda_{De}^2) (1 + k^2 \lambda_{De}^2 + T_e/T_i)}$$

- and then solve for  $N_e$ . So an estimate of electron density,  $\hat{N}_e$  is given by:

$$\hat{N}_e = \frac{2R^2}{\zeta \tau_p K_{\text{sys}} P_{\text{Tx}}} \hat{P}_{\text{Rx}} = A \hat{P}_{\text{Rx}}$$

- with variance:

$$(\delta \hat{N}_e)^2 = \text{Var} \left\{ \hat{N}_e \right\} = A^2 \text{Var} \left\{ \hat{P}_{\text{Rx}} \right\}$$

# Reporting Electron Density Estimates and Fractional Error

## Corrected/Uncorrected:

$$\text{Temperature Correction: } \zeta = \frac{2}{(1 + k^2 \lambda_{De}^2) (1 + k^2 \lambda_{De}^2 + T_e/T_i)}$$

In SRI data hdf5 files:

- Uncorrected  $N_e$ , **known as “Ne\_noTr”**:  $\zeta = 1$ .
  - $T_e/T_i = 1$
  - $k^2 \lambda_{De}^2 \ll 1$ .
- $N_e$  with model, **known as “Ne\_Mod”**:
  - Compute  $\zeta$  using an empirical model of  $T_e/T_i$  as a function of altitude.
- For SRI hdf5 files, electron density error is reported as fractional error, **known as “dNeFrac”**:
  - $\delta \hat{N}_e / \hat{N}_e$

# Plasma parameters from ISR

So what about the rest of the plasma parameters?



## Inverse Problems

# Parameter Estimation and Inverse Problems

**Inverse Problem:** Given observations and a forward model, what are the parameters of the model that best represent the observation?

- A forward model,  $h$ , predicts observations,  $\mathbf{y}$ , for a given set of parameters  $\mathbf{p}$

$$\mathbf{y} = h(\mathbf{p})$$

- But real measurements are noisy! So, a forward model for noisy measurements,  $\mathbf{z}$ , can be written:

$$\mathbf{z} = h(\mathbf{p}) + \mathbf{e}$$

- where  $\mathbf{e}$  is the error in the measurement due to noise. We can construct a covariance matrix of the errors:

$$\text{Cov}\{\mathbf{e}\} = \Sigma_e$$

How do we determine the best estimate of the parameters  $\mathbf{p}$ ?

# Methods for Solving Inverse Problems

An incomplete list of methods:

- Least-squares
- Maximum Likelihood
- Bayesian Inference
- Maximum Entropy

# Least-Squares Estimation

Least-Squares Estimate:

$$\hat{\mathbf{p}}_{\text{LS}} : \min_{\mathbf{p}} [\mathbf{h}(\mathbf{p}) - \mathbf{z}]^T \Sigma_e^{-1} [\mathbf{h}(\mathbf{p}) - \mathbf{z}]$$

For a diagonal  $\Sigma_e^{-1}$  (uncorrelated measurement errors):

$$\hat{\mathbf{p}}_{\text{LS}} : \min_{\mathbf{p}} \sum_i \frac{[h_i(\mathbf{p}) - z_i]^2}{\sigma_i^2}$$

- If  $\mathbf{z}$  is jointly gaussian, then the least-squares estimate is equivalent to the maximum likelihood estimate.
- A commonly used numerical technique for iteratively solving nonlinear least squares problems is the Levenberg-Marquardt algorithm
- Standard Levenberg-Marquardt packages:
  - FORTRAN: MINPACK lmdif.f and lmder.f
  - Python: scipy.optimize.leastsq (wrapper around lmdif and lmder)
  - Matlab: Optimization Toolbox lsqnonlin
  - IDL: LMFIT
- Levenberg-Marquardt requires a good initial guess

# Error Propagation (Linear Least-Squares)

Linear Least-Squares  $h(\mathbf{p}) = H\mathbf{p}$

$$\begin{aligned}\hat{\mathbf{p}}_{\text{LS}} &= \left[ H^T \Sigma_e^{-1} H \right]^{-1} H^T \Sigma_e^{-1} \mathbf{z} \\ &= \left[ \tilde{H}^T \tilde{H} \right]^{-1} \tilde{H}^T \tilde{\mathbf{z}}\end{aligned}$$

where  $\tilde{H} = \Sigma_e^{-1/2} H$  and  $\tilde{\mathbf{z}} = \Sigma_e^{-1/2} \mathbf{z}$

Recall the property of jointly Gaussian random variables:

$$\mathbf{Y} = A\mathbf{X} \Rightarrow \text{Cov}\{\mathbf{Y}\} = A \text{Cov}\{\mathbf{X}\} A^T$$

Thus

$$\begin{aligned}\text{Cov}\{\hat{\mathbf{p}}_{\text{LS}}\} &= \left[ \tilde{H}^T \tilde{H} \right]^{-1} \tilde{H}^T \text{Cov}\{\tilde{\mathbf{z}}\} \tilde{H} \left[ \tilde{H}^T \tilde{H} \right]^{-1} \\ &= \left[ \tilde{H}^T \tilde{H} \right]^{-1}\end{aligned}$$

(Note  $\text{Cov}\{\tilde{\mathbf{z}}\} = \Sigma_e^{-1/2} \text{Cov}\{\mathbf{z}\} \Sigma_e^{-1/2} = I$ )

# Error Propagation (Nonlinear Least Squares)

Suppose we are minimizing

$$\hat{\mathbf{p}}_{\text{LS}} \min_{\mathbf{p}} \sum_i \frac{[h_i(\mathbf{p}) - z_i]^2}{\sigma_i^2}$$

Linearize the problem in the vicinity of the solution

$$\text{Cov} \{ \hat{\mathbf{p}}_{\text{LS}} \} \approx [\tilde{J}^T \tilde{J}]^{-1}$$

where the Jacobian  $\tilde{J}$  is evaluated at the solution  $\mathbf{p} = \hat{\mathbf{p}}_{\text{LS}}$

$$\tilde{J} = \begin{pmatrix} \frac{1}{\sigma_0} \frac{\partial h_0}{\partial p_0} & \frac{1}{\sigma_0} \frac{\partial h_0}{\partial p_1} & \cdots & \frac{1}{\sigma_0} \frac{\partial h_0}{\partial p_{M-1}} \\ \frac{1}{\sigma_1} \frac{\partial h_1}{\partial p_0} & \frac{1}{\sigma_1} \frac{\partial h_1}{\partial p_1} & \cdots & \frac{1}{\sigma_1} \frac{\partial h_1}{\partial p_{M-1}} \\ \vdots & \vdots & \vdots & \vdots \\ \frac{1}{\sigma_{N-1}} \frac{\partial h_{N-1}}{\partial p_0} & \frac{1}{\sigma_{N-1}} \frac{\partial h_{N-1}}{\partial p_1} & \cdots & \frac{1}{\sigma_{N-1}} \frac{\partial h_{N-1}}{\partial p_{M-1}} \end{pmatrix}$$

$\tilde{J}$  is  $N \times M$  (tall and skinny)

- Levenberg-Marquart computes  $\tilde{J}$  at every iteration internally
- Standard packages usually have an option to return either  $\tilde{J}$ , and/or  $[\tilde{J}^T \tilde{J}]^{-1}$  from the last iteration

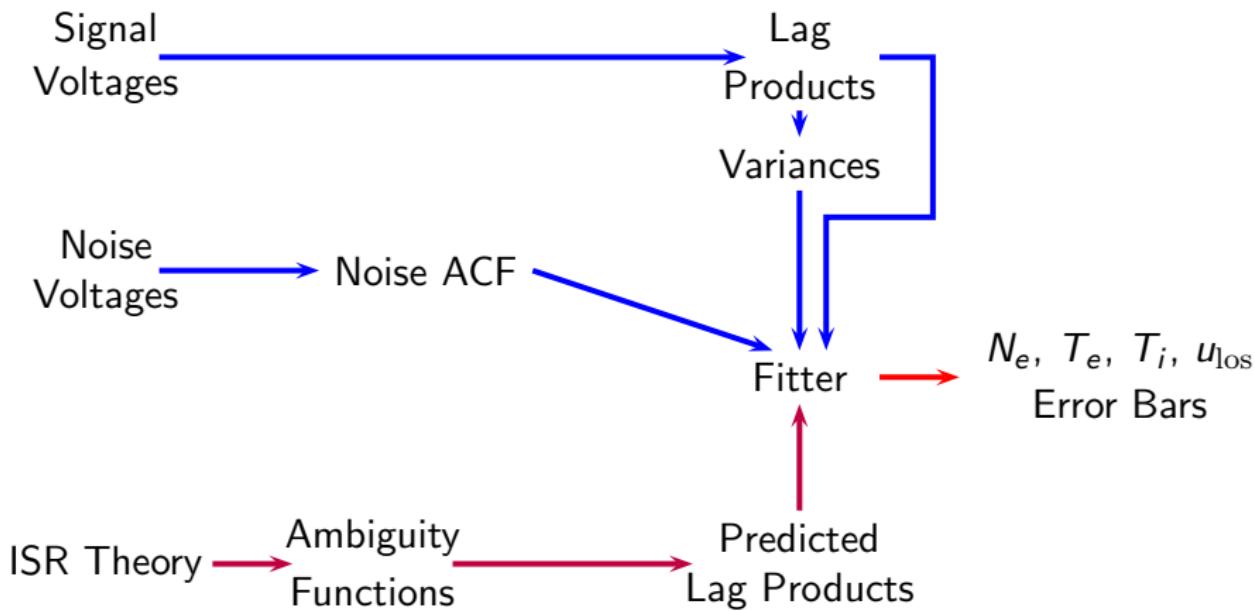
# Is the result meaningful?

Two important things to check:

- evaluate the “goodness of fit”: for AMISR, reduced chi-squared is provided
- check covariance of fitted parameters

## Fitting ISR Data

# The ISR Signal Processing Chain



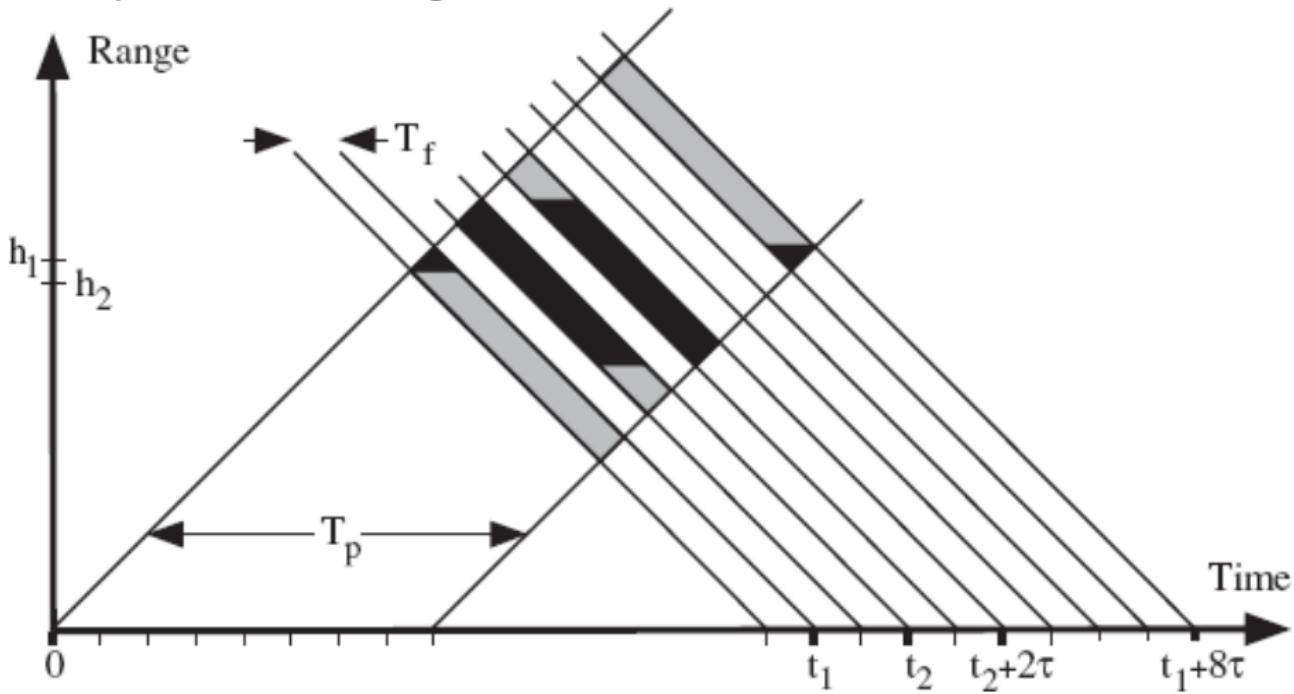
(Seen previously in Radar 3: Statistical Signal Processing by Roger Varney)

# The ISR Signal Processing Chain

- Run the experiment, collect voltage samples
- Estimate ACFs from the voltage samples for each range gate
- Gate the ACFs
- Set up forward model
- Solve the inverse problem

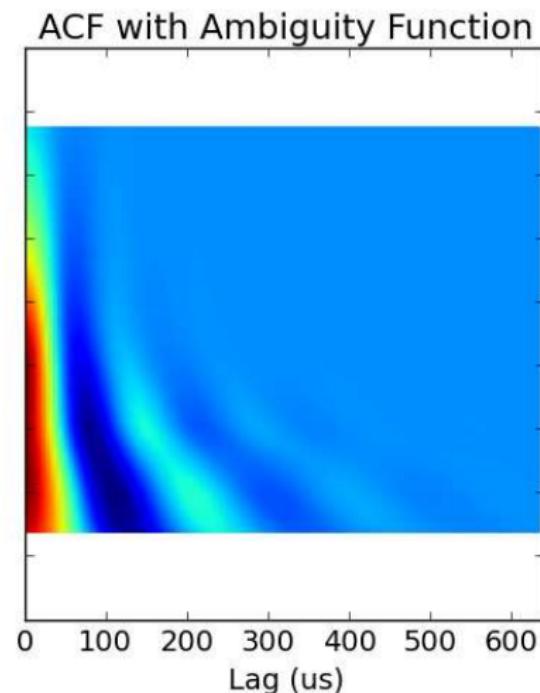
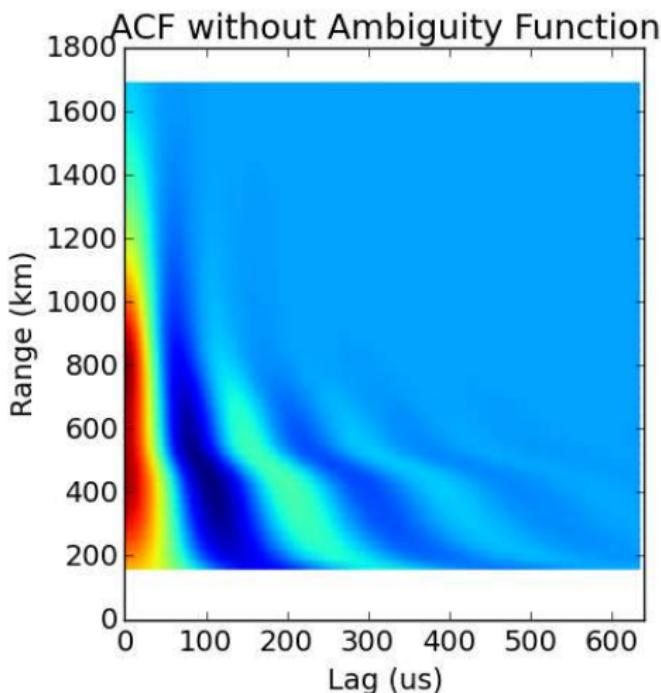
# ACF Estimation

Example: Uncoded Long Pulse



# Blurring of ACFs by Ambiguity Functions

A particular exaggerated example using 1.5 ms long pulses and a profile with a sharp  $T_e$  gradient at 500 km.



# ACF Gating

Uncoded Long Pulse ACF Gating:

- Different lags have different range ambiguity
- For fitting, want all ACF lags to share common range extent

**Solution:** Sum Rules

- Sum lags so that all lags have the same range extent

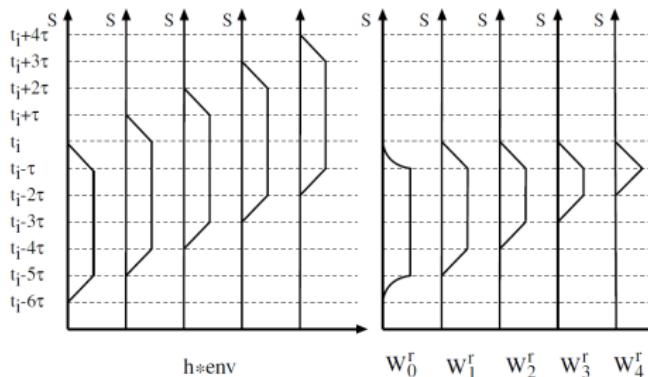
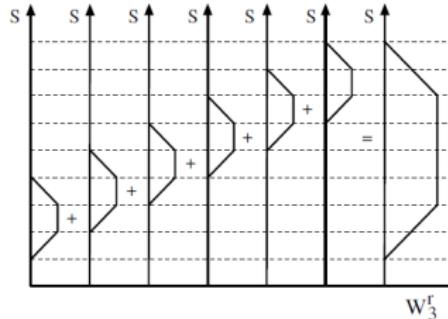
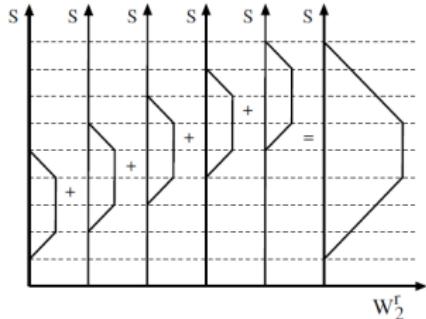
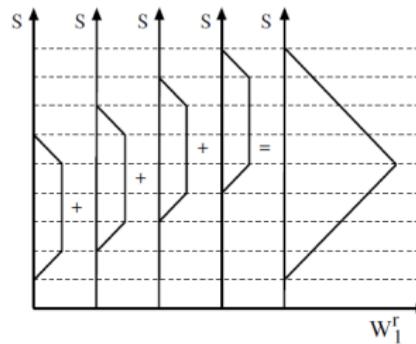
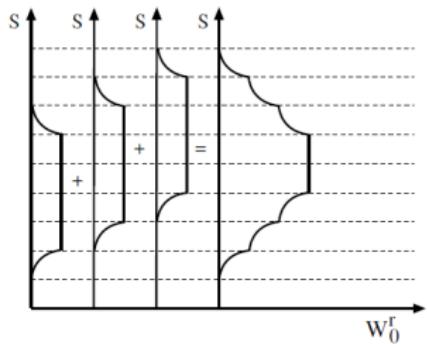


Figure from Nygrén ISR Book

# ACF Gating and Sum Rules



Figures from Nygrén ISR Book

# ACF Gating and Sum Rules

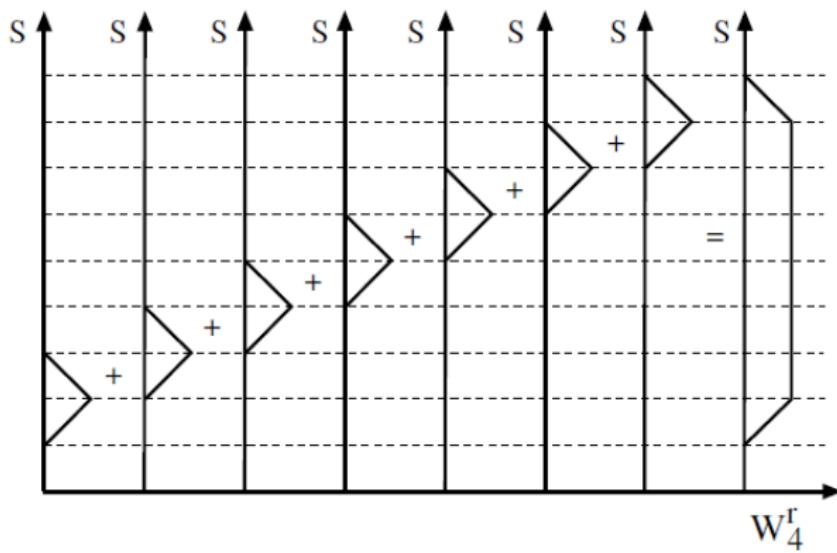


Figure from Nygrén ISR Book

# Creating Forward Models

The forward model has two portions

## ① Physics and Chemistry (ISR Theory)

- Assume Maxwellian distributions?
- Constraints on  $T_e$  and  $T_i$ ?
- Constraints on ion composition? Chemistry model?
- Magnetic field effects

## ② Instrumental Effects and Signal Processing

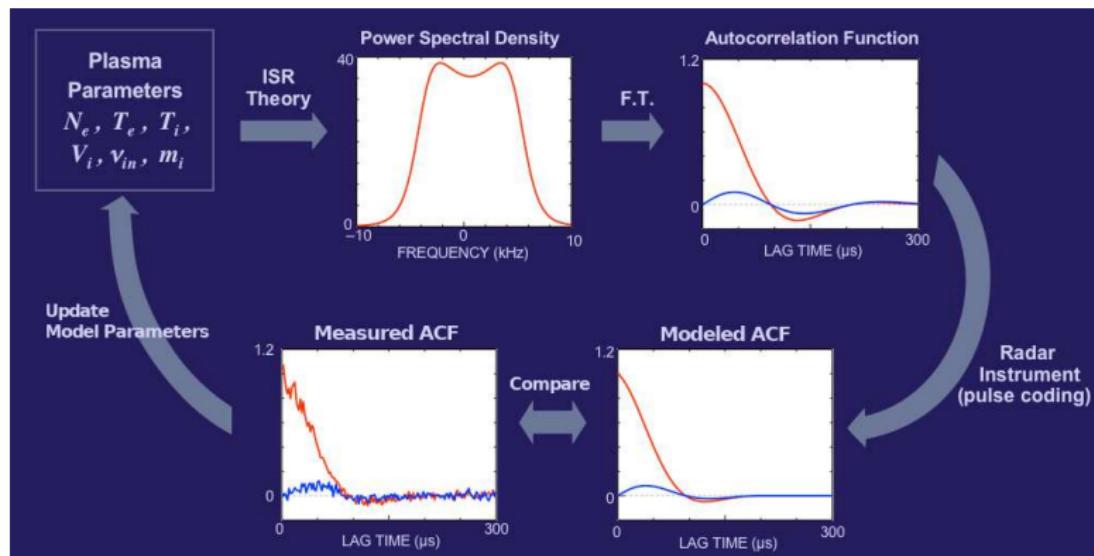
- Sampling and Aliasing
- Windowing
- Ambiguity Functions

Best practice is to build the instrumental effects into the forward model.

**Do not manipulate the data in an attempt to undo the instrumental effects!**

# Solving the Inverse Problem

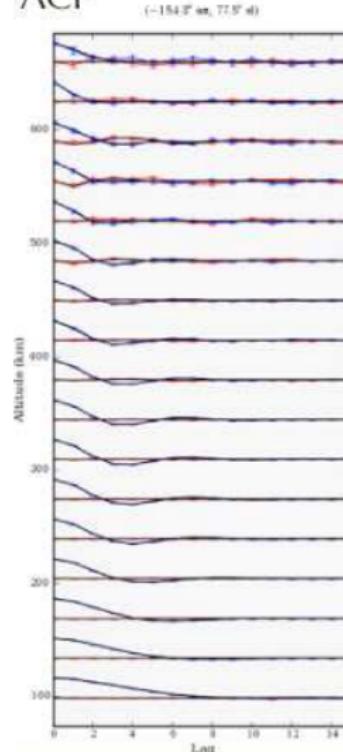
- gate ACF integrate in time
- evaluate ISR theory, FFT, and convolved with ambiguity functions
- Levenburg-Marquardt least-squares solver



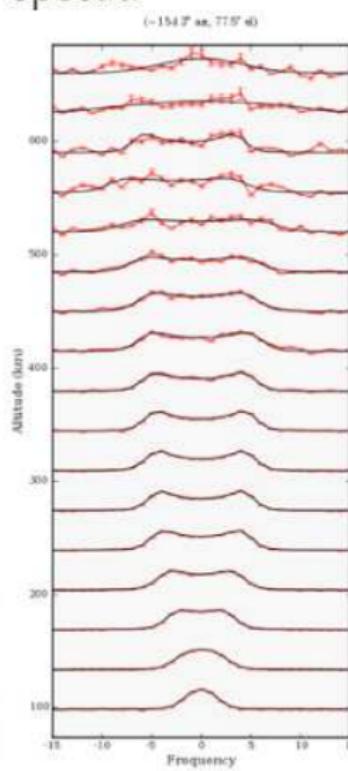
# Example PFISR Long Pulse Fits

6-14-2007 22:332-22:584 UT

ACF

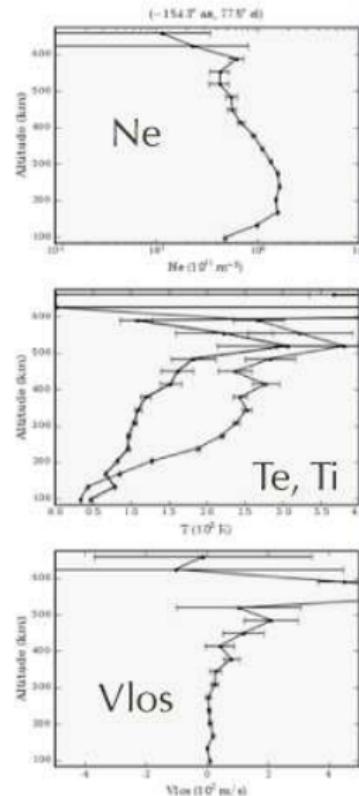


Spectra



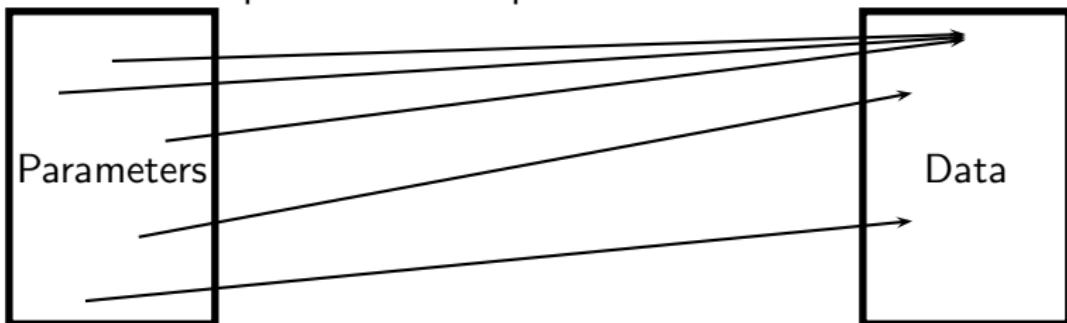
6-14-2007 22:332-22:584

(~154.3° az, 77.5° d)



## III-Posed and III-Conditioned Problems

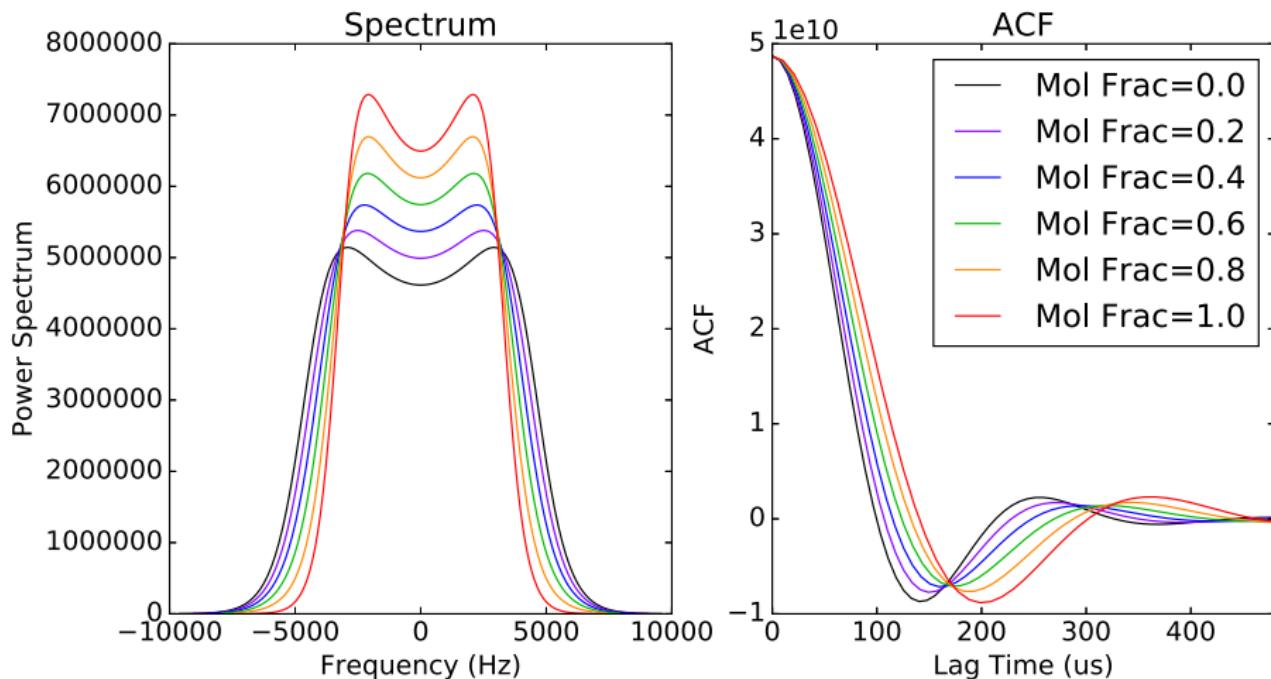
What happens if my forward model maps different points in parameter space to almost the same points in data space?



- **III-Posed Problem:** Multiple points in parameters space map to exactly the same point in data space.  
 $[\tilde{H}^T \tilde{H}]$  is singular, inverse problem is impossible
- **III-Conditioned Problem:** Multiple points in parameters space map to nearly the same point in data space.  
 $[\tilde{H}^T \tilde{H}]$  is nearly singular, inverse problem is unstable given noisy data

### III-Conditioned ISR Theory: Molecular Ion Chemistry

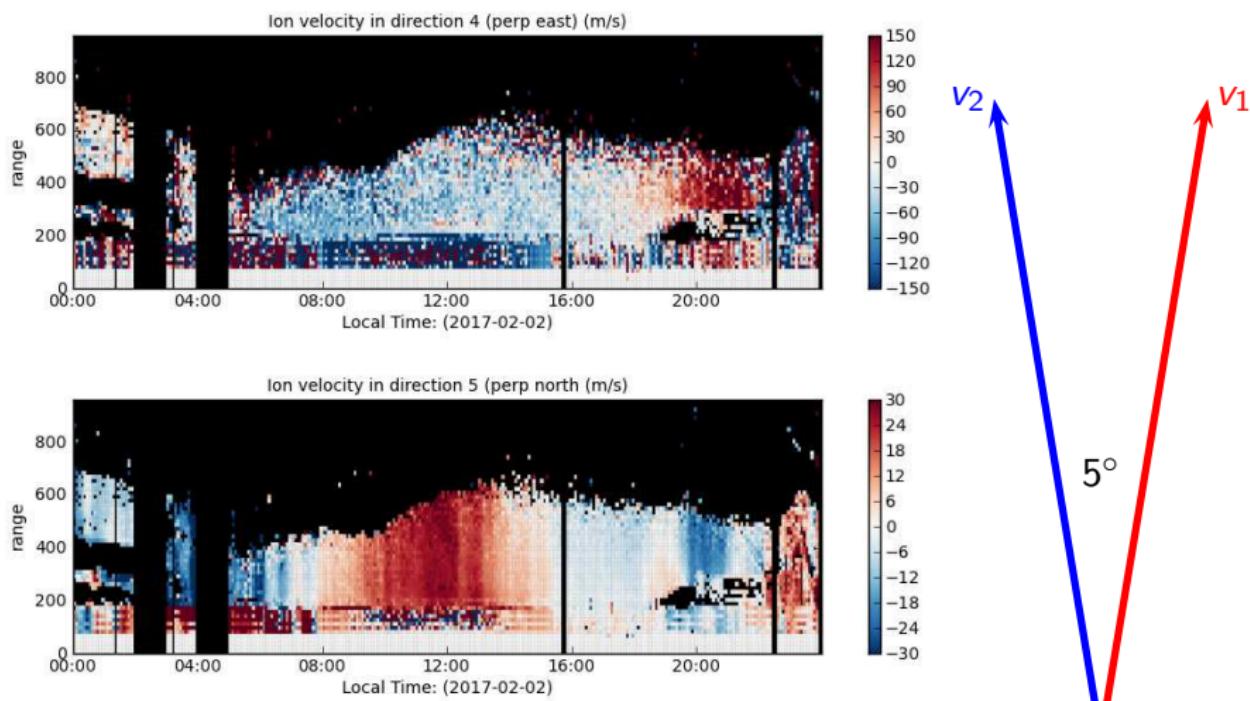
Mixtures of  $O^+$  and  $O_2^+$  using  $N_e = 10^{11}$ ,  $T_e = T_i = 1000$  K



ISR spectrum measures  $\sqrt{\frac{T_i}{m_i}}$ , ambiguity between  $T_i$  and  $m_i$

## Derived Parameters

# Jicamarca $\mathbf{E} \times \mathbf{B}$ Drifts



Why does zonal ( $u$ ) look noisier than vertical ( $w$ )?

# Error Analysis of Jicamarca Drifts

$$\begin{bmatrix} v_1 \\ v_2 \end{bmatrix} = \begin{bmatrix} \sin(2.5^\circ) & \cos(2.5^\circ) \\ -\sin(2.5^\circ) & \cos(2.5^\circ) \end{bmatrix} \begin{bmatrix} u \\ w \end{bmatrix}$$

$$\begin{bmatrix} u \\ w \end{bmatrix} = \frac{1}{2\sin(2.5^\circ)\cos(2.5^\circ)} \begin{bmatrix} \cos(2.5^\circ) & -\cos(2.5^\circ) \\ \sin(2.5^\circ) & \sin(2.5^\circ) \end{bmatrix} \begin{bmatrix} v_1 \\ v_2 \end{bmatrix}$$

$$\begin{bmatrix} u \\ w \end{bmatrix} = \mathbf{A} \begin{bmatrix} v_1 \\ v_2 \end{bmatrix}$$

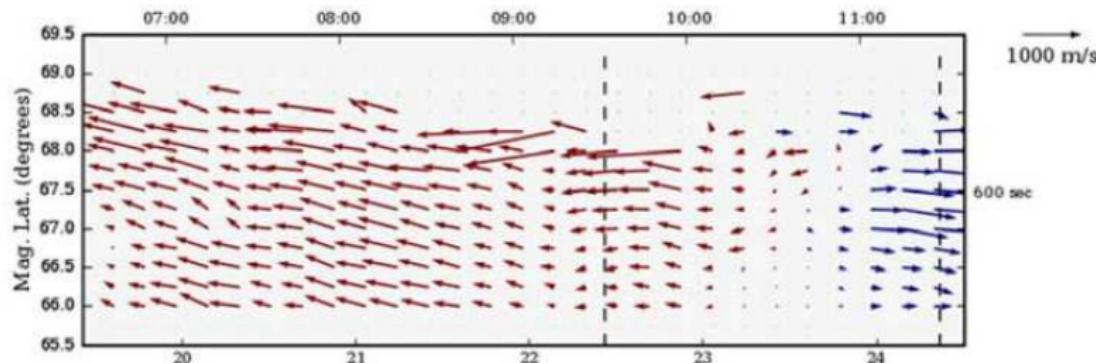
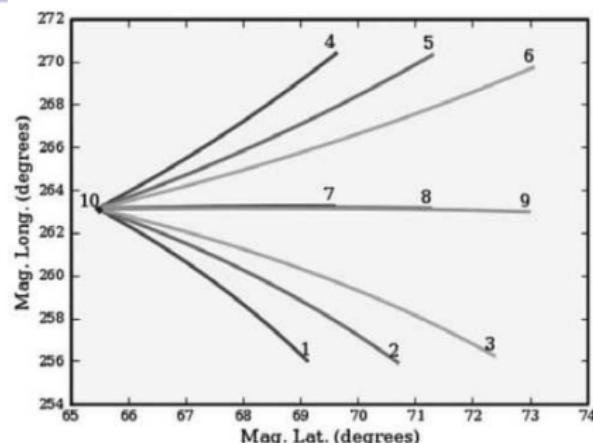
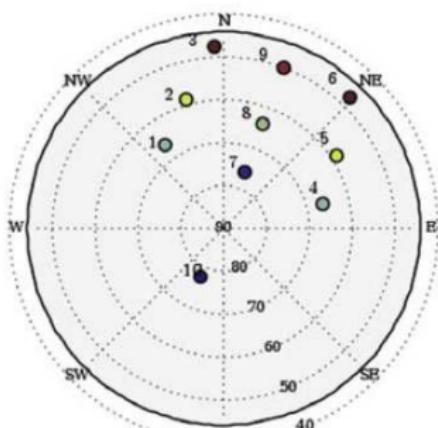
$$\begin{bmatrix} \text{Var}\{u\} & \text{Cov}\{u, w\} \\ \text{Cov}\{u, w\} & \text{Var}\{w\} \end{bmatrix} = \mathbf{A} \begin{bmatrix} \sigma_v^2 & 0 \\ 0 & \sigma_v^2 \end{bmatrix} \mathbf{A}^T$$

$$\begin{bmatrix} \text{Var}\{u\} & \text{Cov}\{u, w\} \\ \text{Cov}\{u, w\} & \text{Var}\{w\} \end{bmatrix} = \begin{bmatrix} \frac{1}{2} \frac{\sigma_v^2}{\sin^2(2.5^\circ)} & 0 \\ 0 & \frac{1}{2} \frac{\sigma_v^2}{\cos^2(2.5^\circ)} \end{bmatrix}$$

# AMISR F-region 1-D Vector Electric Fields

- In F-region assume  $\mathbf{v}_i = \frac{\mathbf{E} \times \mathbf{B}}{B^2}$
- Assume  $\mathbf{E} \cdot \mathbf{B} = 0$  (no parallel fields)
- LOS velocity is related to  $\mathbf{E}$  perpendicular to LOS and  $\mathbf{B}$
- Assume  $\mathbf{E}$  is uniform in magnetic longitude, but varies with magnetic latitude
- Assume  $\mathbf{E}$  fields map along equipotential field lines
- Different range gates correspond to different magnetic latitudes
- Fit for 2-components of  $\mathbf{E}$  as a function of magnetic latitude

# 1-D Electric Field Estimation



# Interpretation of E-region Ion Velocities

Ion Momentum Equation:

$$0 = e(\mathbf{E} + \mathbf{u}_i \times \mathbf{B}) - m_i \nu_{in} (\mathbf{u}_i - \mathbf{u}_n)$$

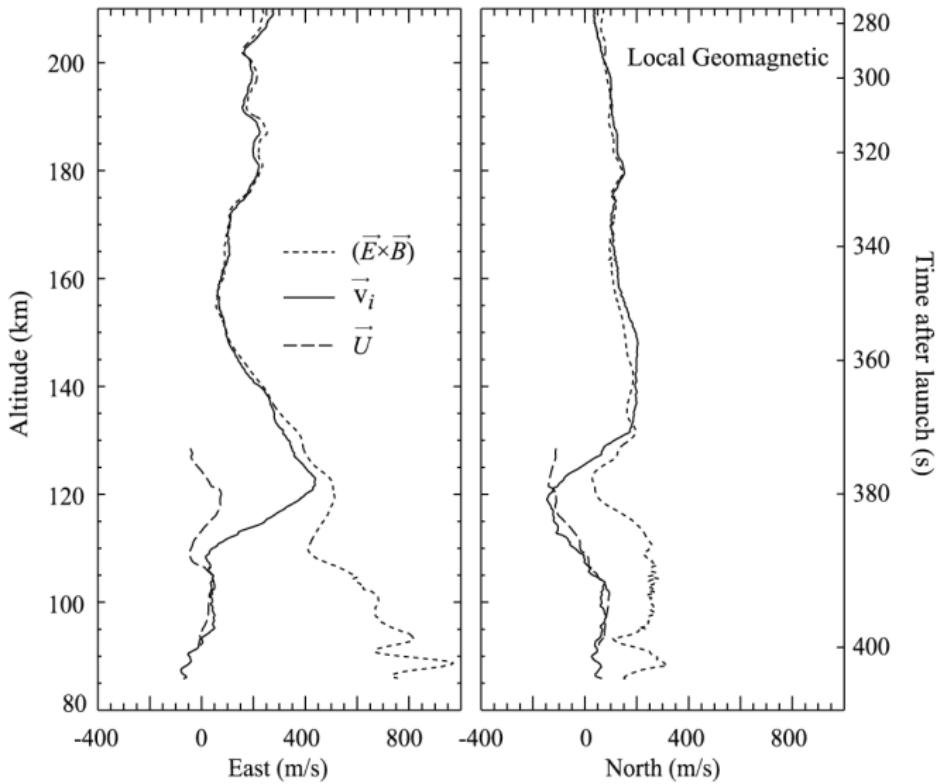
Collisional Limit (D-region):  $\mathbf{u}_i = \mathbf{u}_n$

Collisionless Limit (F-region):  $\mathbf{u}_i = \frac{\mathbf{E} \times \mathbf{B}}{B^2}$

$$\text{E-region: } \mathbf{u}_i = \begin{pmatrix} \frac{1}{1+\kappa_i^2} & \frac{-\kappa_i}{1+\kappa_i^2} & 0 \\ \frac{\kappa_i}{1+\kappa_i^2} & \frac{1}{1+\kappa_i^2} & 0 \\ 0 & 0 & 1 \end{pmatrix} \left[ \mathbf{u}_n + \frac{e}{m_i \nu_{in}} \mathbf{E} \right]$$

$$\kappa_i \equiv \frac{eB}{m_i \nu_{in}}$$

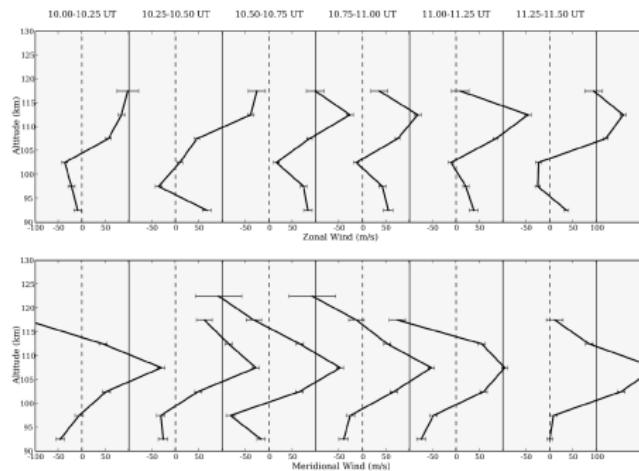
# Joule II Rocket Results (Sangali et al. 2009)



# E-region Neutral Wind Estimation

- Estimate vector E-region ion velocities from E-region LOS velocity
- Estimate vector F-region electric fields from F-region LOS velocity
- Map electric fields from F-region to E-region along equipotential field lines
- Solve for  $\mathbf{u}_n$

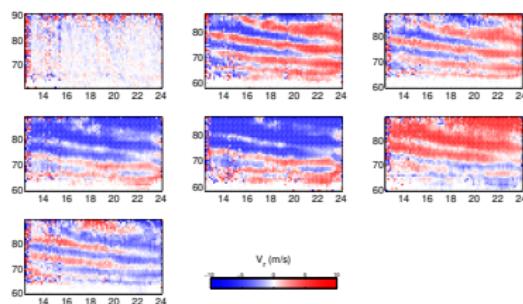
$$\mathbf{u}_n = \mathbf{u}_i - \frac{e}{m_i \nu_{in}} (\mathbf{E} + \mathbf{u}_i \times \mathbf{B})$$



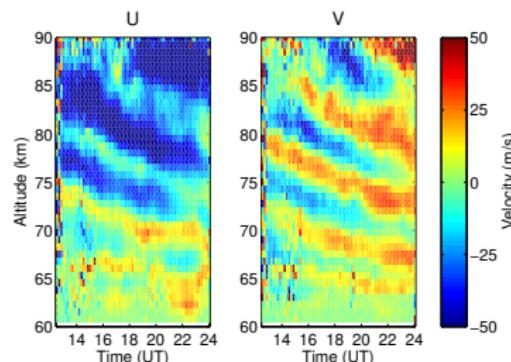
Heinselman and Nicolls (2008) Radio Sci.

# Mesospheric Vector Neutrals Winds

## Line of Sight Velocities



## Fitted Horizontal Velocities



$$\begin{pmatrix} V_{r,1} \\ \vdots \\ V_{r,7} \end{pmatrix} = \begin{pmatrix} \cos(\theta_1) \sin(\phi_1) & \cos(\theta_1) \cos(\phi_1) & \sin(\theta_1) \\ \vdots & \vdots & \vdots \\ \cos(\theta_7) \sin(\phi_7) & \cos(\theta_7) \cos(\phi_7) & \sin(\theta_7) \end{pmatrix} \begin{pmatrix} u \\ v \\ w \end{pmatrix}$$

$$\mathbf{v}_r = \mathbf{D}\mathbf{U}$$

$$\mathbf{U} = (\mathbf{D}^T \Sigma_{V_r}^{-1} \mathbf{D})^{-1} \mathbf{D}^T \Sigma_{V_r}^{-1} \mathbf{v}_r$$

# Derived Electrodynamical Parameters

- Conductivity

$$\sigma_P = N_e e^2 \left( \frac{\nu_{en}/m_e}{\nu_{en}^2 + \Omega_e^2} + \frac{\nu_{in}/m_i}{\nu_{in}^2 + \Omega_i^2} \right)$$

$$\sigma_H = N_e e^2 \left( \frac{\Omega_e/m_e}{\nu_{en}^2 + \Omega_e^2} - \frac{\Omega_i/m_i}{\nu_{in}^2 + \Omega_i^2} \right)$$

- Horizontal Currents

$$\mathbf{J} = \sigma_P (\mathbf{E} + \mathbf{u}_n \times \mathbf{B}) - \sigma_H \left[ (\mathbf{E} + \mathbf{u}_n \times \mathbf{B}) \times \frac{\mathbf{B}}{B} \right]$$

- Joule Heating

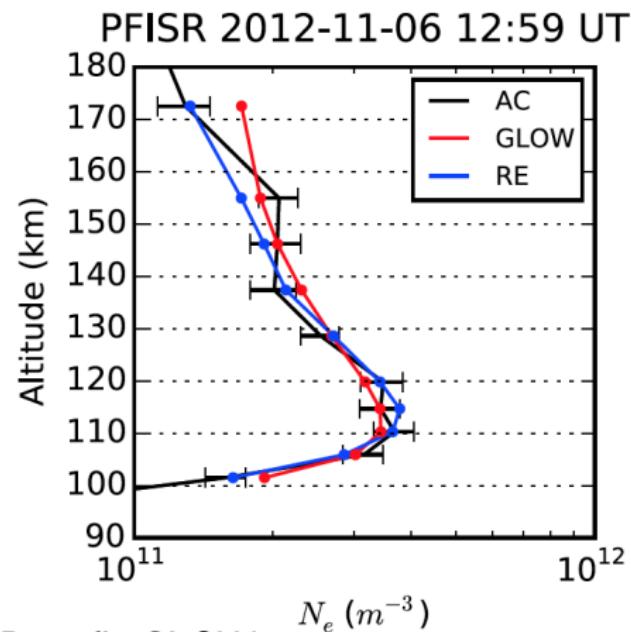
$$\begin{aligned} Q_J &= \mathbf{J} \cdot (\mathbf{E} + \mathbf{u}_n \times \mathbf{B}) \\ &= \sigma_P |\mathbf{E} + \mathbf{u}_n \times \mathbf{B}|^2 \end{aligned}$$

See Thayer (1998) *JGR* and Thayer and Semeter (2004) *JASTP*

# Precipitation Characteristics from $N_e$ Profile Inversion

- Input  $N_e$  profiles vs altitude (up-B beam)
- Estimate precipitating energy flux and characteristic energy
- Use a forward model of energetic electron transport, impact ionization, and recombination (e.g. GLOW).

Kaepller et al. (2015) *JGR*.



Best fit GLOW parameters:  
 $Q = 7.3 \pm 0.8 \text{ mW/m}^2$ ,  
 $\mathcal{E}_0 = 5.0 \pm 0.2 \text{ keV}$

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We infer plasma parameters from processed voltage samples:

- send pulses, record voltage samples: zero mean Gaussian random variables
- estimate the autocorrelation function, mean value and variance of lag products
- gate ACFs so all lag products have same range extent
- construct a forward model: convolution ambiguity function with the FFT of theoretical ISR spectrum
- use non-linear weighted least-squares to find parameters that minimize the difference between the forward model and the estimated (“measured”) autocorrelation function