IS Radar Data Examples: Basic and Derived Parameters

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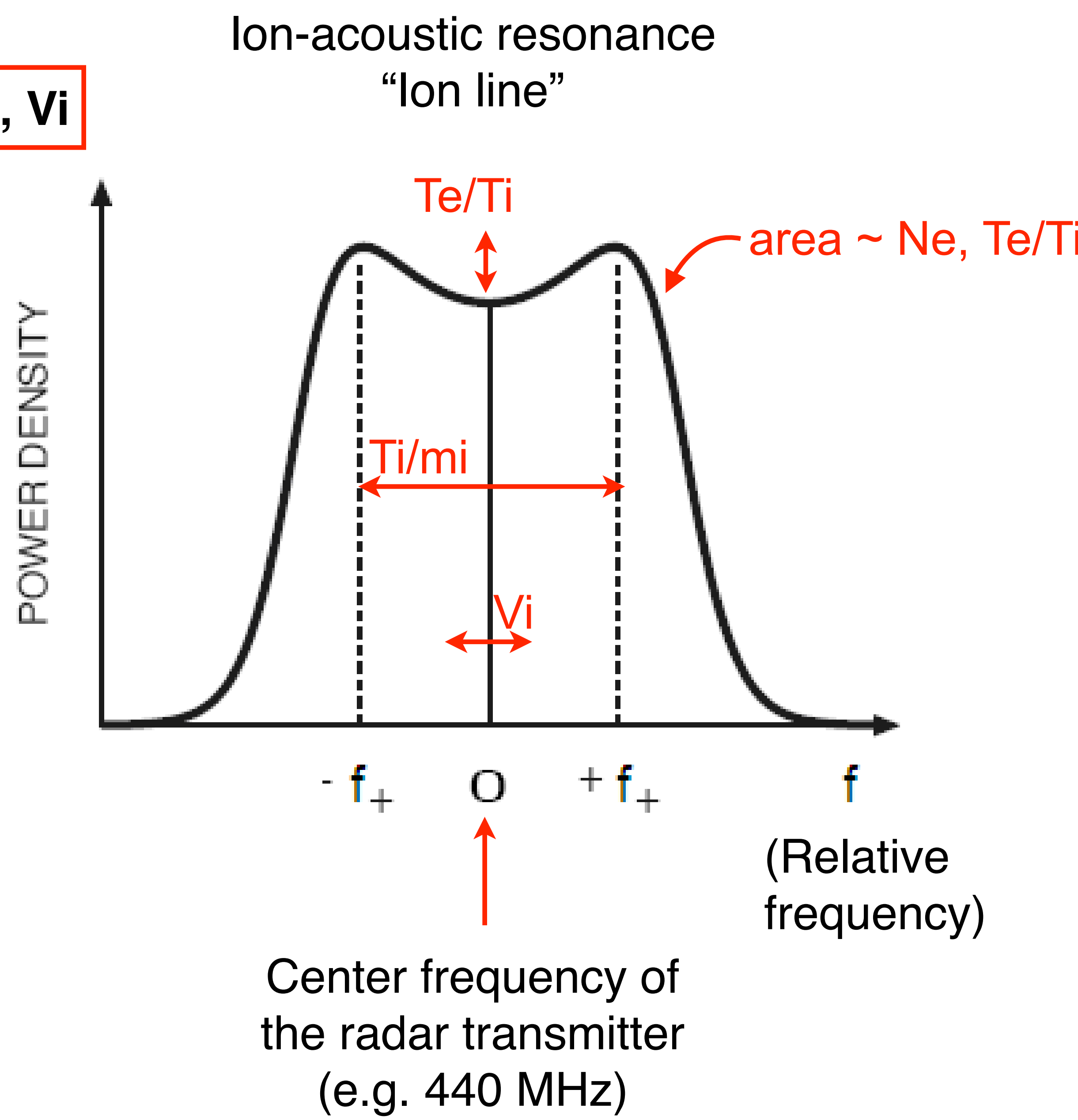
Topics covered:

- Basic measured plasma parameters
- Derived plasma and neutral parameters
- Range vs. altitude: Pointing considerations
- Altitude profiles vs. 2D altitude/time profiles
- Science examples using ISR data

Basic IS Radar Measured Parameters (Ion Line)

Ne, Te, Ti, Vi



Rules of Thumb

Ion temperature (Ti) to ion mass (mi) ratio from the width of the spectra

Electron to ion temperature ratio (Te/Ti) from "peak_to_valley" ratio

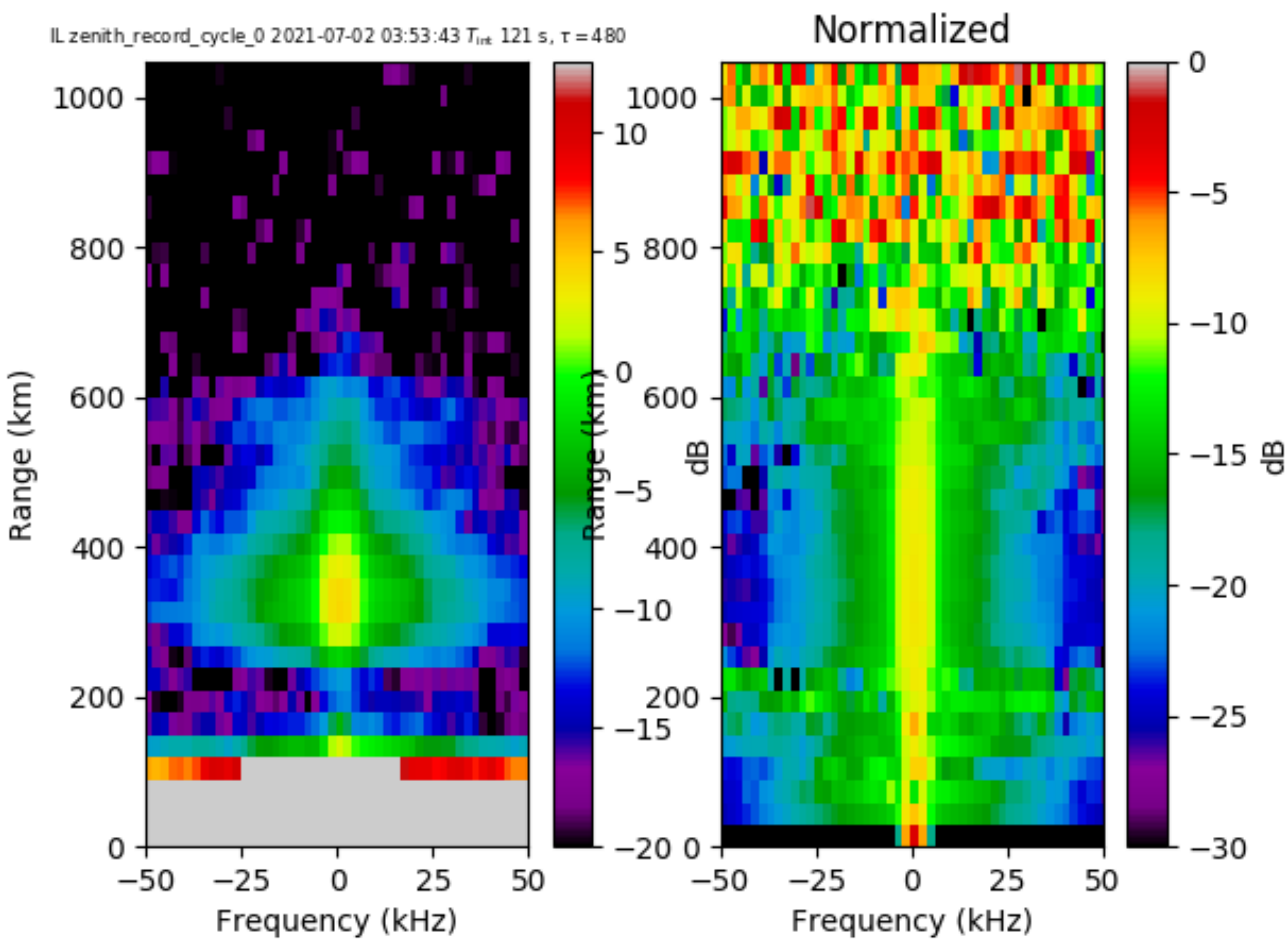
Electron (= total ion) density from total area (corrected for temperatures)

Line-of-sight ion velocity (Vi) from the Doppler shift

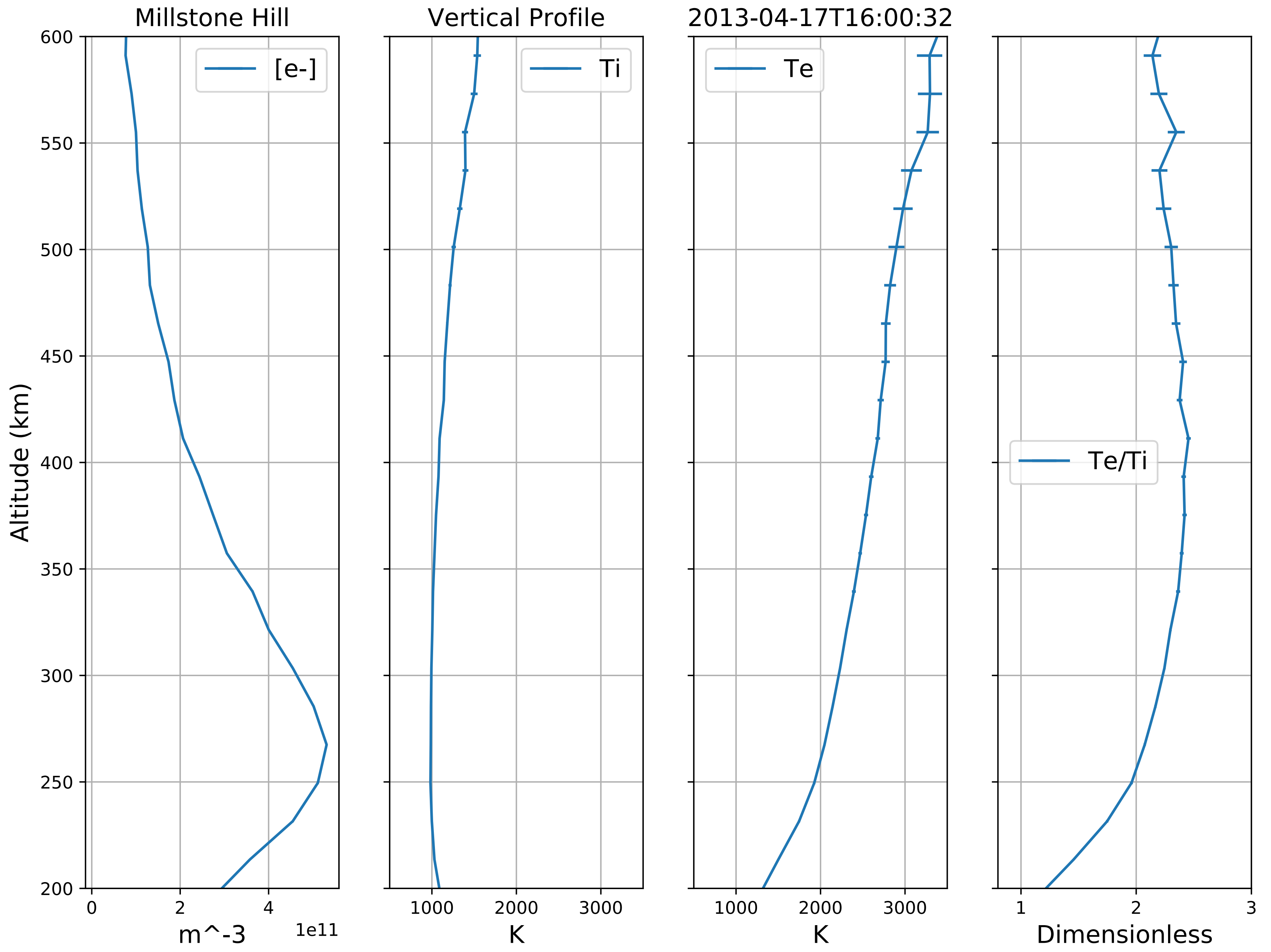
(NB: only one ion species here)

Ion Line Spectral Example as a Function of Altitude

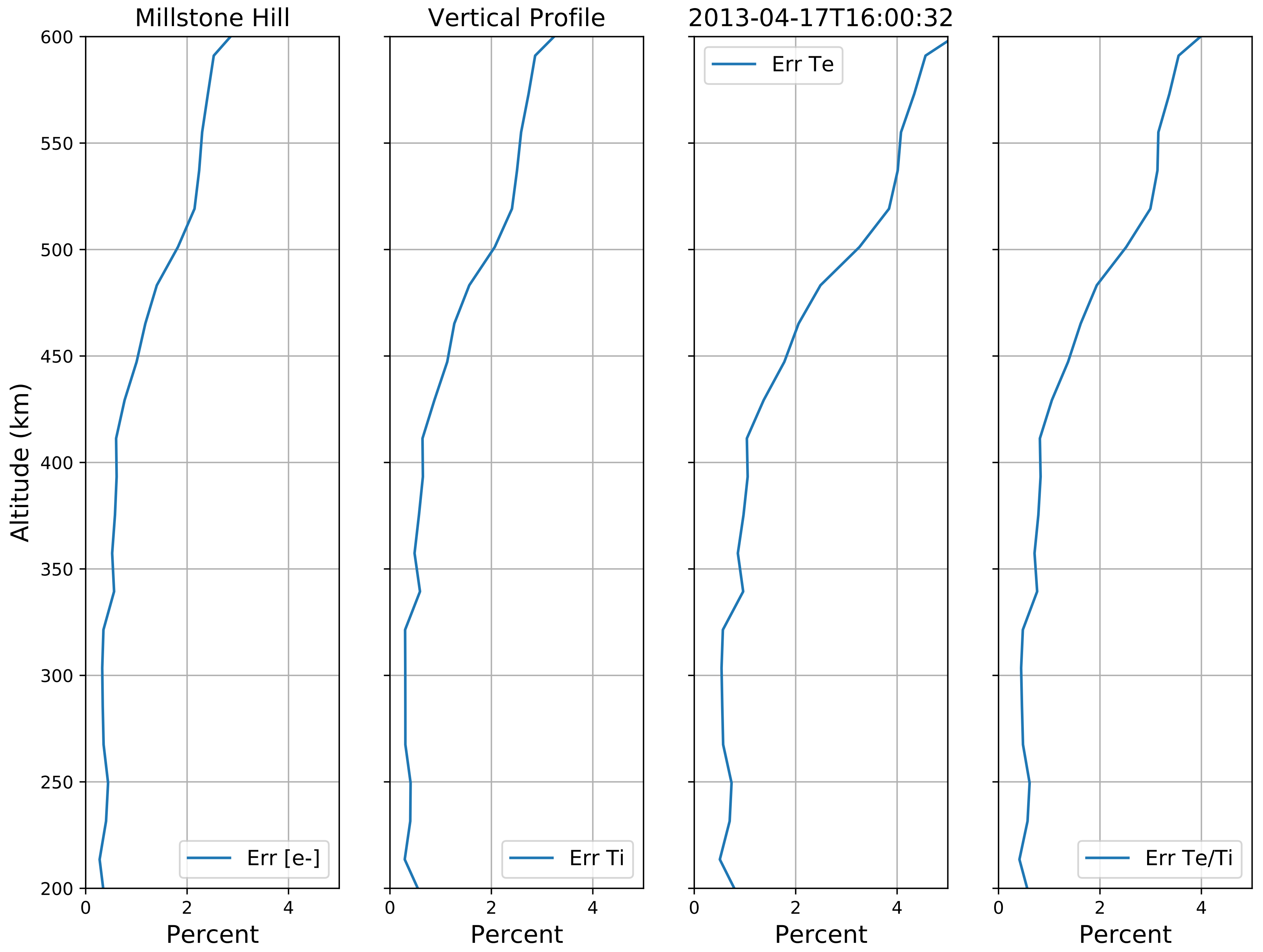
Millstone Hill vertical direction



Fitted Example: Ionospheric Parameter Altitude Profile

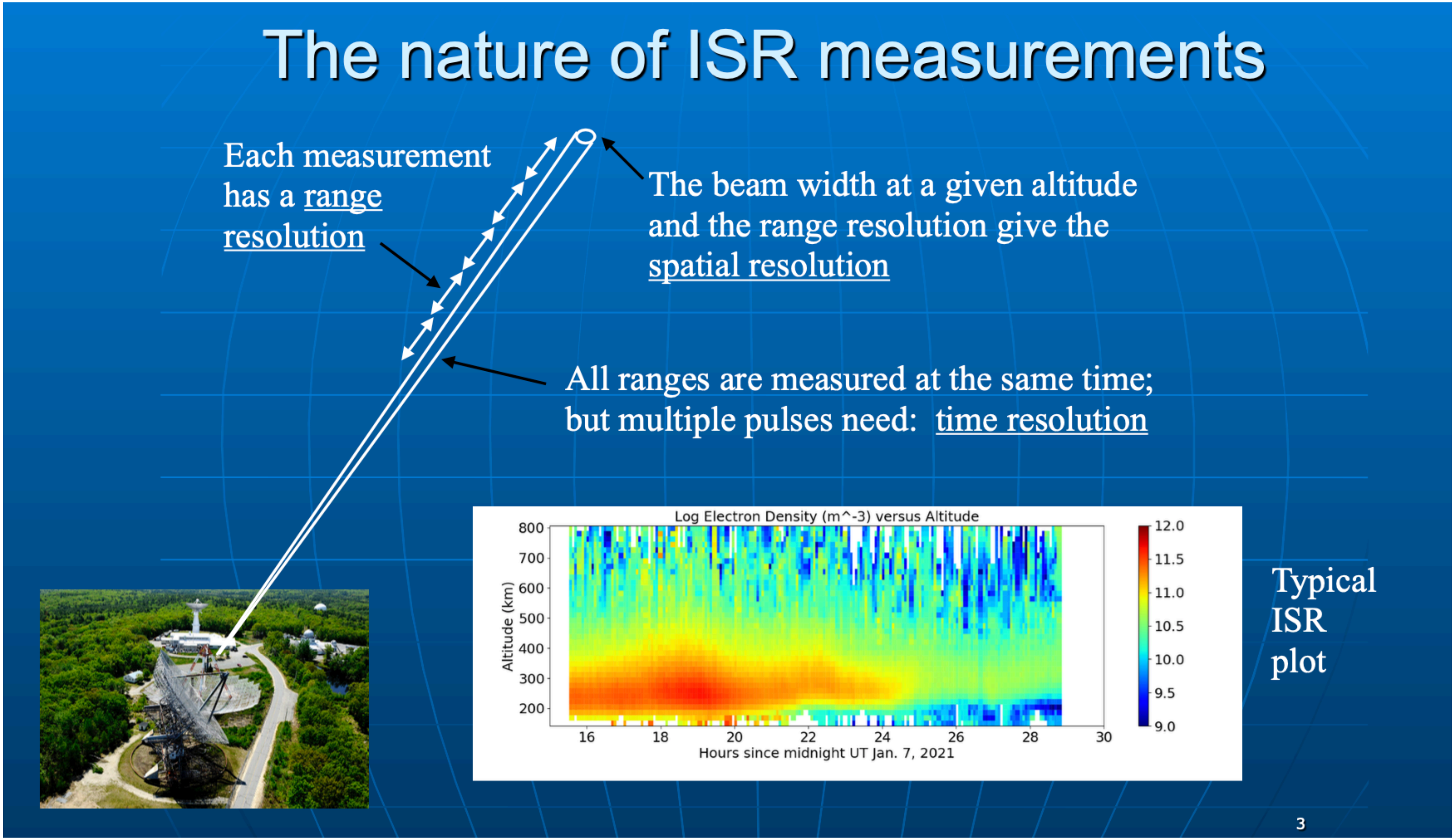


Fitted Example: Ionospheric Parameter Uncertainties

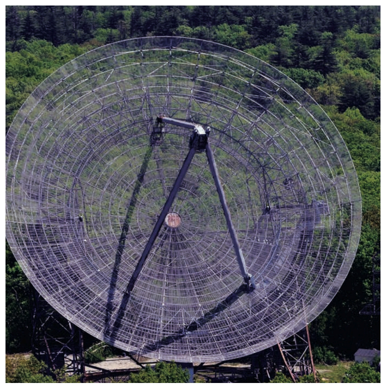


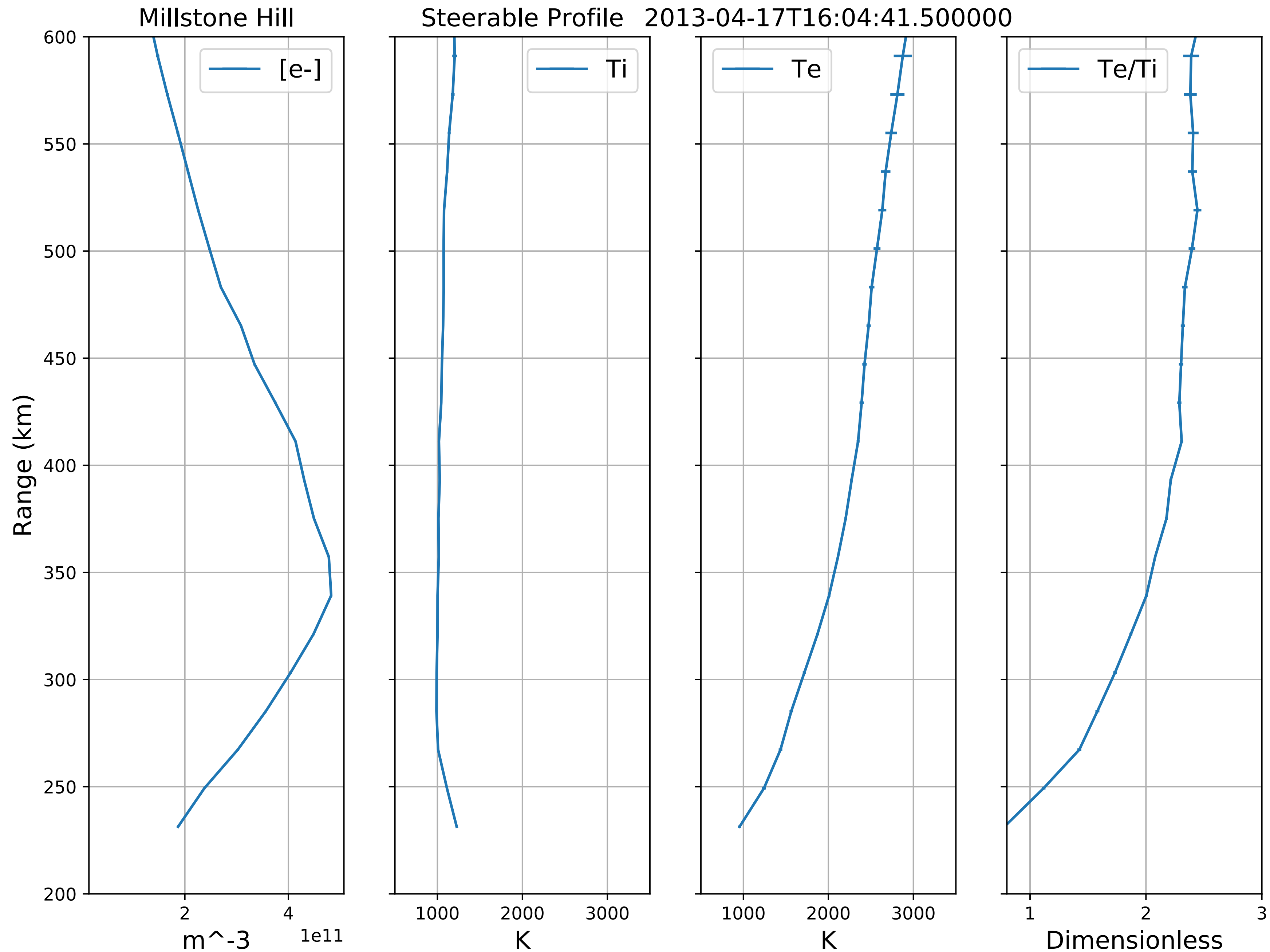
Range vs Altitude

(From Bill Rideout's "Black Box" talk)



Fitted Example: Steerable Antenna **Range** Profile



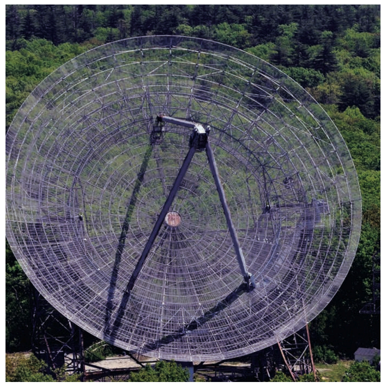


Peak Range

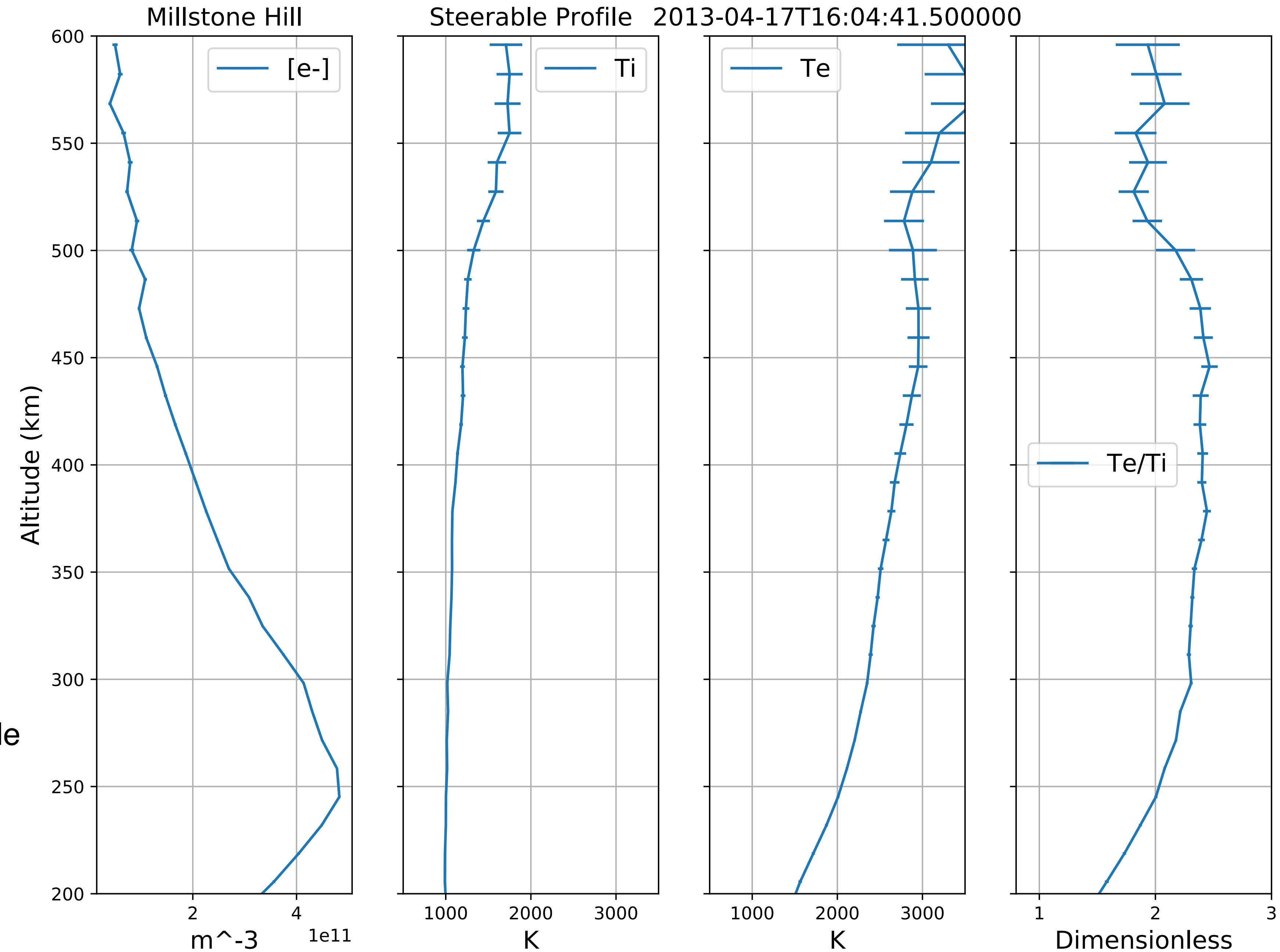
Mid-latitude daytime does not often have peak altitudes this high… Hmm.

.. But elevation angle = 45 degrees

Fitted Example: Steerable Antenna **Altitude** Profile



SAME DATA as previous slide

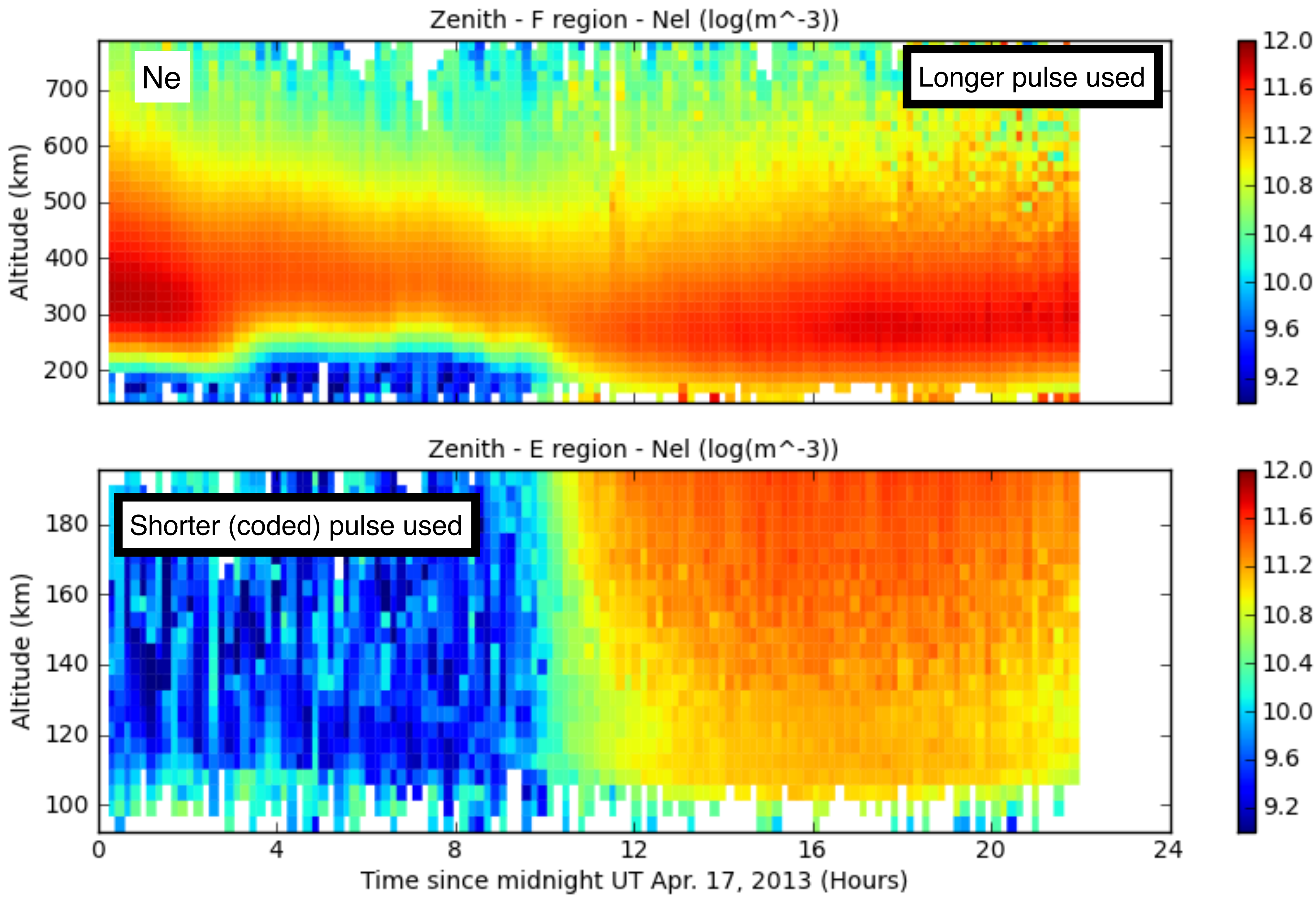


True Peak Altitude

Elevation Angle Taken Into Account

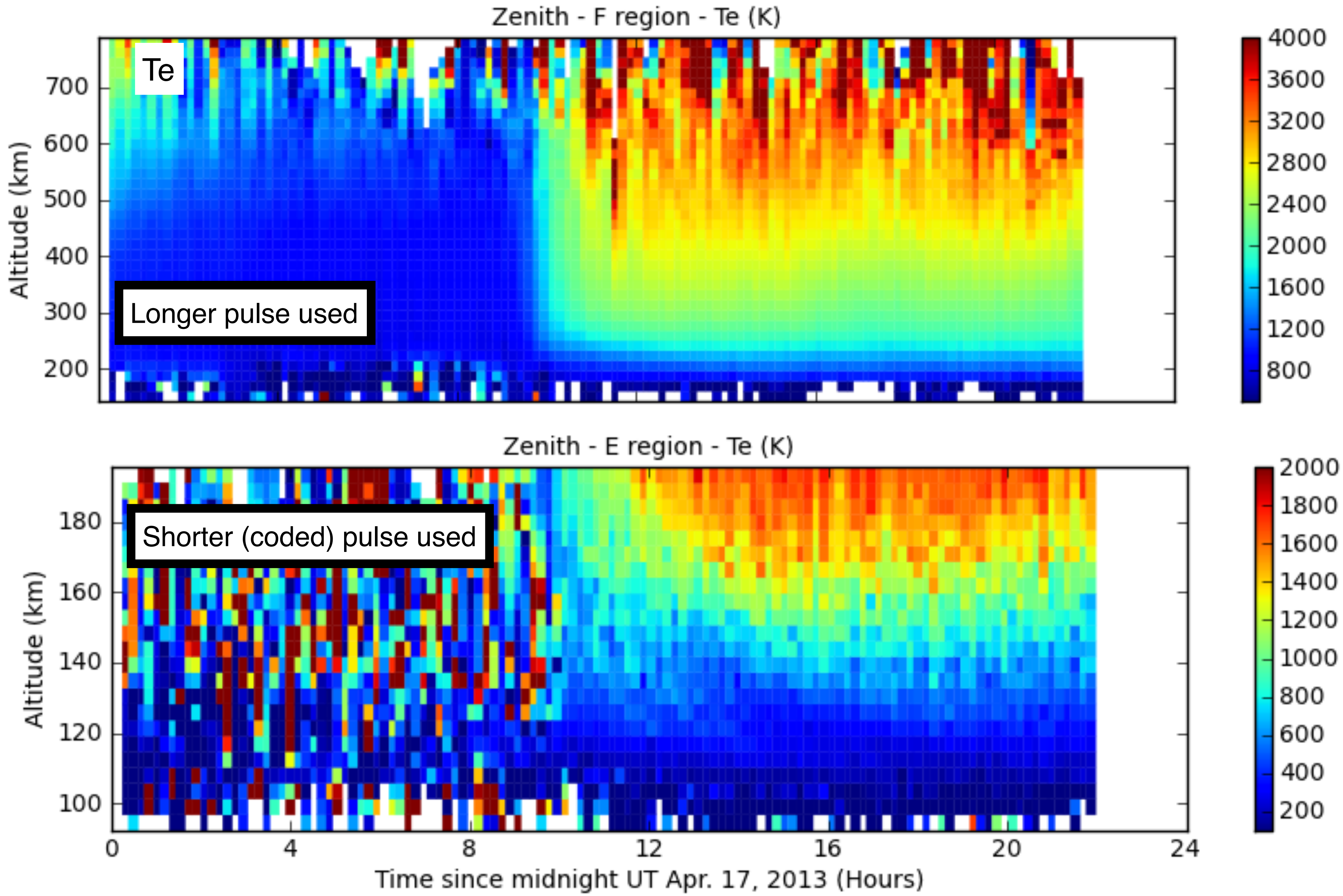
Basic IS Radar Measured Parameters (Ion Line)

Millstone Hill Vertical



Basic IS Radar Measured Parameters (Ion Line)

Millstone Hill Vertical



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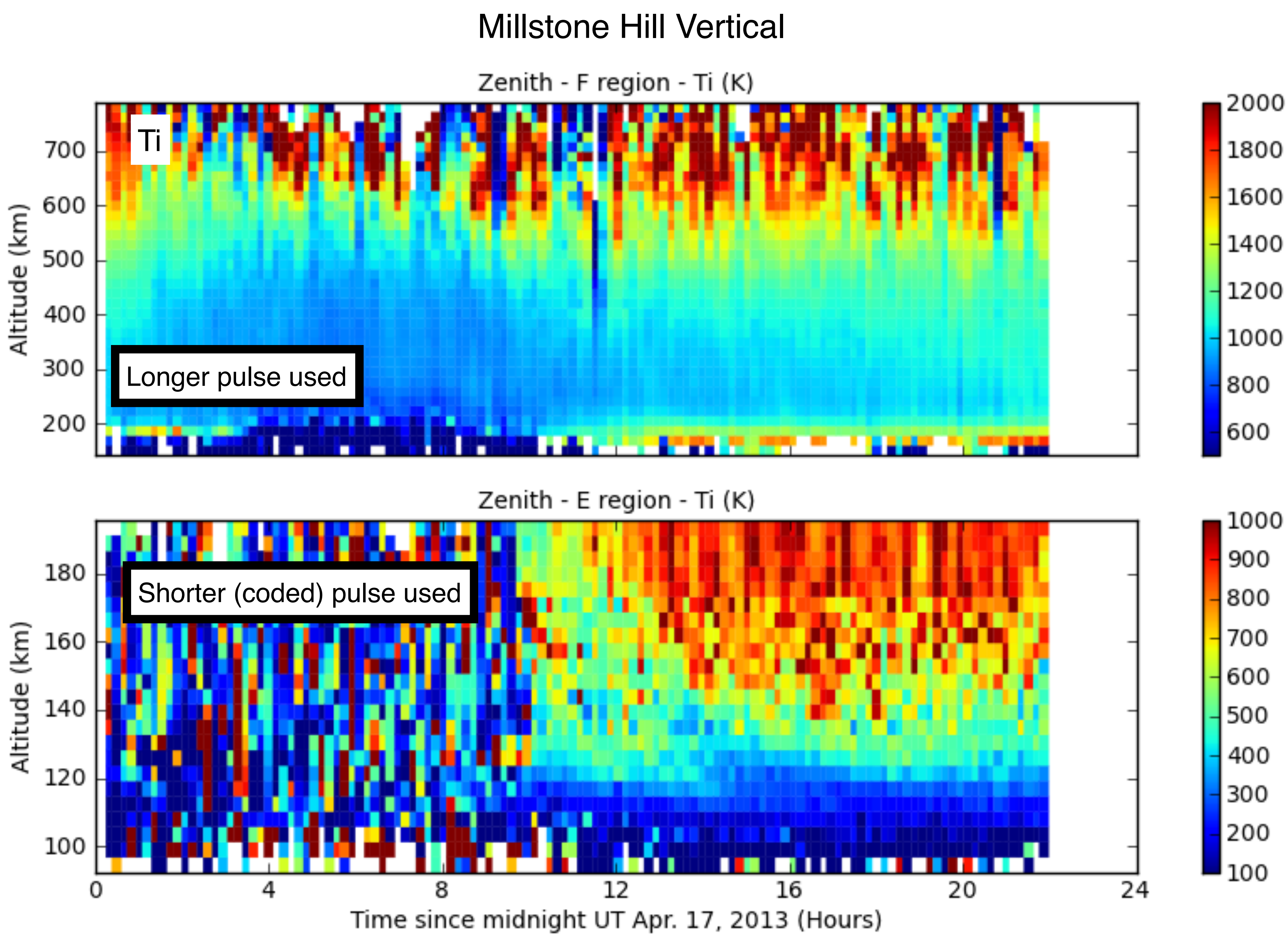

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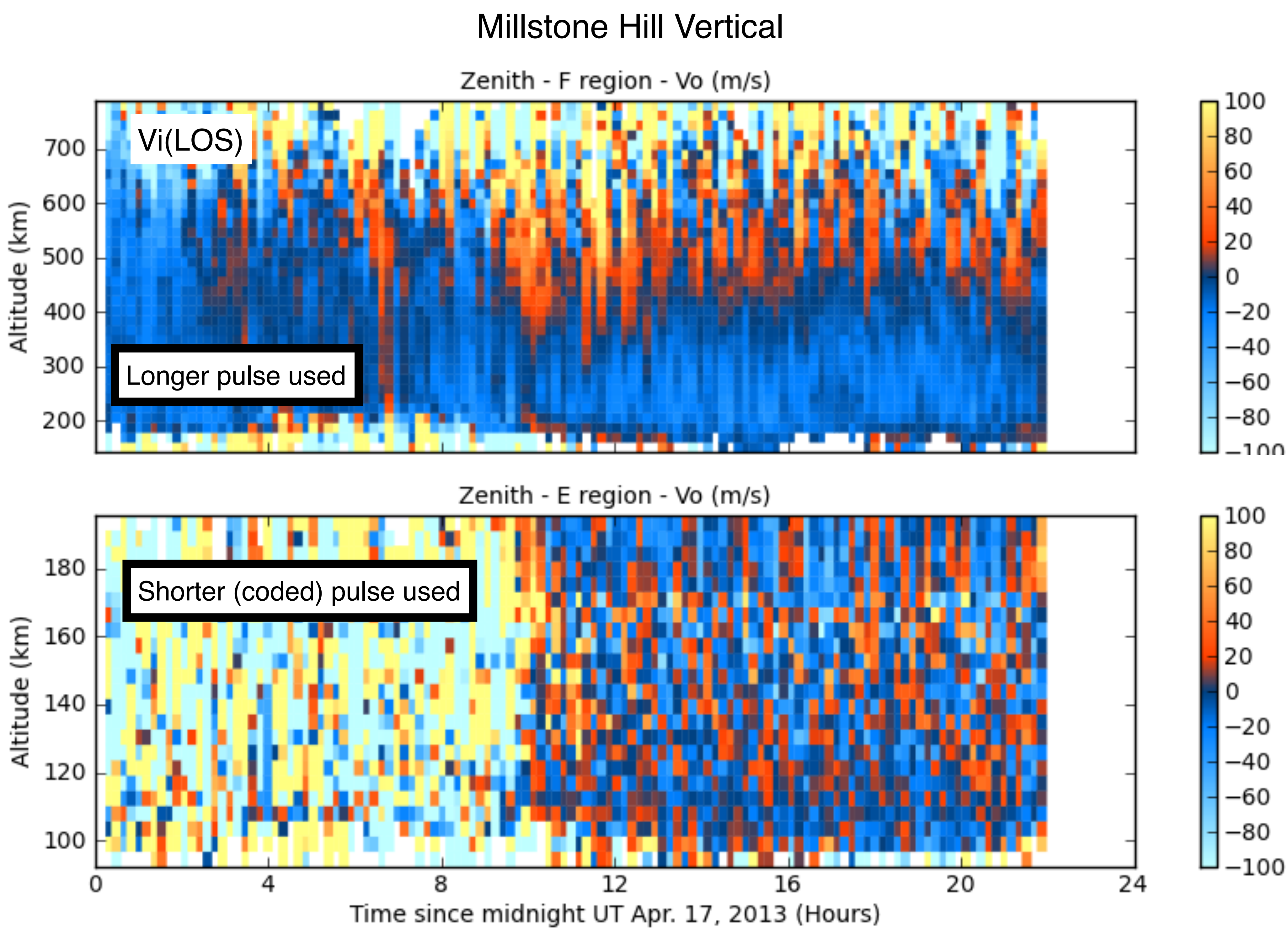

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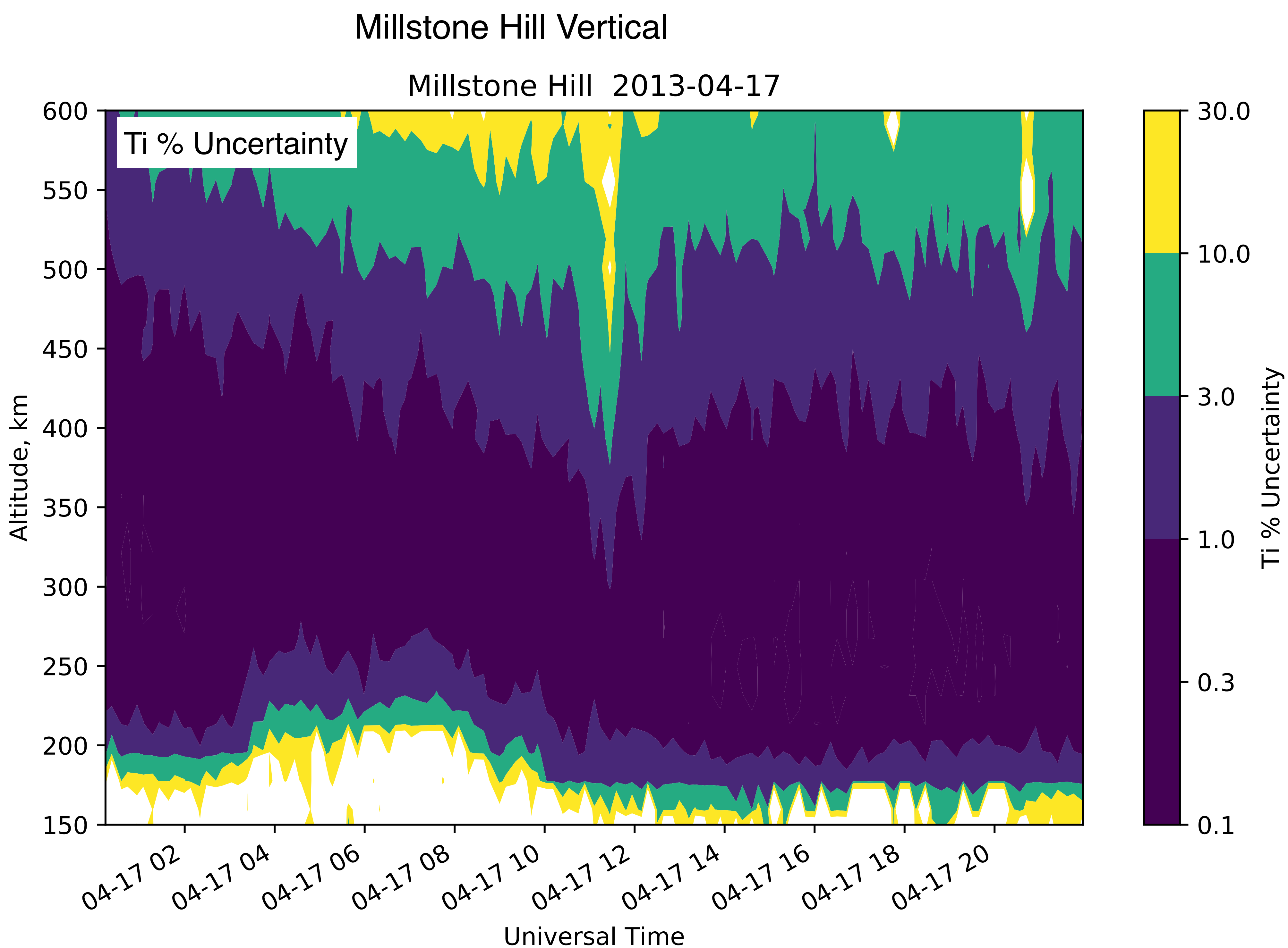
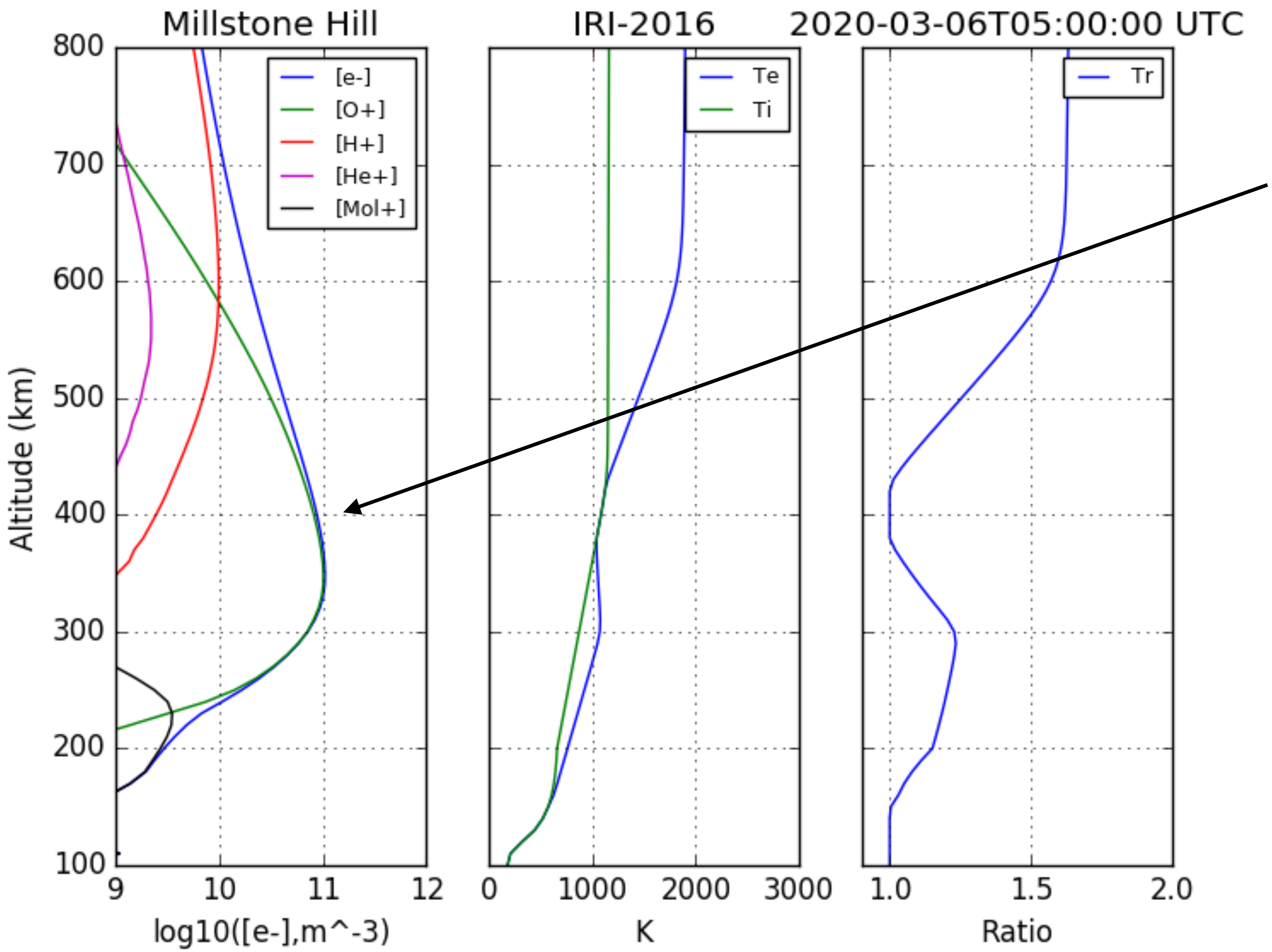


Uncertainties are available on each parameter

Ion Composition Fractions

Geophysical Location Dependent



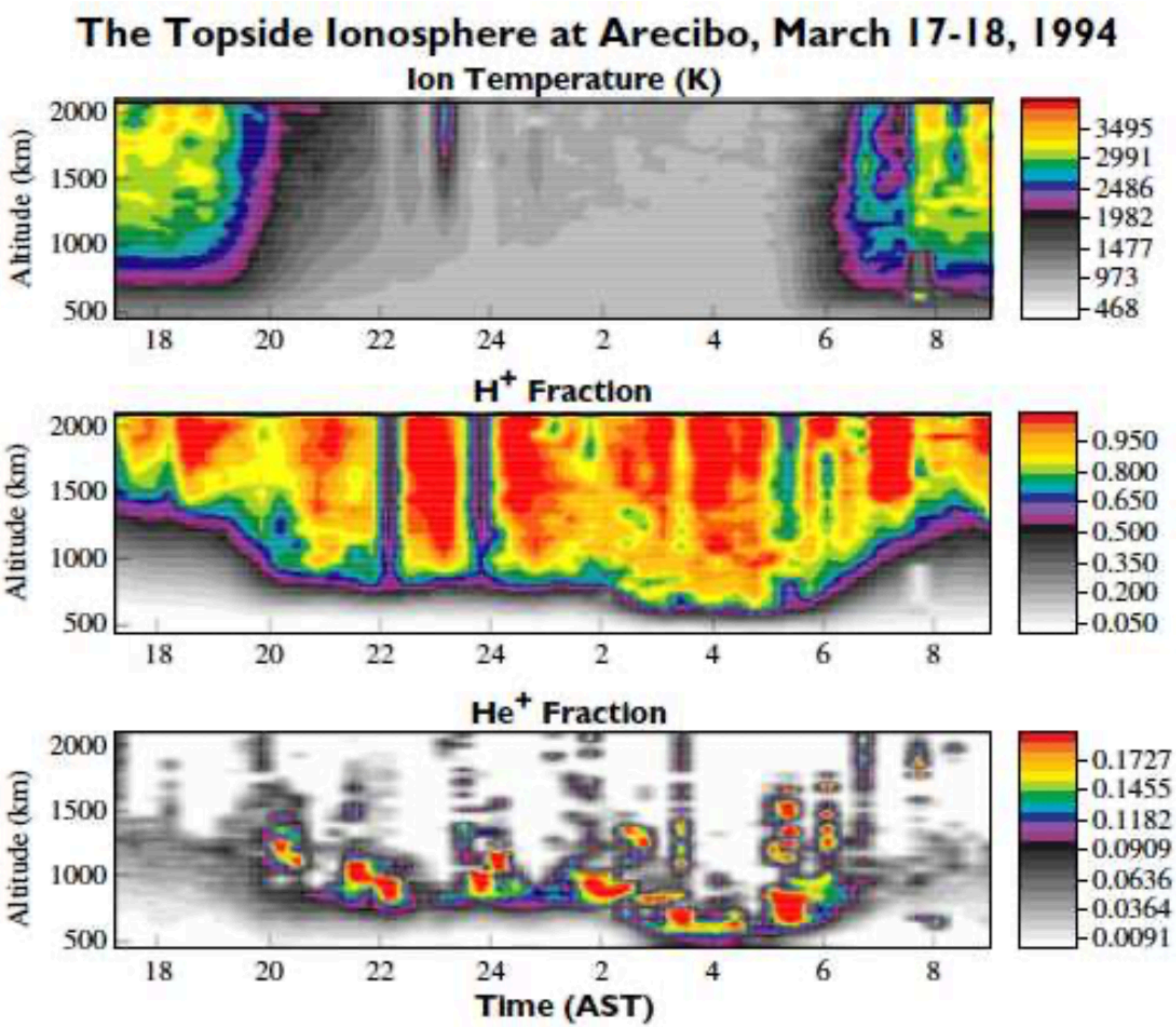
Ion composition (which ions allowed at what altitude) can be set by priors from e.g. modeling

Good assumptions can be also made - e.g. O+ is the only ion species near the F2 region

Allows Ti measurement through resolution of Ti/mi ratio ambiguity inherent in ion-acoustic resonance

Fraction of each ion can be fit in most 2-ion cases (occasionally 3-ions at Arecibo)

Ion Composition Fractions



Arecibo:

Topside fractions of O+, He+, H+ can be measured: high enough SNR (excellent statistics) through the F region ionosphere and into the topside ionosphere

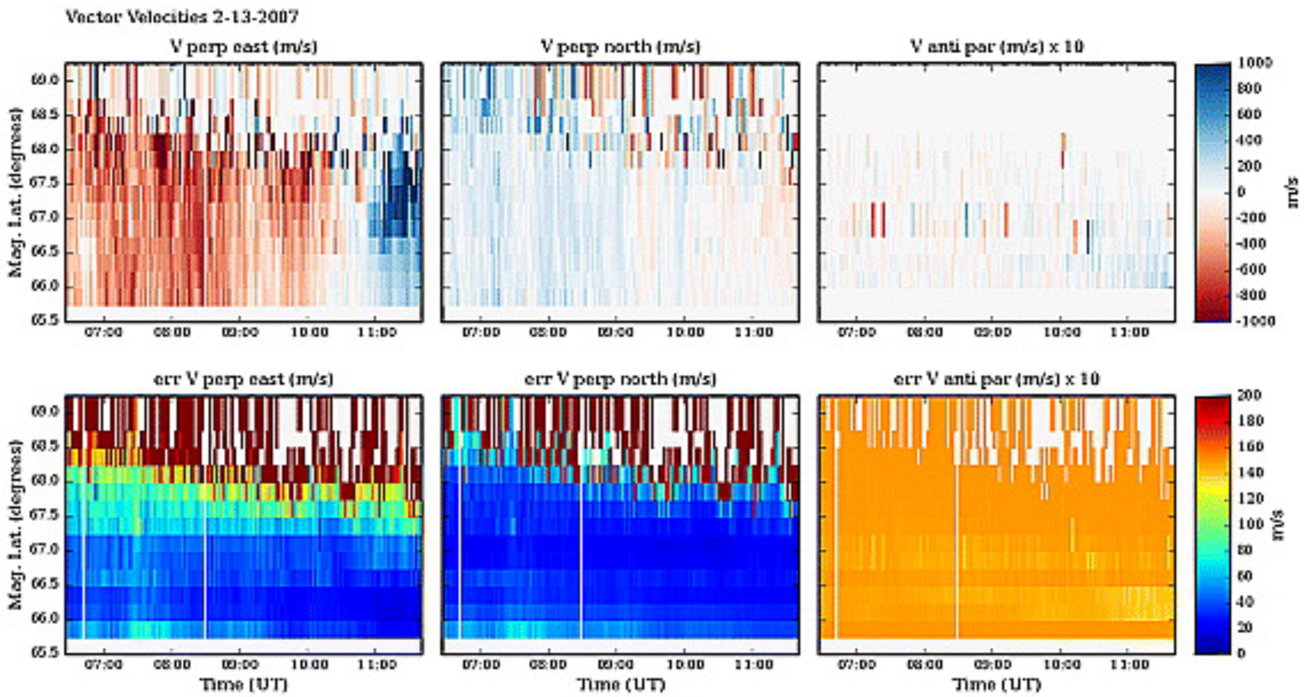
Gonzalez and Sulzer 1996
doi:10.1029/96GL02212

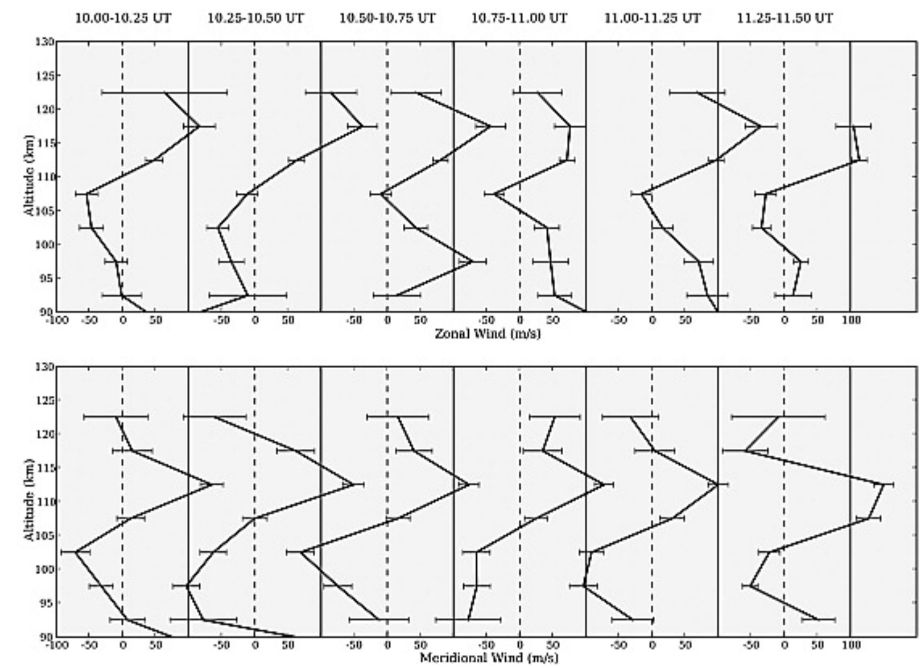
Derived Parameters

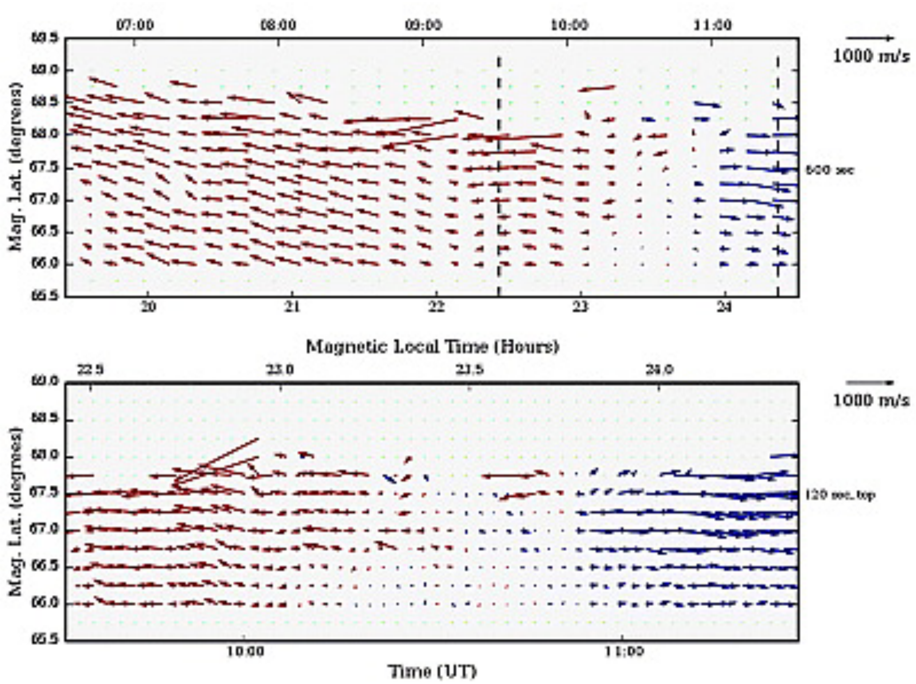
With the basic parameters as a function of space and time, one can go farther: apply constraint and physics to produce additional ionospheric and plasma physics parameters.

We'll call those Derived parameters because they are not produced from the basic incoherent scatter radar observables we've just covered.

Derived Parameters: Vector Ion Velocities, Neutral Winds







$$\begin{bmatrix} v_{los}^1 \\ v_{los}^2 \\ \vdots \\ v_{los}^n \end{bmatrix} = \begin{bmatrix} k_{pe}^1 & k_{pn}^1 & k_{ap}^1 \\ k_{pe}^2 & k_{pn}^2 & k_{ap}^2 \\ \vdots & \vdots & \vdots \\ k_{pe}^n & k_{pn}^n & k_{ap}^n \end{bmatrix} \begin{bmatrix} v_{pe} \\ v_{pn} \\ v_{ap} \end{bmatrix} + \begin{bmatrix} e_{los}^1 \\ e_{los}^2 \\ \vdots \\ e_{los}^n \end{bmatrix}$$

$$0 = e\left(\mathbf{E} + \mathbf{v}_i \times \mathbf{B}\right) - m_i \nu_{in} \left(\mathbf{v}_i - \mathbf{u}\right)$$

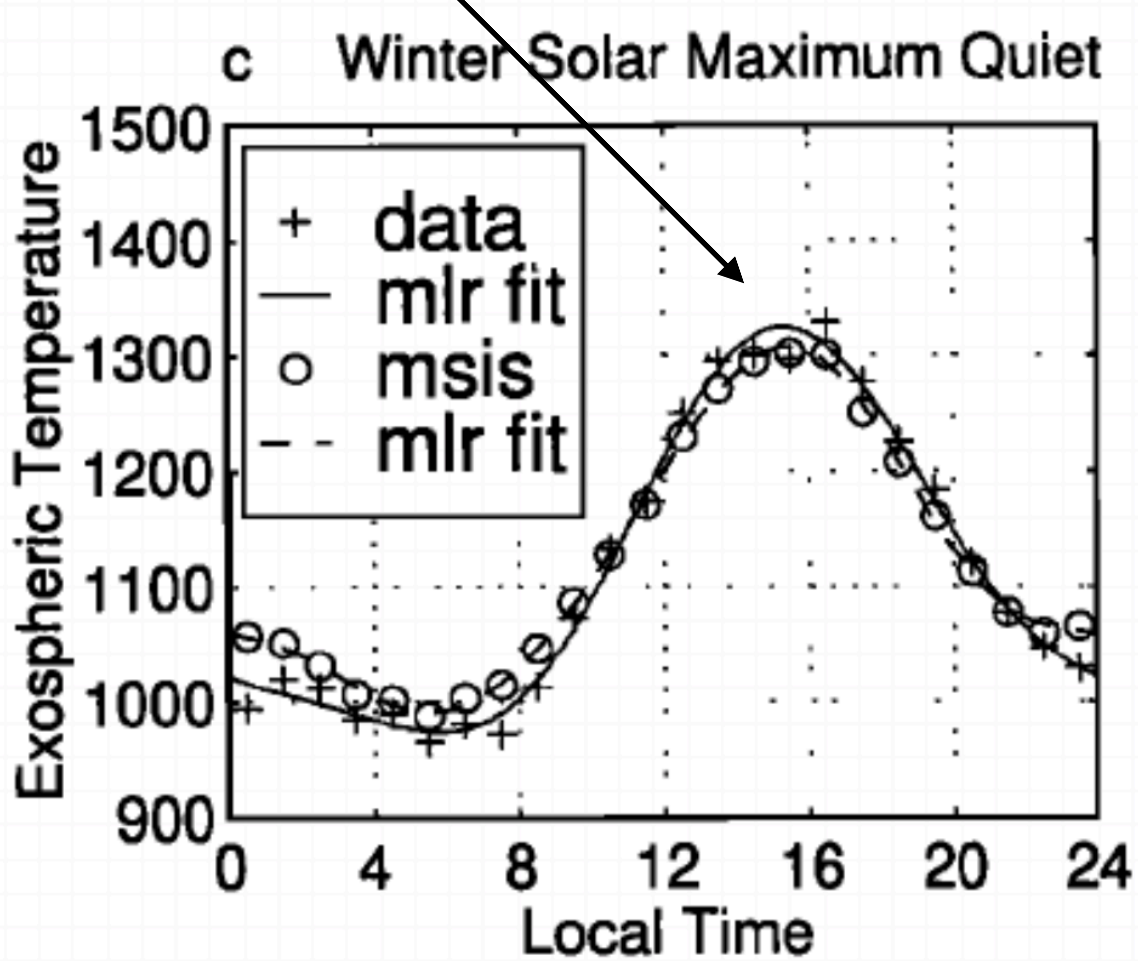
$$\mathbf{v}_i = b_i C\mathbf{E} + C\mathbf{u} \qquad C = \begin{bmatrix} \left(1+\kappa_i^2\right)^{-1} & -\kappa_i\left(1+\kappa_i^2\right)^{-1} & 0 \\ \kappa_i\left(1+\kappa_i^2\right)^{-1} & \left(1+\kappa_i^2\right)^{-1} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

Heinselman, C. J., and Nicolls, M. J. (2008), *Radio Sci.*, doi:10.1029/2007RS003805.

Derived Parameters: Neutral Temperatures

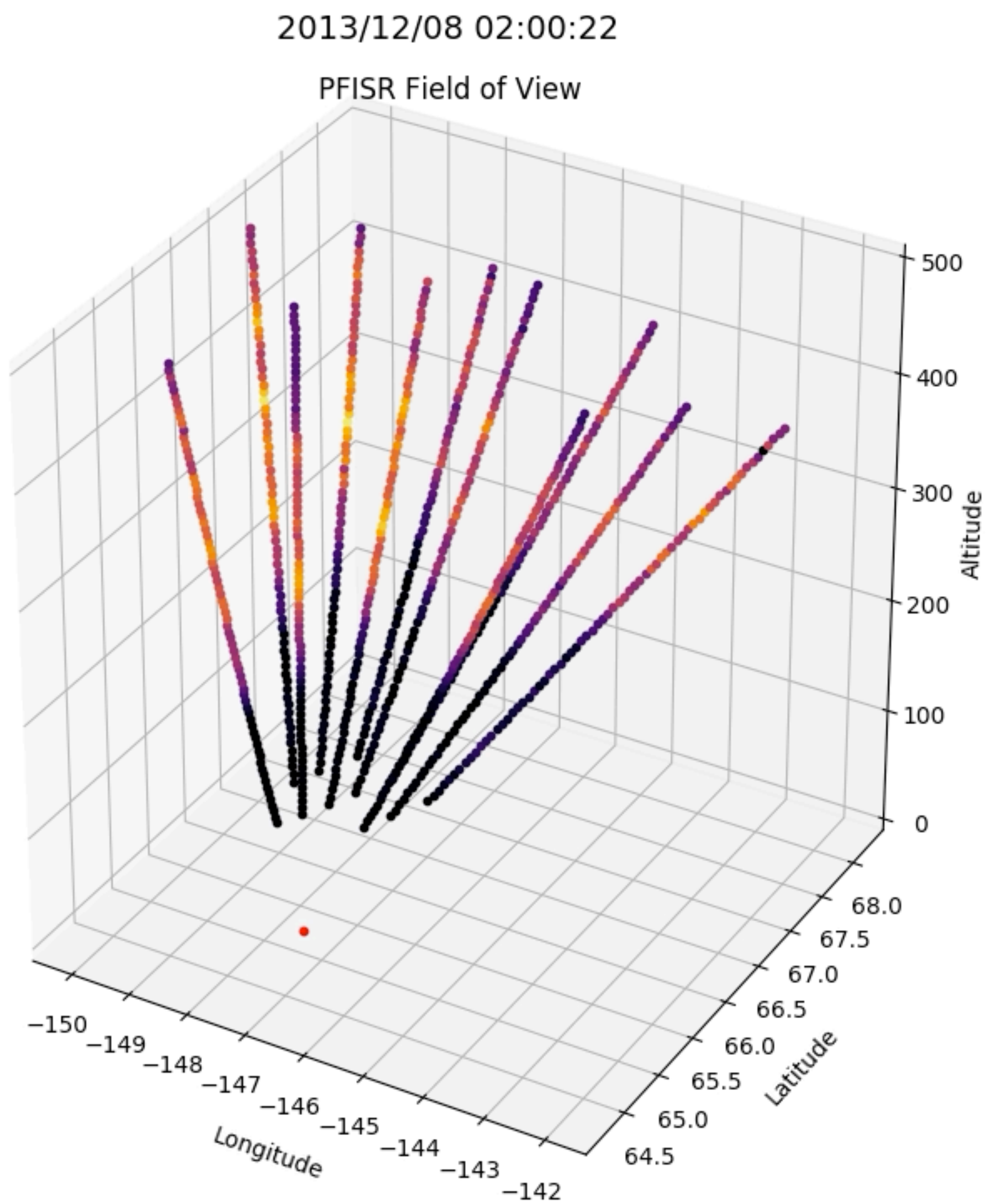
$$\sum_i L_{ei} = \sum_{i,n} L_{in} \qquad a(T_e - T_i) = b(T_i - T_n)$$

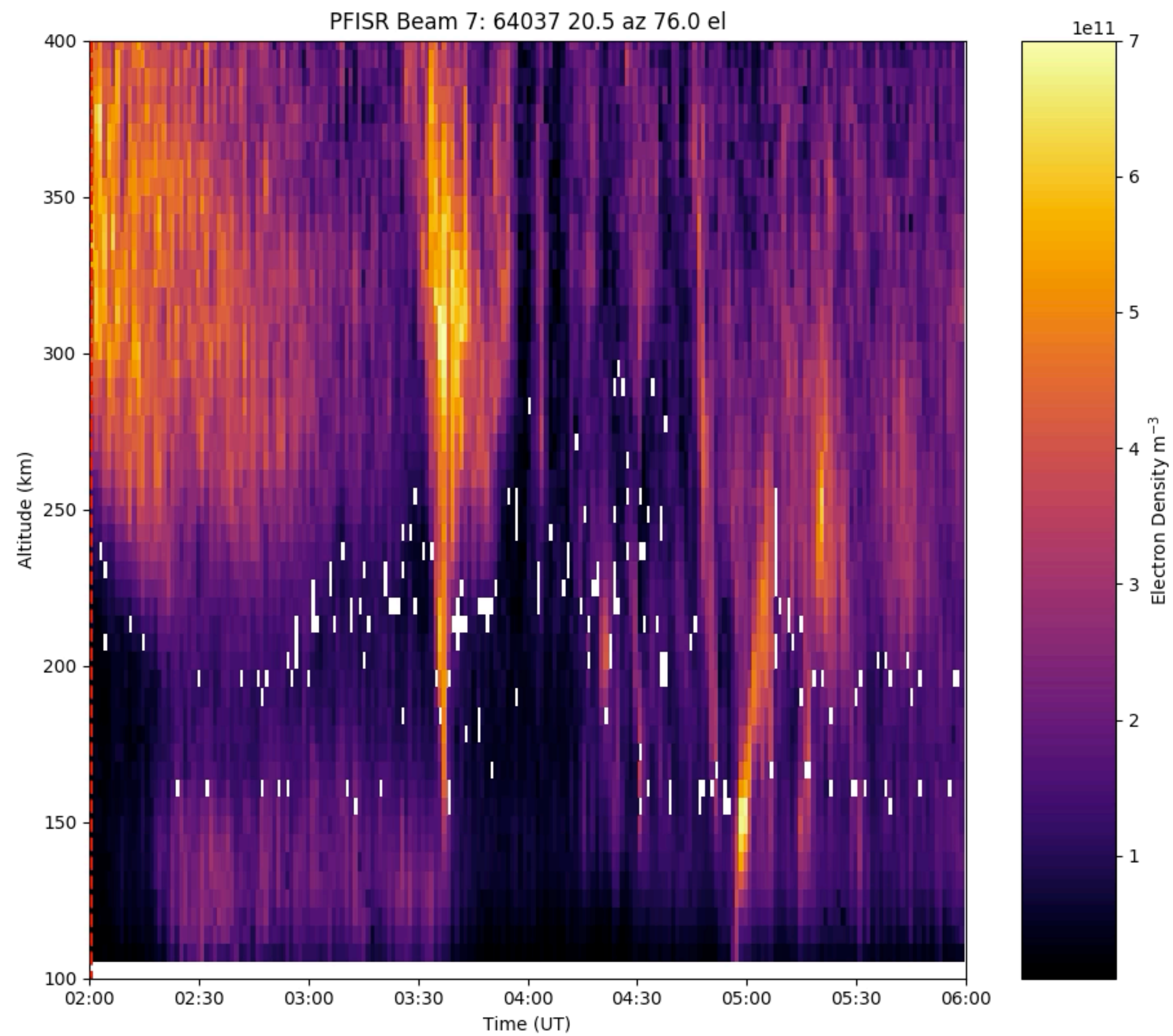
$$T_n(z) = T_\infty - (T_\infty - T_0) exp[-s(z - z_0)(R_E + z_0)/(R_E + z)] \quad (3)$$



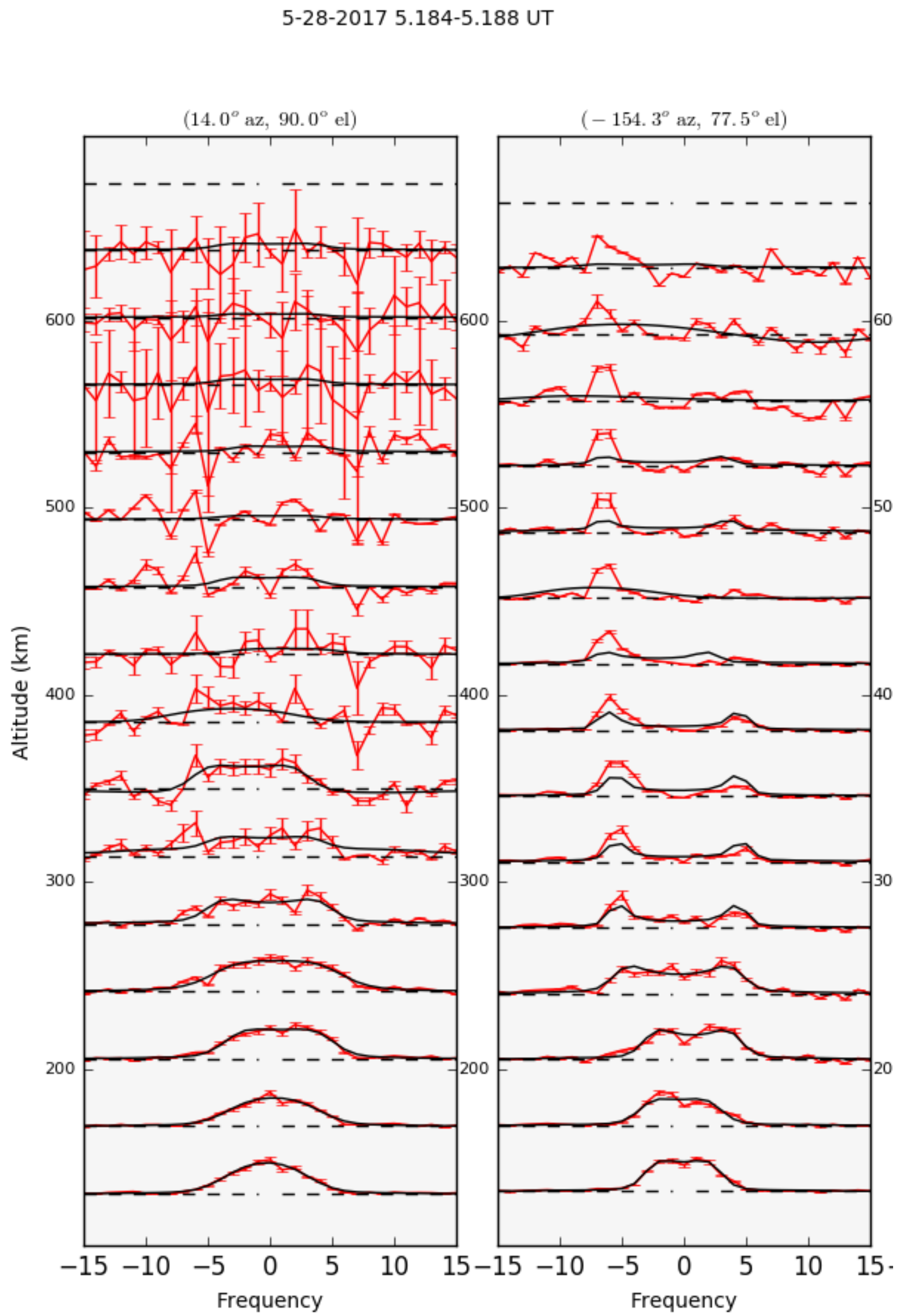
Buonsanto, M. J., and Pohlman, L. M. (1998), Climatology of neutral exospheric temperature above Millstone Hill, *J. Geophys. Res.*, 103(A10), 23381–23392, doi:10.1029/98JA01919.

PFISR Multi-beam Measurements of Auroral Ionization



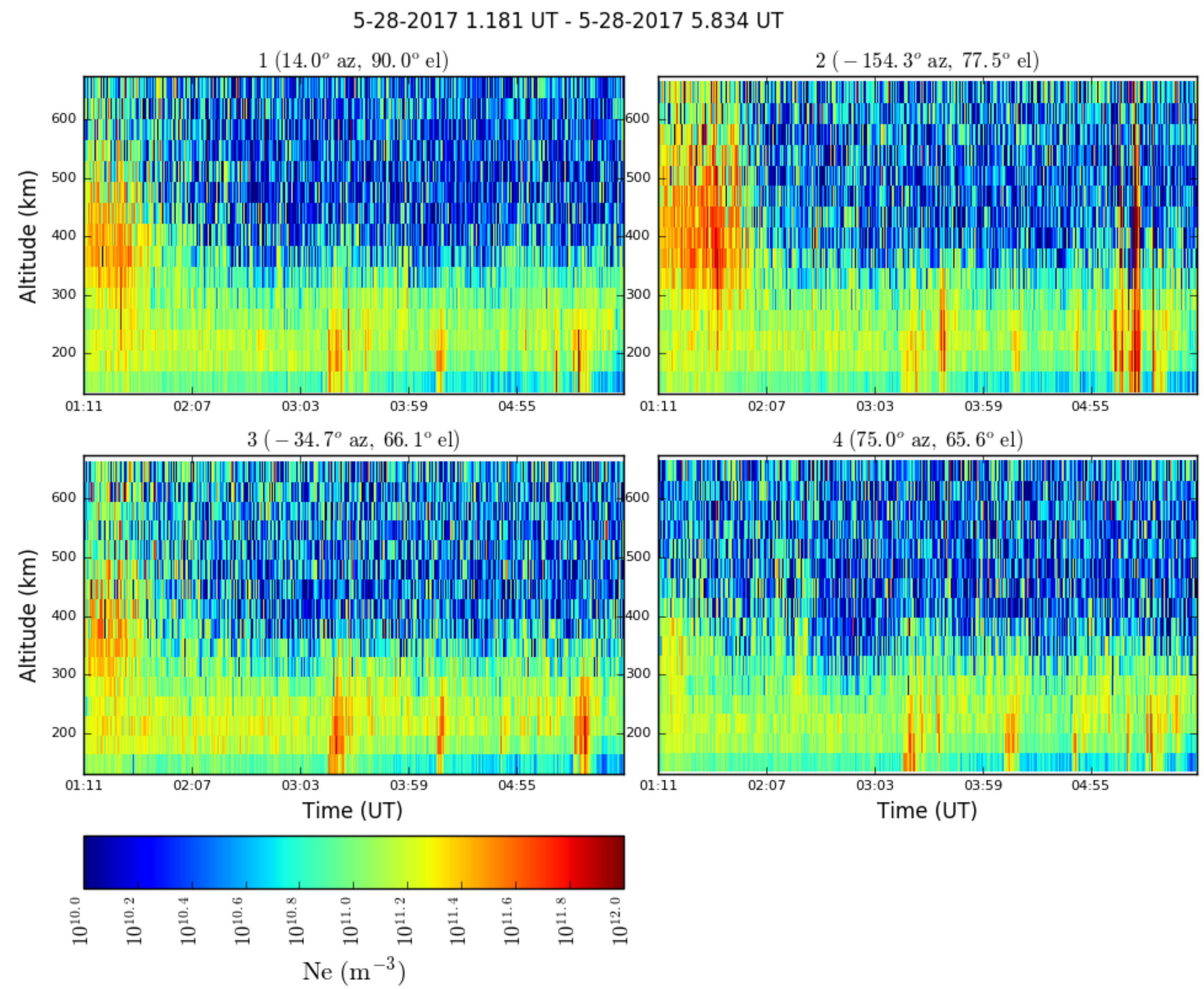


PFISR Naturally Enhanced Ion-Acoustic Lines

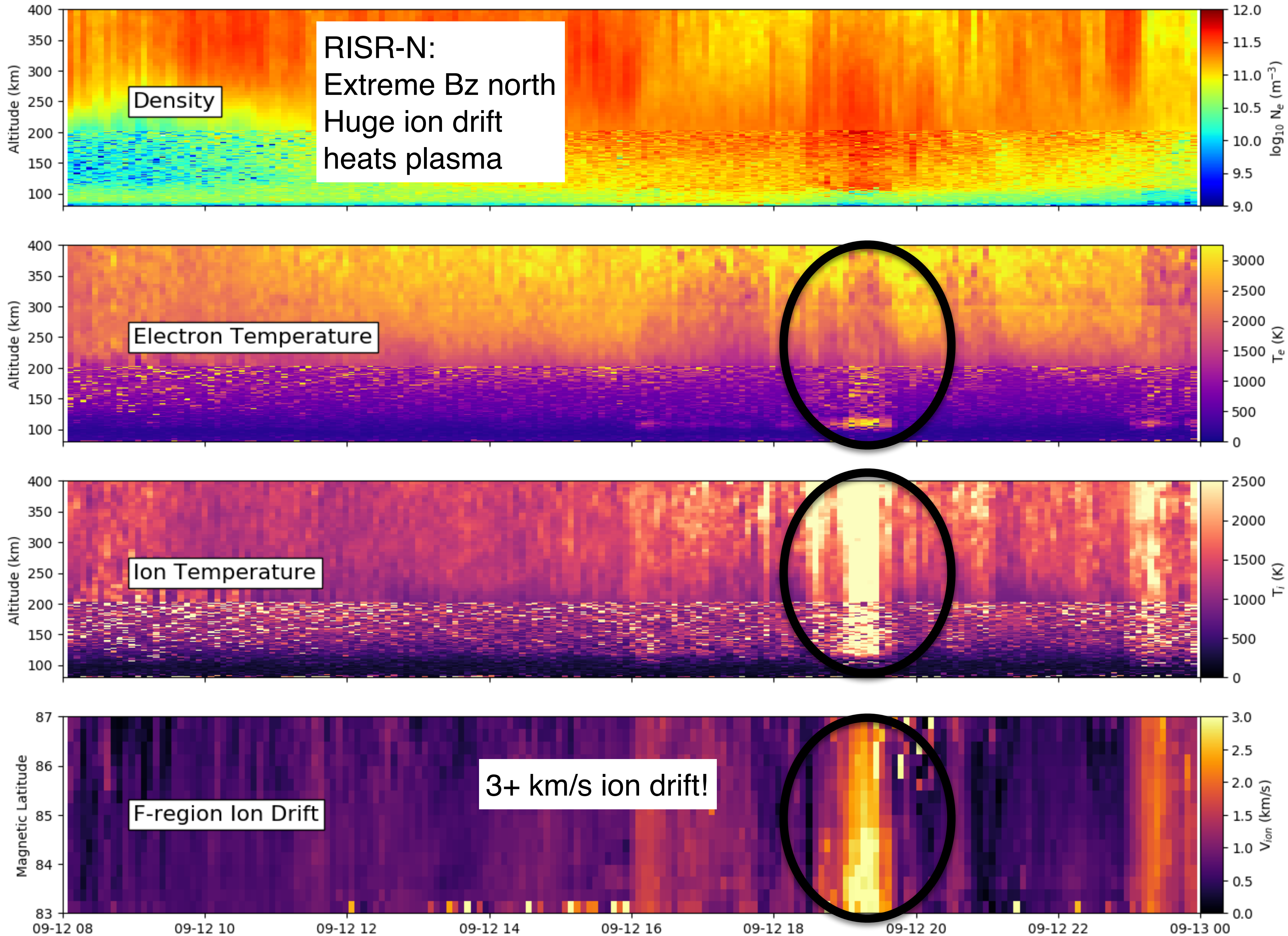


Plasma instability process [current driven F region instabilities, ion-ion 2 stream, Langmuir wave parametric decay]

Particle precipitation is directly or indirectly involved
E.g. Lunde et al, 2007, doi:10.5194/angeo-25-1323-2007

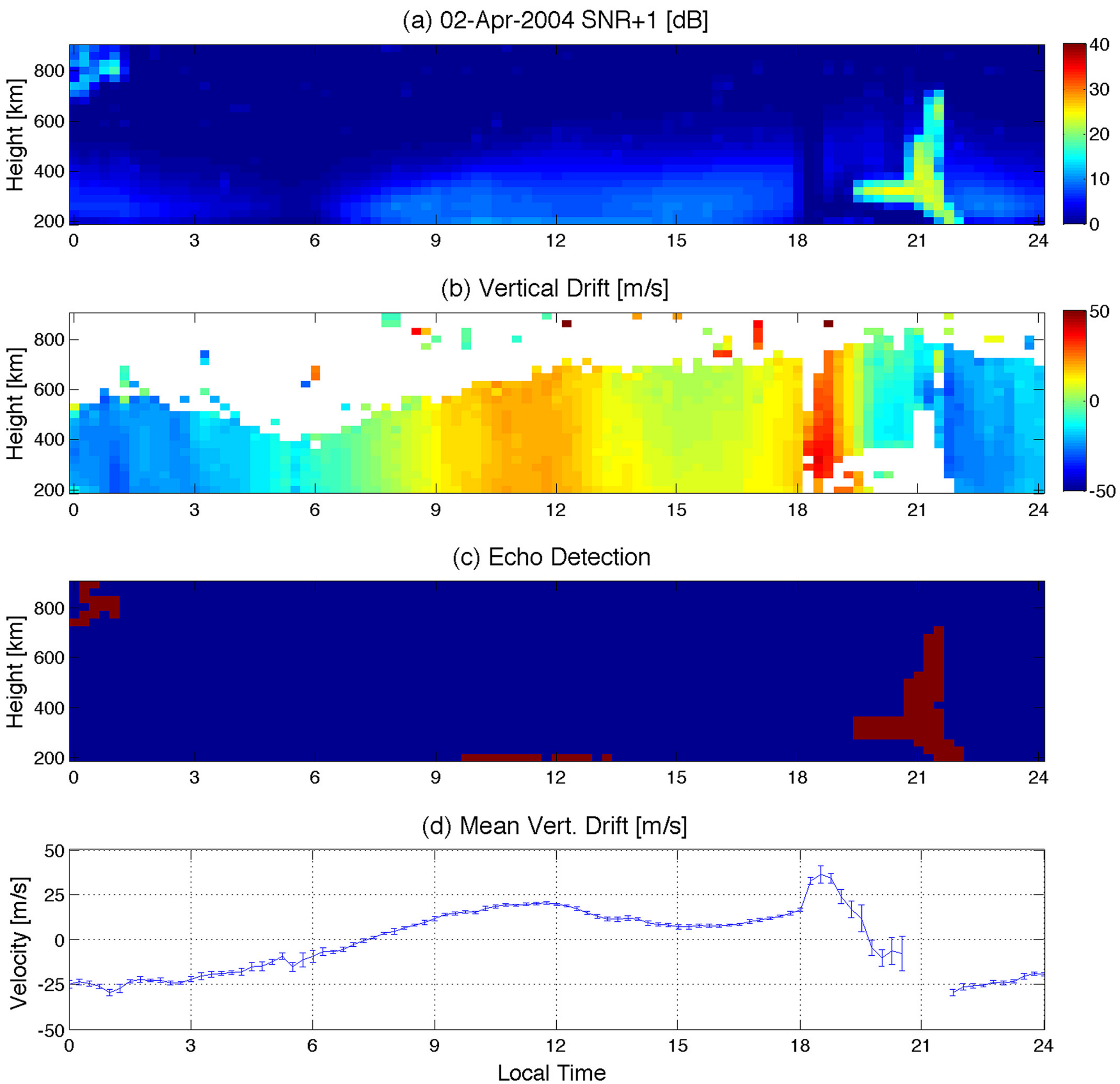


Polar Cap Response to Non-Saturated Potential Drop



cf.
Clauer et al. JGR 2016
doi:10.1002/2016JA022557

Equatorial spread F simultaneous with vertical drifts



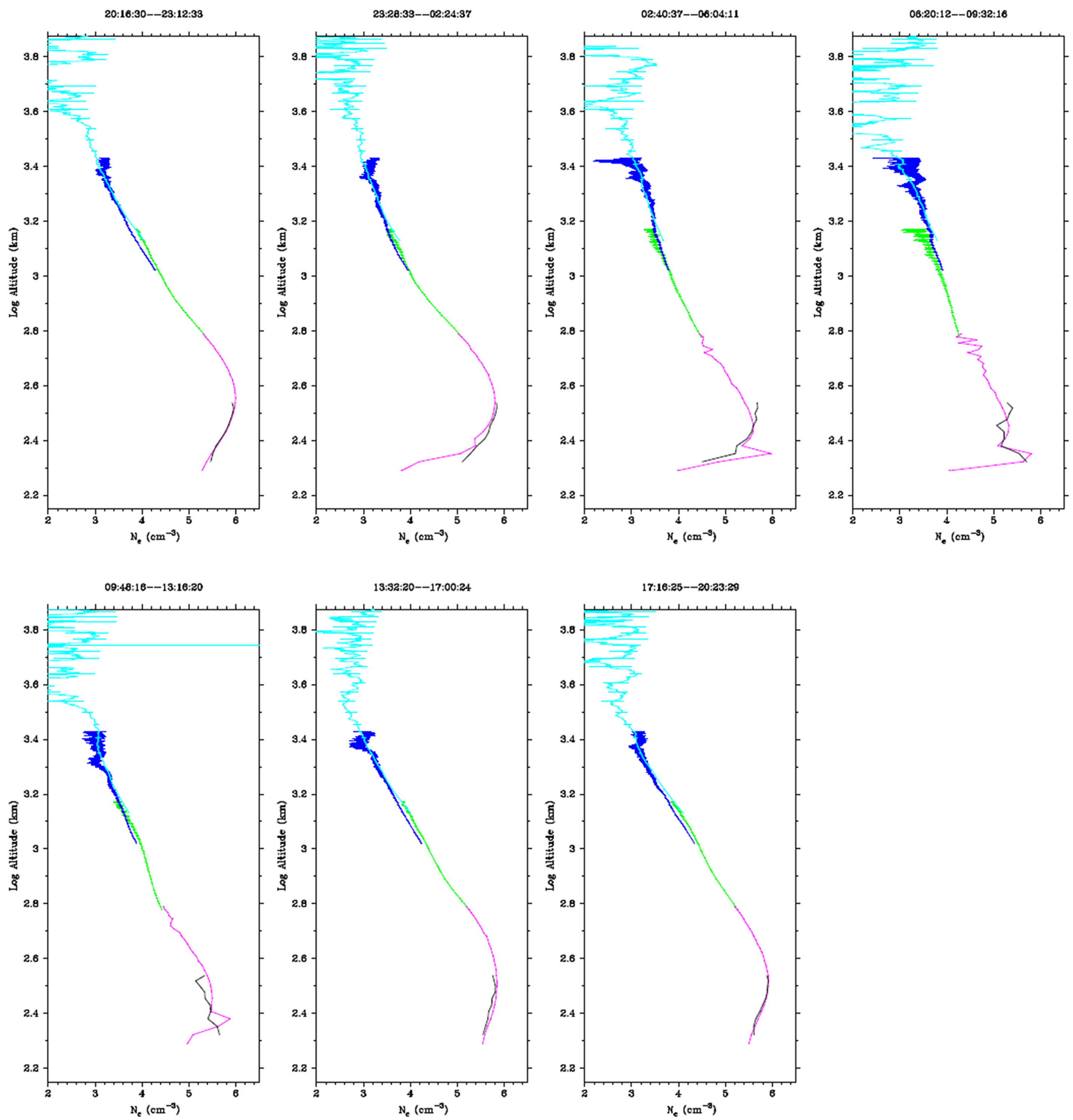
Jicamarca

Near-perpendicular to B = highly accurate vertical drifts

Dusk sector: spread-F coherent irregularities (very bright)

Smith, J. M., Rodrigues, F. S., Fejer, B. G., and Milla, M. A. (2016), Coherent and incoherent scatter radar study of the climatology and day-to-day variability of mean F region vertical drifts and equatorial spread F , J. Geophys. Res. Space Physics, 121, 1466– 1482, doi:10.1002/2015JA021934.

Extreme High-Altitude Equatorial Electron Density



Jicamarca
4 transmitters (6 MW peak) @ 50 MHz

Electron density to L=2!
(~6,000 km altitude)

Possible to 10,000 km in daytime

(2500 km threads through Arecibo field of view @ 400 km altitude)

Larger system would be able to perhaps see plasmapause from the ground (connections to SED plume)

Hysell, D. L., Milla, M. A., and Woodman, R. F. (2017), High-altitude incoherent-scatter measurements at Jicamarca, J. Geophys. Res. Space Physics, 122, 2292– 2299, doi:10.1002/2016JA023569.

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